

# First insights into ITB physics

A progress report for TSVV1 workshop

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Justin Ball, Stephan Brunner



**EUROfusion**

# EPFL Outline

- Internal and edge transport barriers
- Internal transport barrier formation
- Milestones and deliverables
- Turbulent self-interaction
- Preliminary results:
  - Linear simulations
  - Nonlinear simulations
- Current work

# Internal and edge transport barriers

## H-mode (edge transport barrier)

- Transport bifurcation
- Gradient steepening
- **ExB shearing flow**
- High magnetic shear

## Internal transport barrier (ITB)

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- Low or negative magnetic shear

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# Core region – easier to investigate numerically

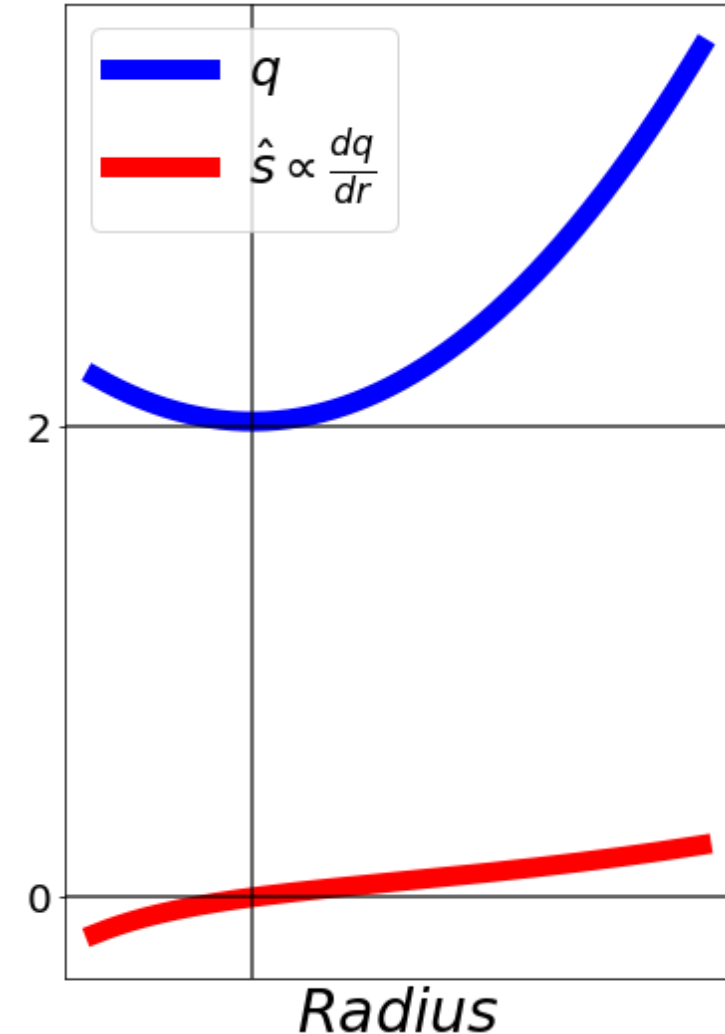
- No coupling to scrape-off layer;
- Relatively simple geometry;
- Low collisionality;
- Low fluctuation levels;
- No neutrals.

# EPFL ITB formation

- Low or negative magnetic shear  $\hat{s}$  is crucial;
- ITBs are often localized around rational  $q$ ;
- Presence of **integer minimum  $q$**  seems to be especially favourable for ITB formation

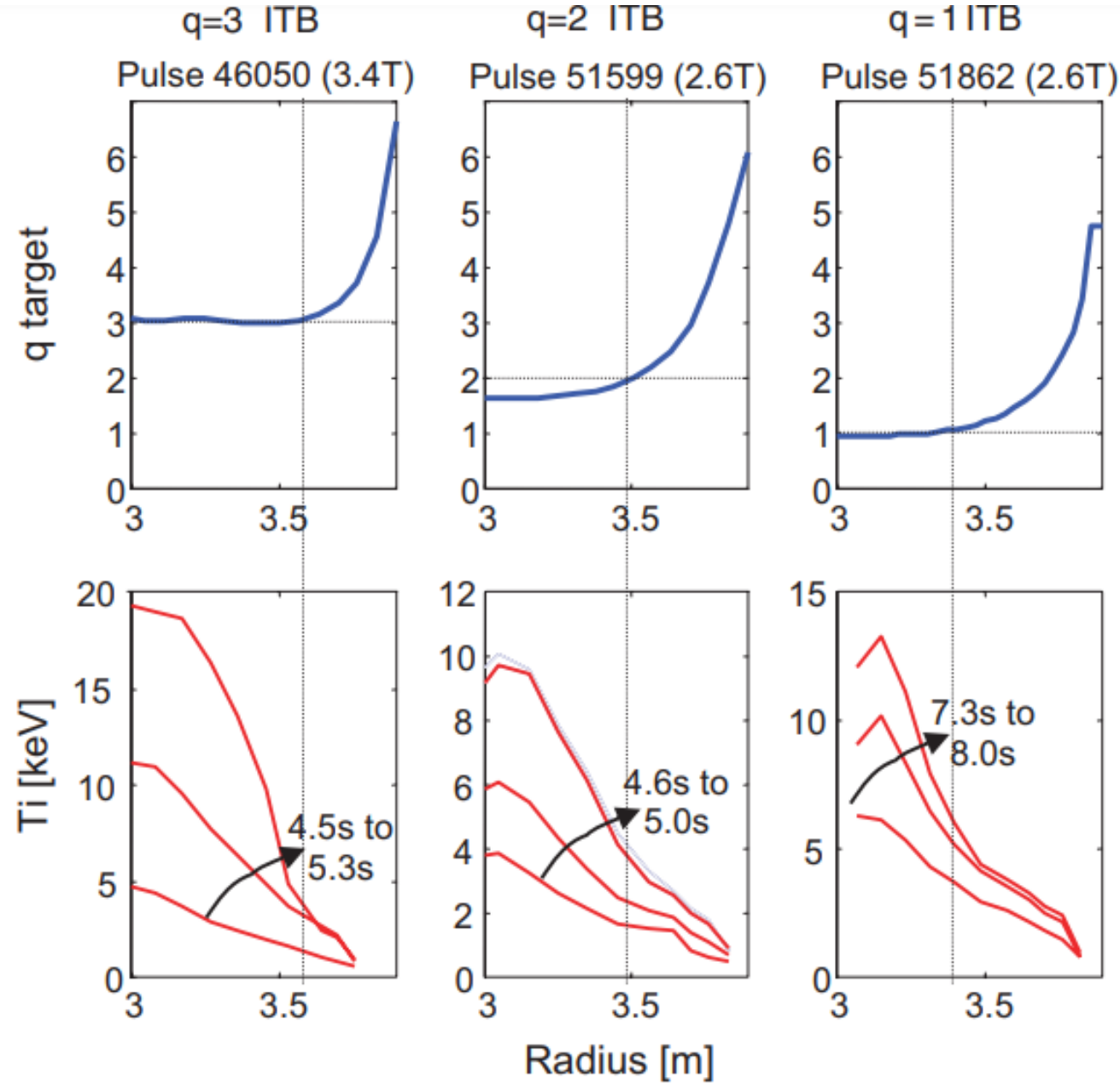
Additionally:

- Heating power threshold

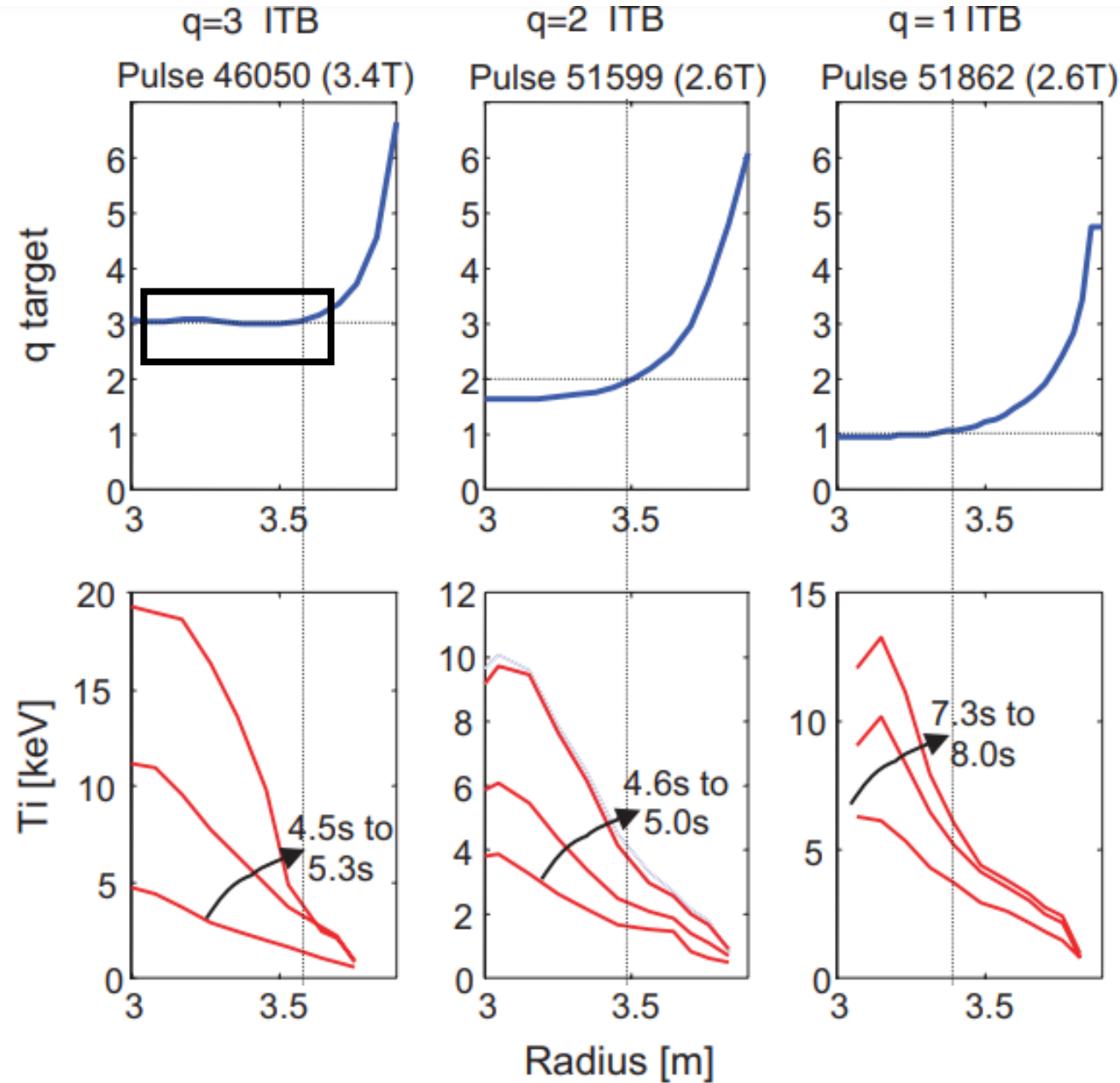


References:

- K Ida and T Fujita 2018 Plasma Phys. Control. Fusion 60 033001
- J.W. Connor et al. 2004 Nucl. Fusion 44 R1
- X. Garbet et al. 2010 Nucl. Fusion 50 043002



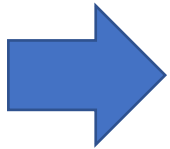




D4.1 Quantification of ITB momentum drive from rational vs irrational surfaces and comparisons to plasma edge	Target date 02/2022
M4.1 Quantify momentum drive from rational vs irrational surfaces in ITBs and compare to momentum drive at plasma edge and determine relationship of parallel correlation length with magnetic shear.	Target date 12/2021

D4.1 Quantification of ITB momentum drive from rational vs irrational surfaces and comparisons to plasma edge

Target date  
02/2022



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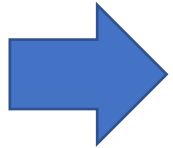
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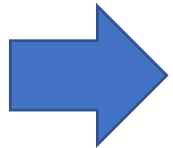


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The proposed plan has not changed substantially and we are on track, working towards the milestones.

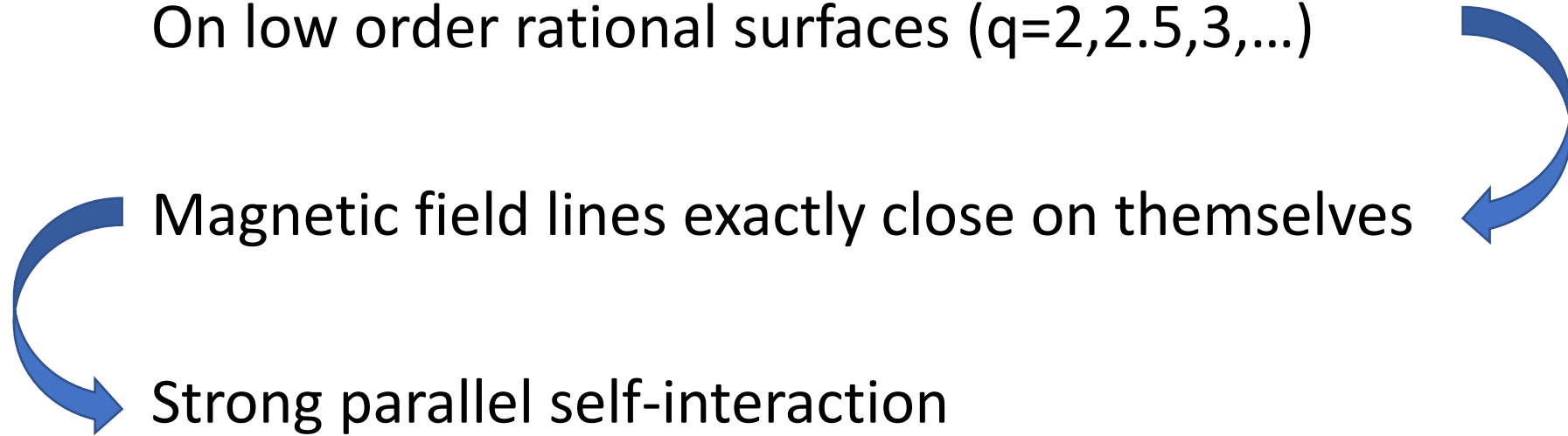
# Internal transport barrier investigation in local gyrokinetic simulations

# Turbulent self-interaction

On low order rational surfaces ( $q=2,2.5,3,\dots$ )

Magnetic field lines exactly close on themselves

Strong parallel self-interaction



## References:

- J. Ball *et al.* 2020 *Journal of Plasma Physics* **86(2)**, 905860207
- Ajay CJ, Studying the effect of non-adiabatic passing electron dynamics on microturbulence self-interaction in fusion plasmas using gyrokinetic simulations, Thesis EPFL Lausanne, 2020

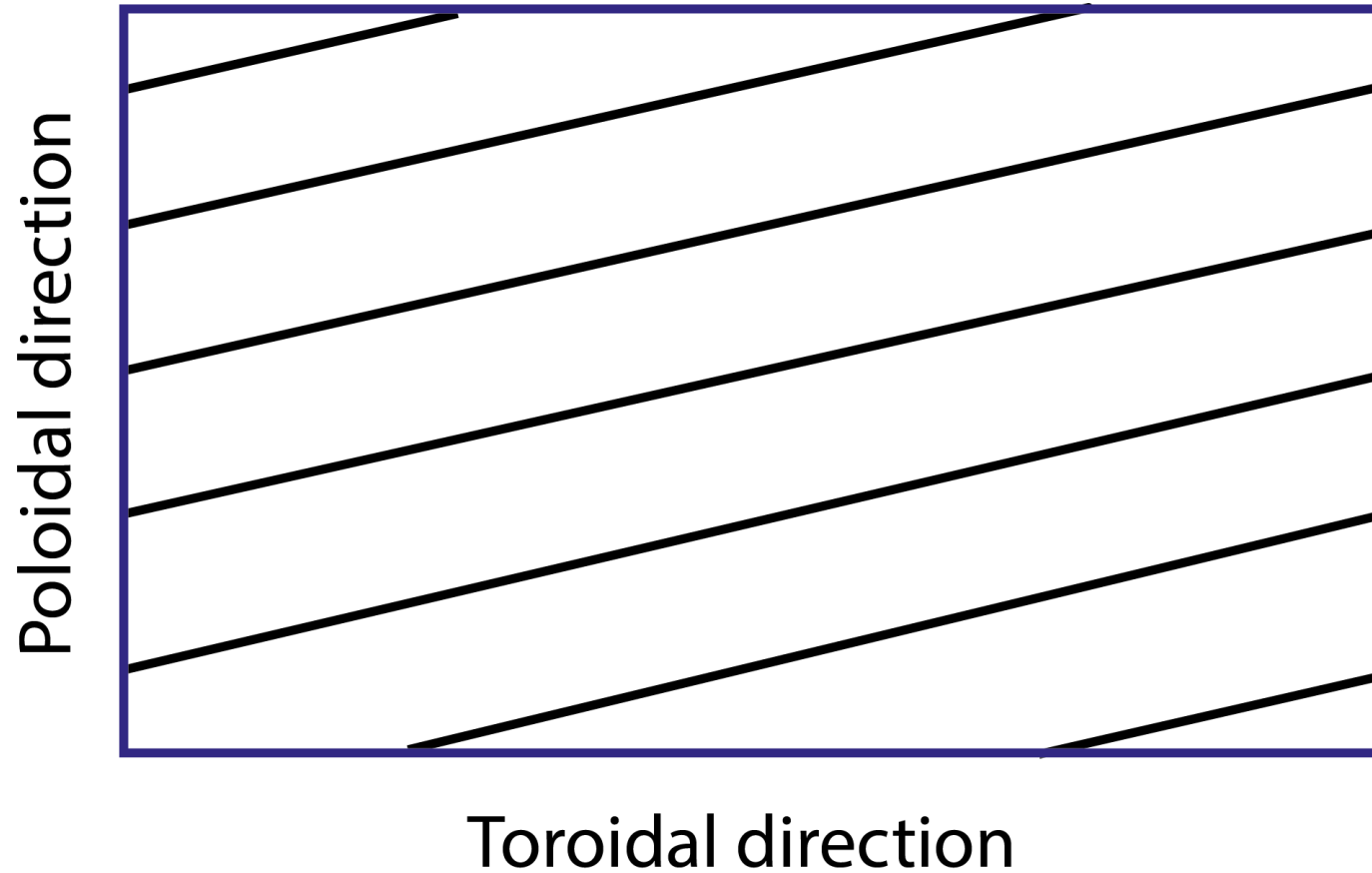


# EPFL Self-interaction

- Self-interaction alters fluctuation behaviour both linearly and non-linearly
- Turbulent self-interaction can be visualized **in real space as “eddy biting its own tail”**

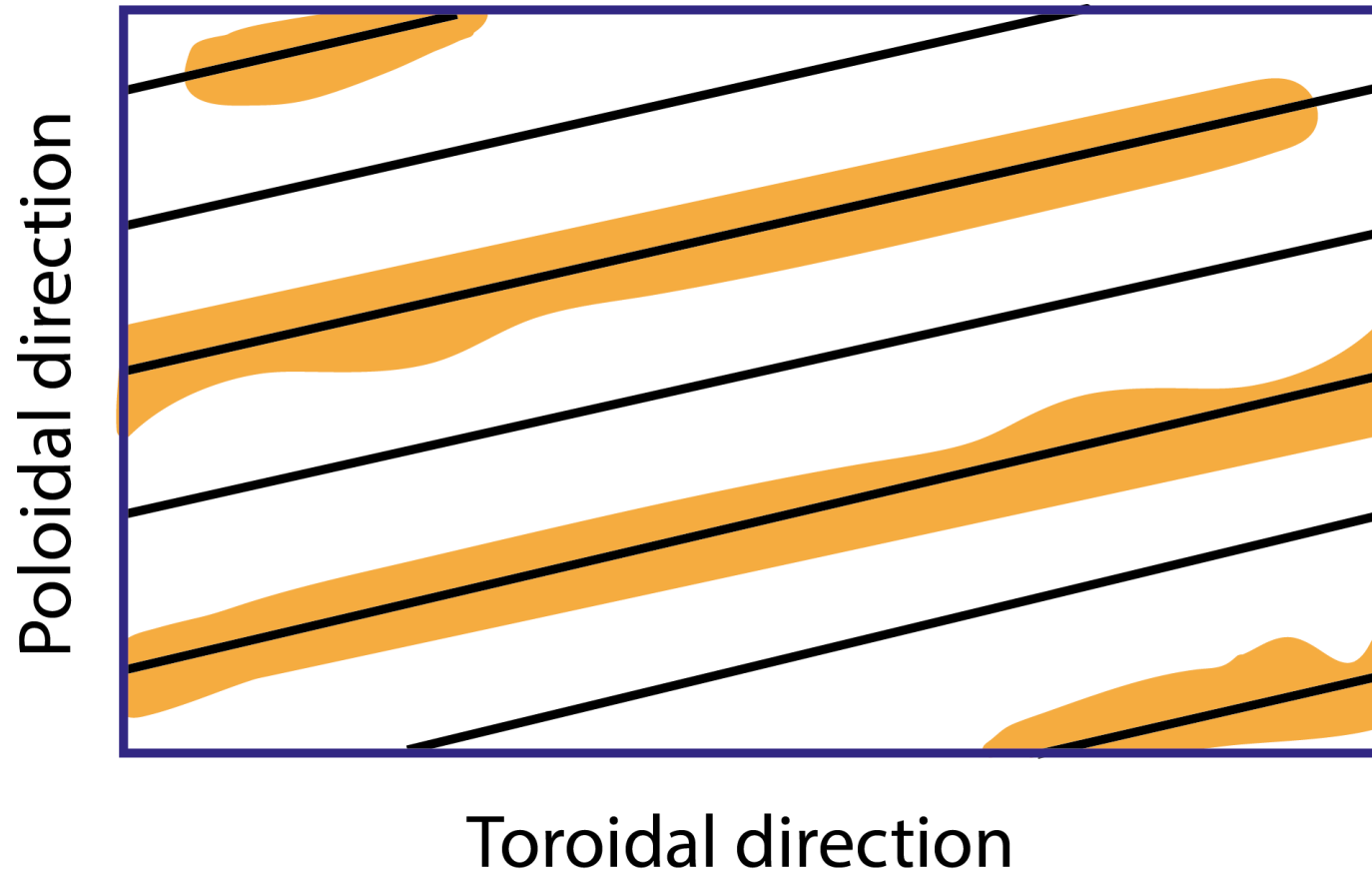
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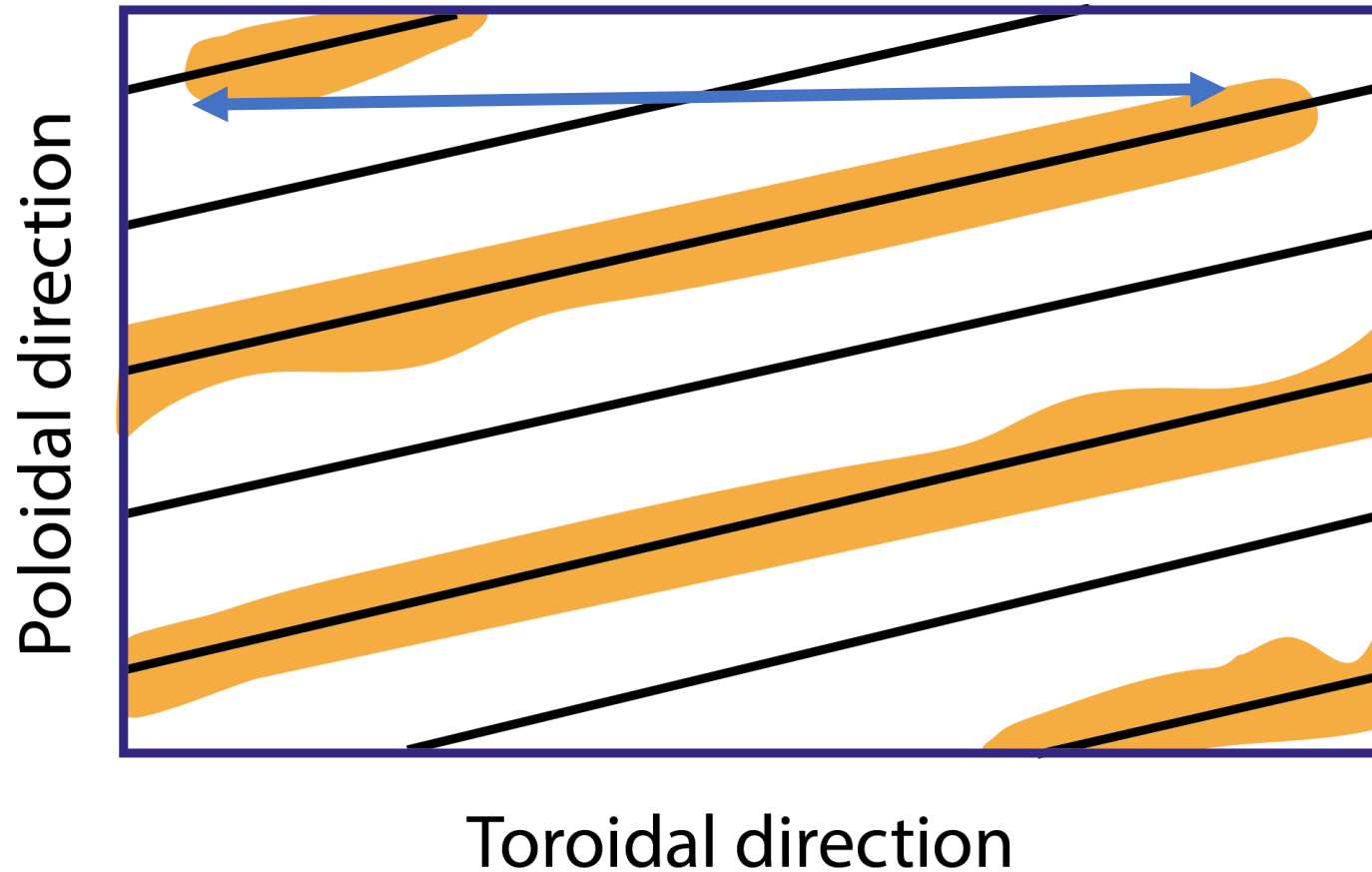


$$q = 2.5$$

# EPFL Self-interaction

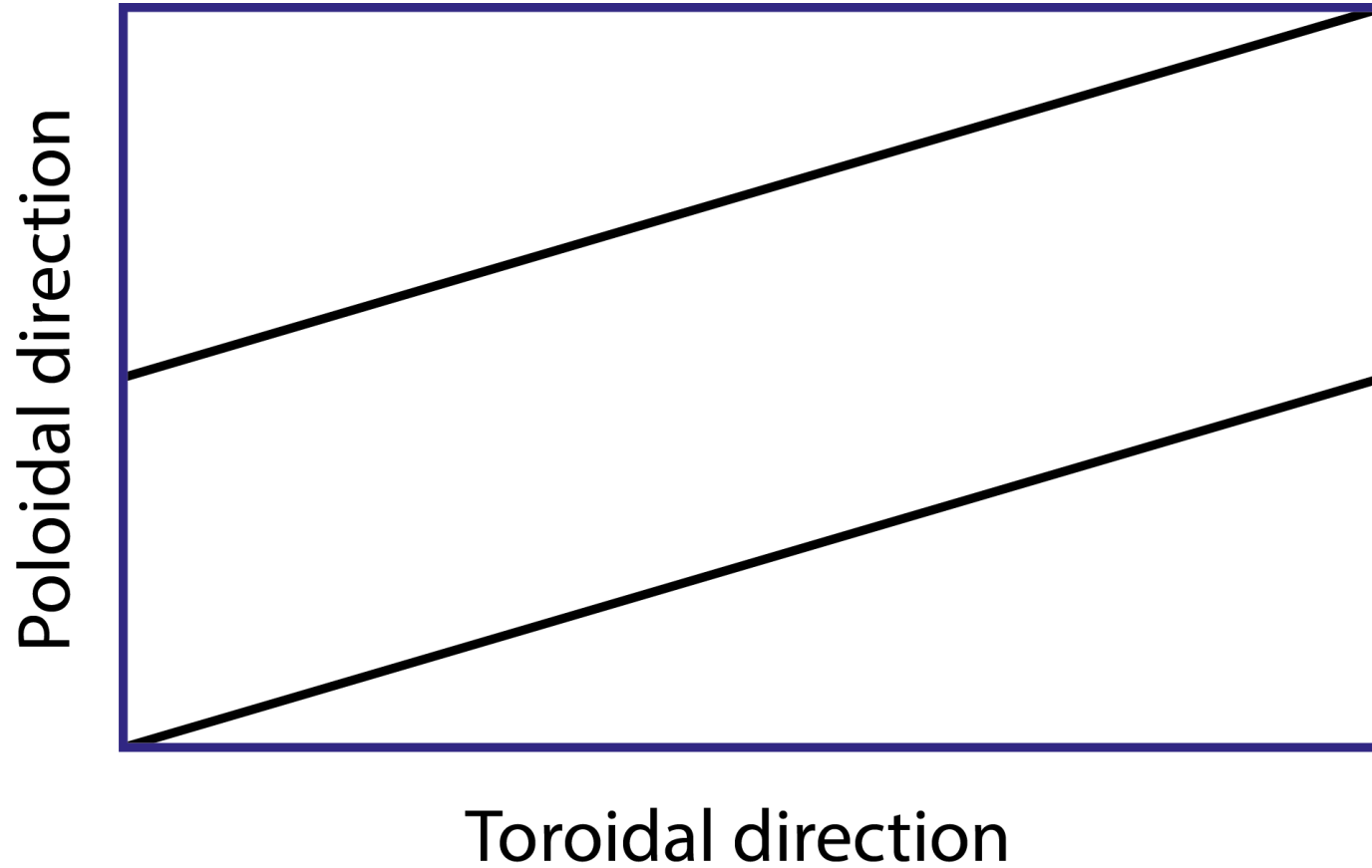


$$q = 2.5$$



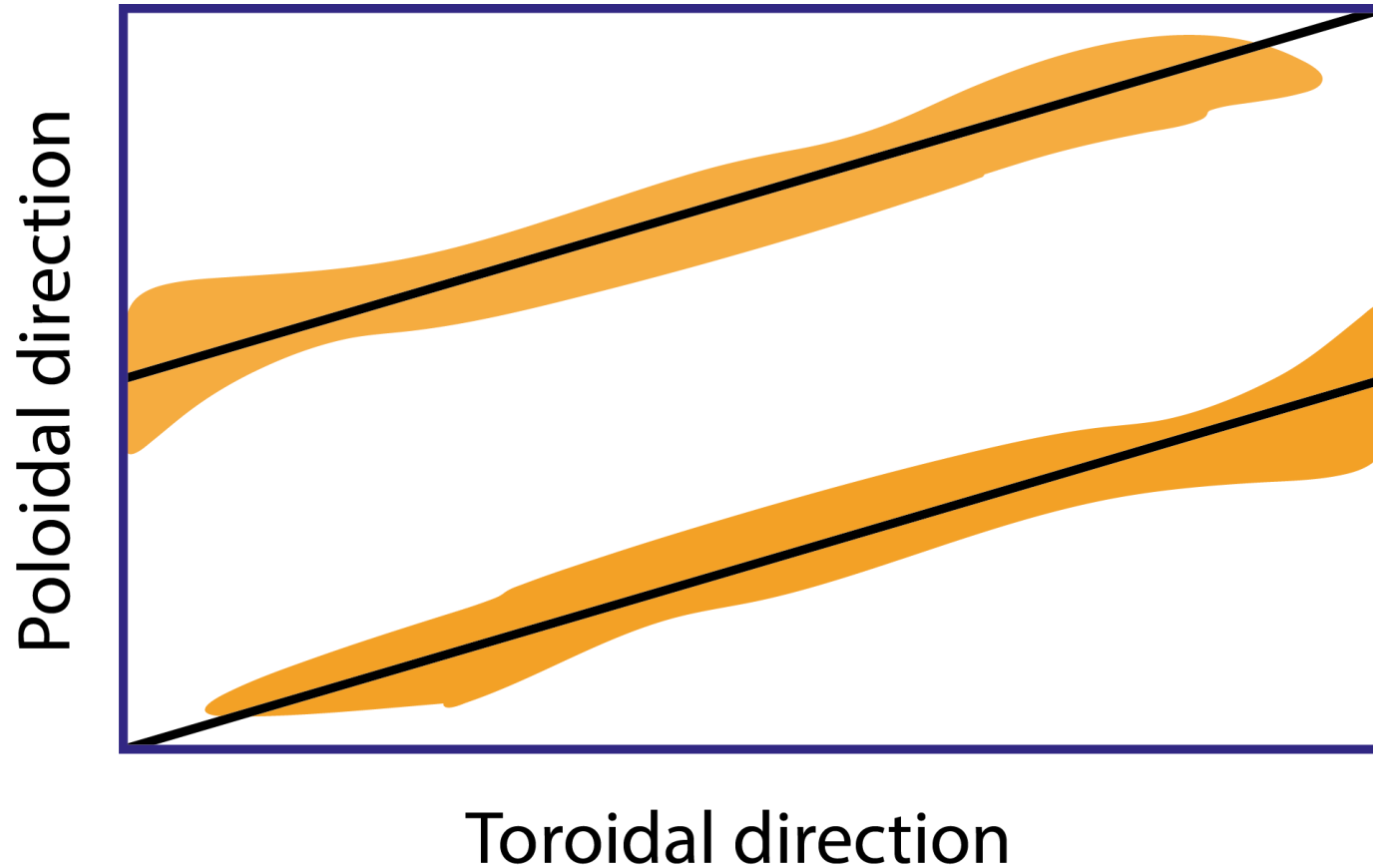
$$q = 2.5$$

# EPFL Self-interaction



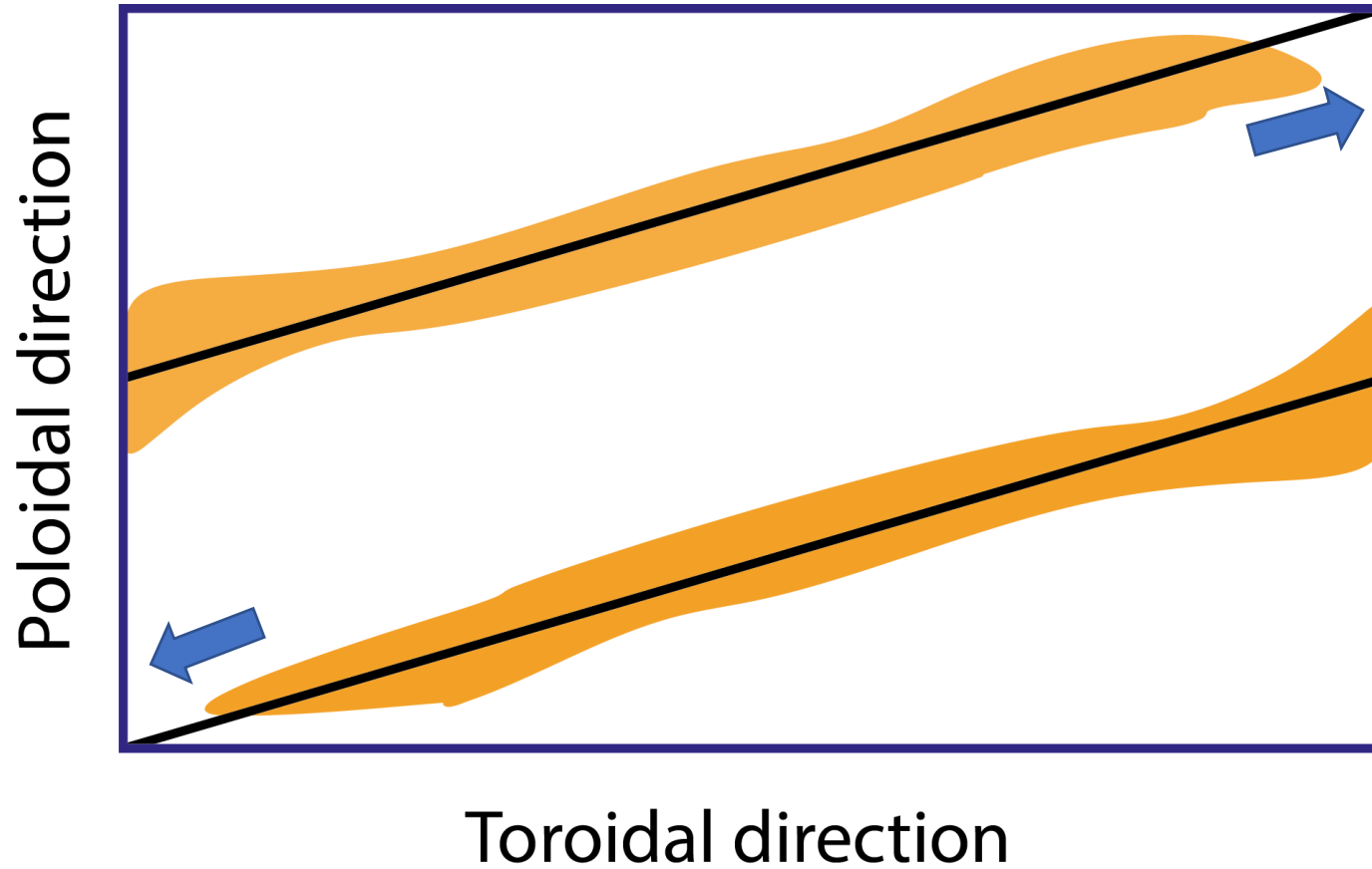
$$q = 2$$

# EPFL Self-interaction



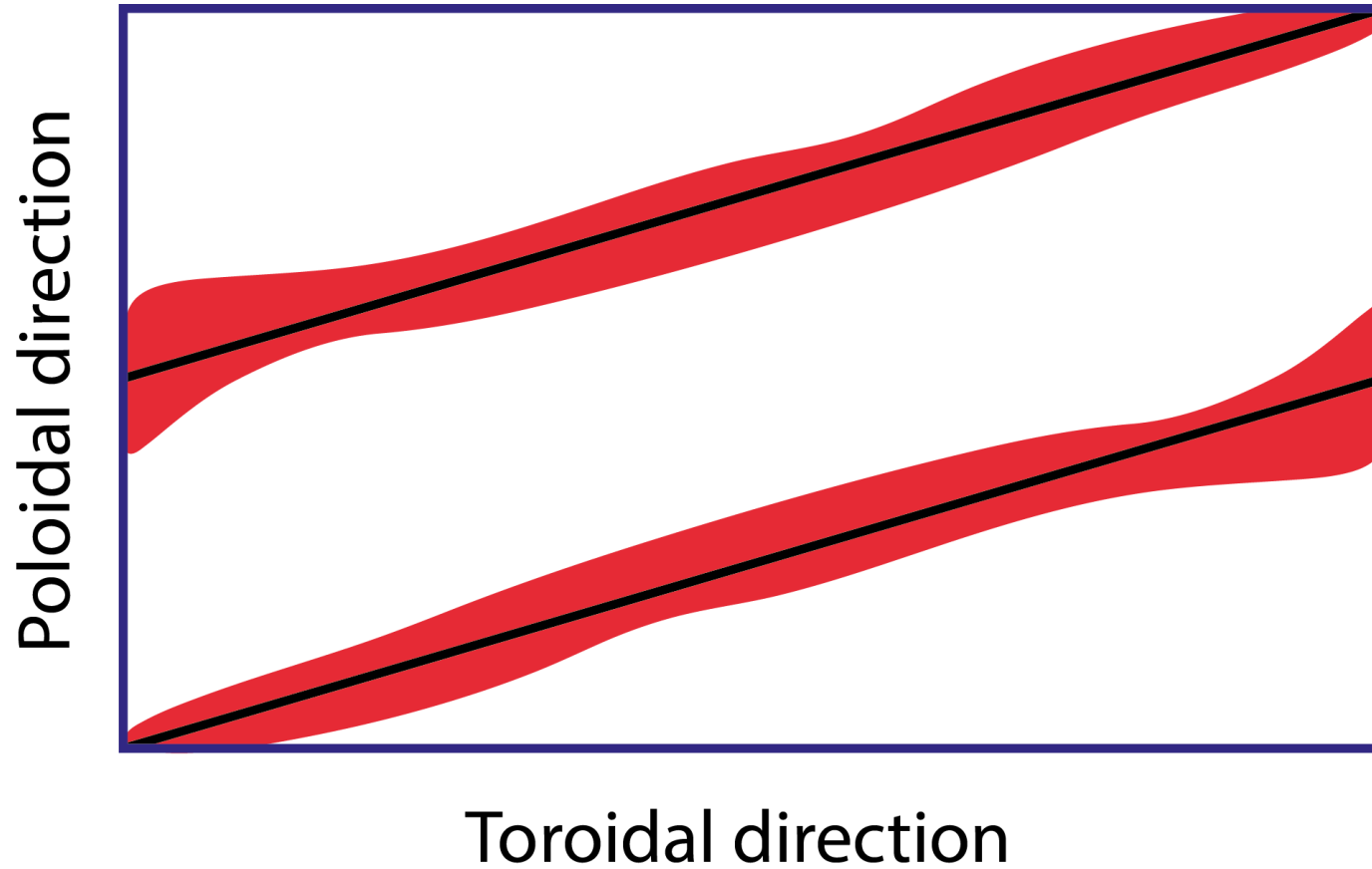
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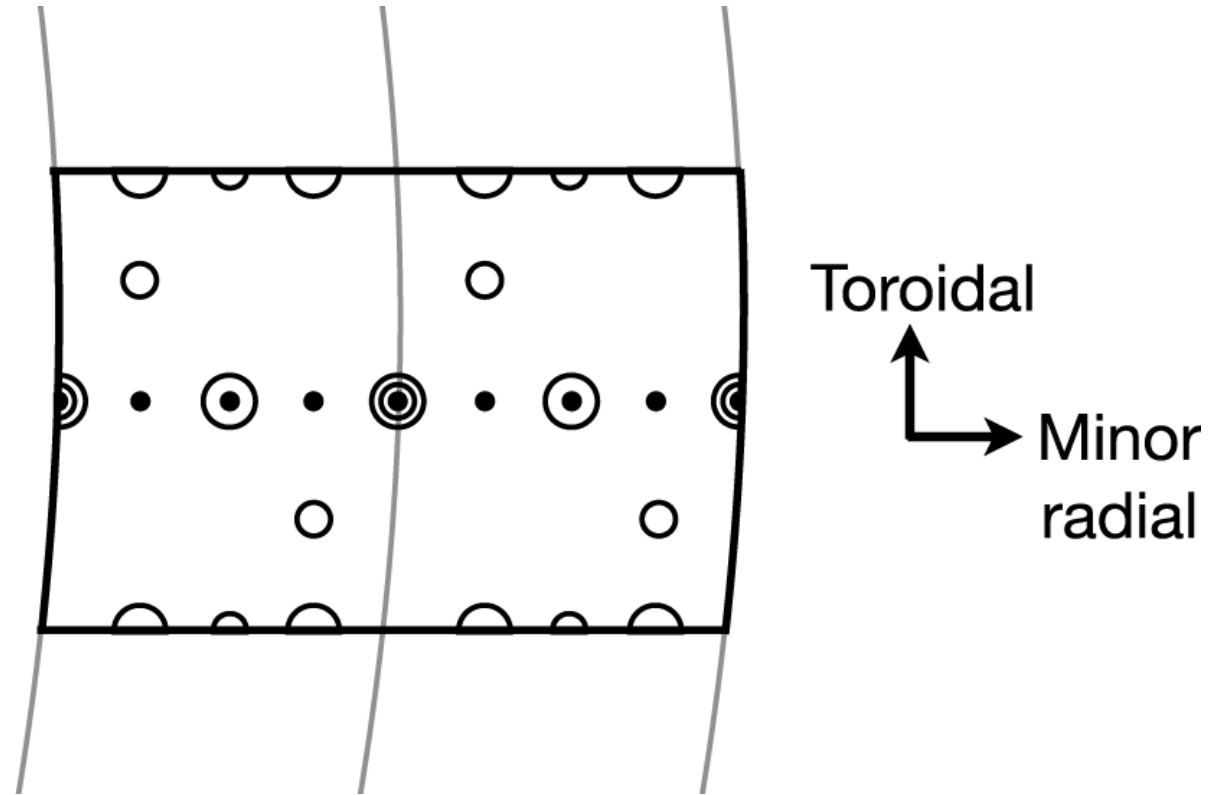


$$q = 2$$



# EPFL Flux-tube

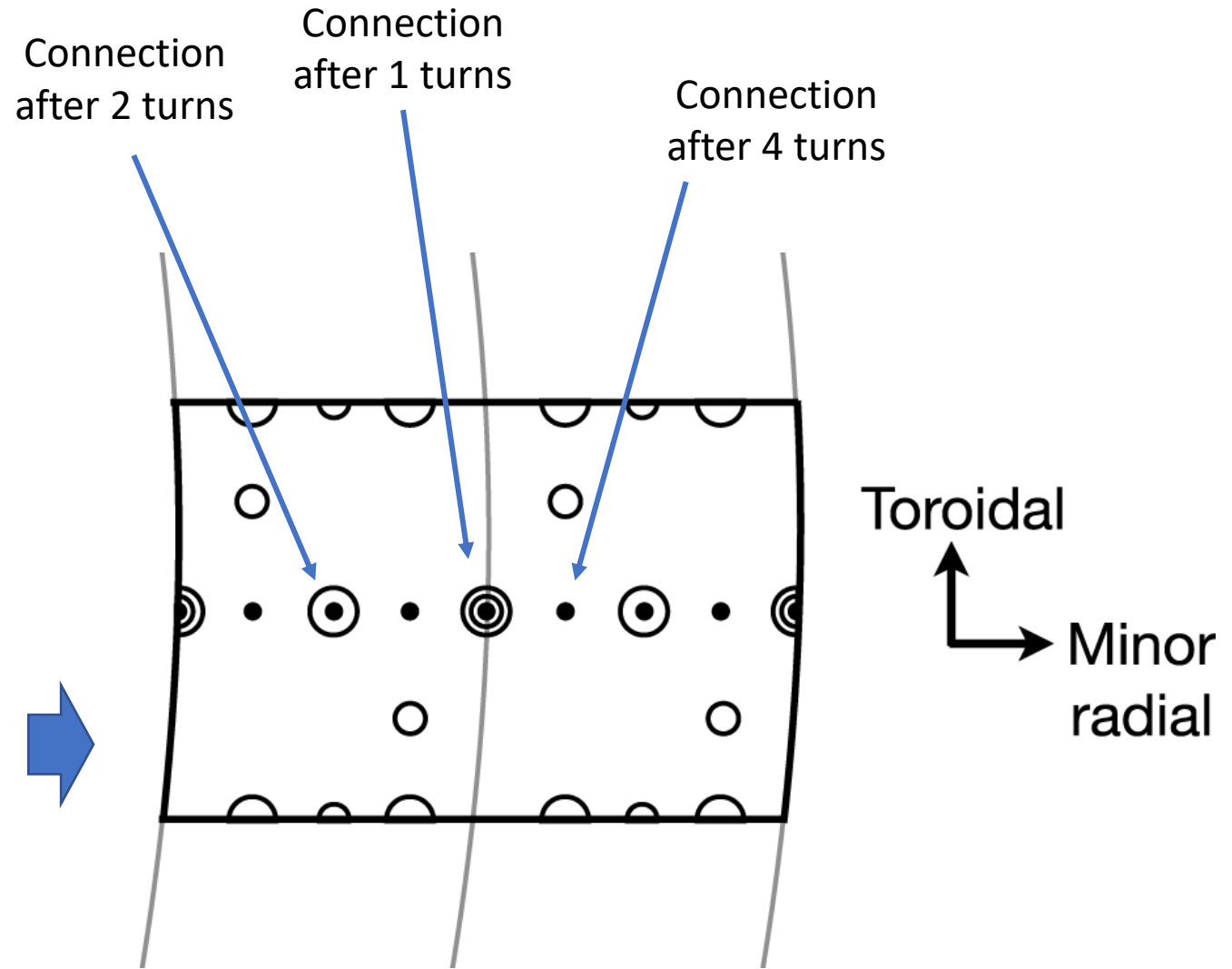
- Simulations using local flux-tube GENE code (Eulerian  $\delta f$  code)
- Twist and shift parallel boundary condition -> special radial locations



J. Ball *et al.* 2020 *Journal of Plasma Physics* **86(2)**, 905860207

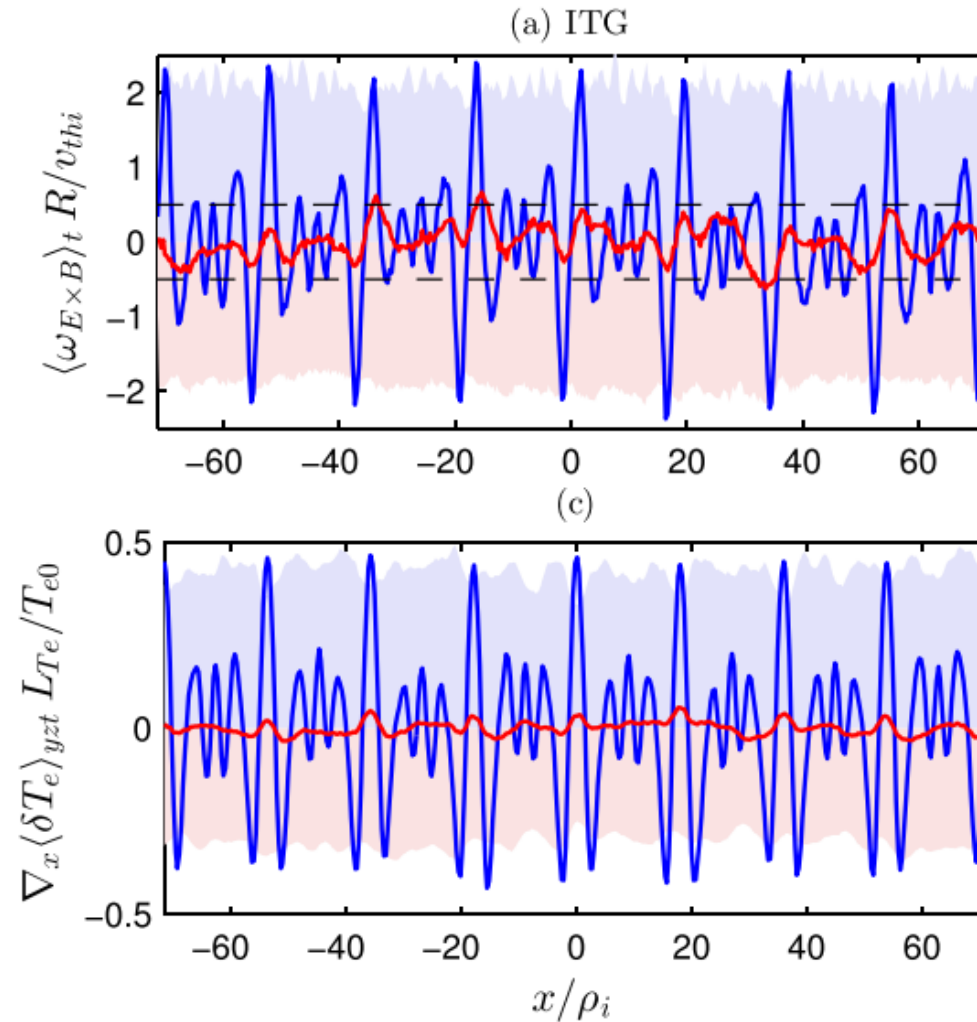
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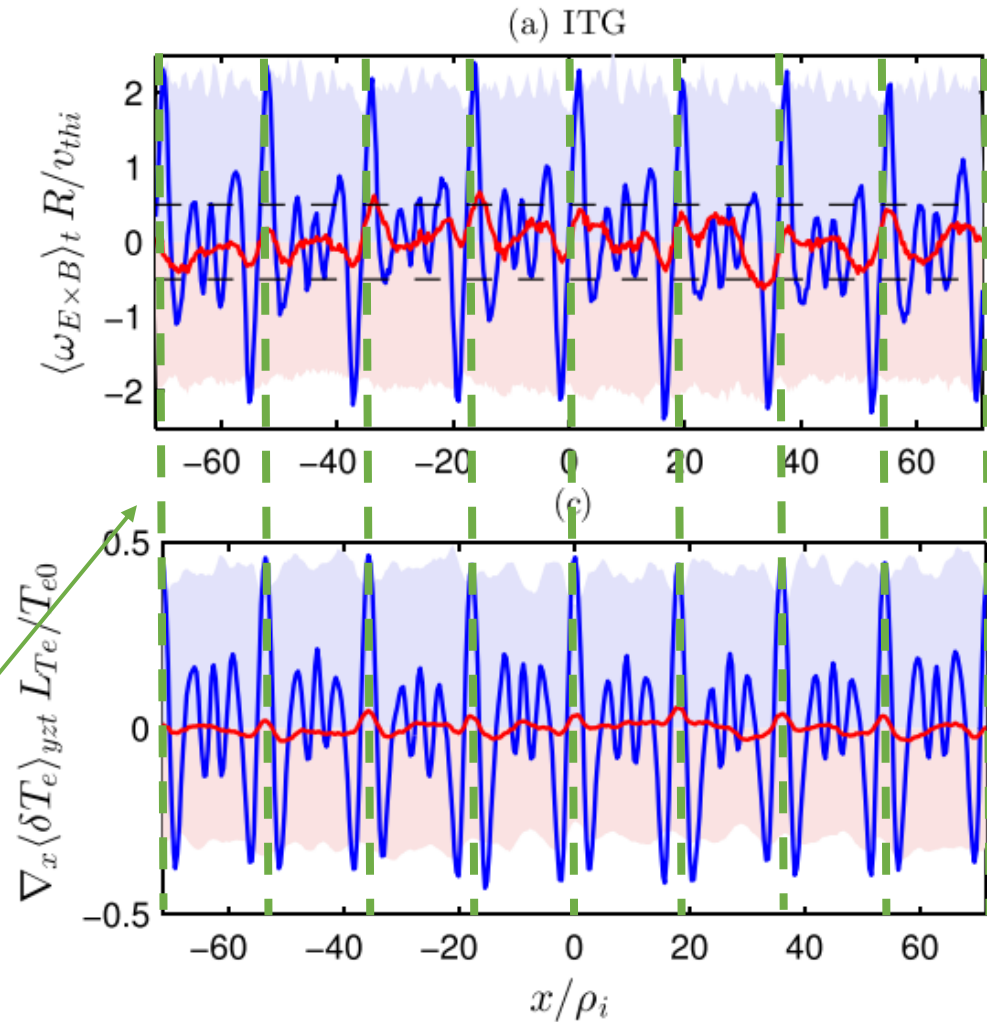
Drives ExB shear flow  
around rational surfaces



J. Dominski *et al.* 2015 *Physics of Plasmas* **22**, 062303

Drives ExB shear flow  
around rational surfaces

Lowest order rational  
surfaces



J. Dominski *et al.* 2015 *Physics of Plasmas* **22**, 062303

# EPFL ITB triggering

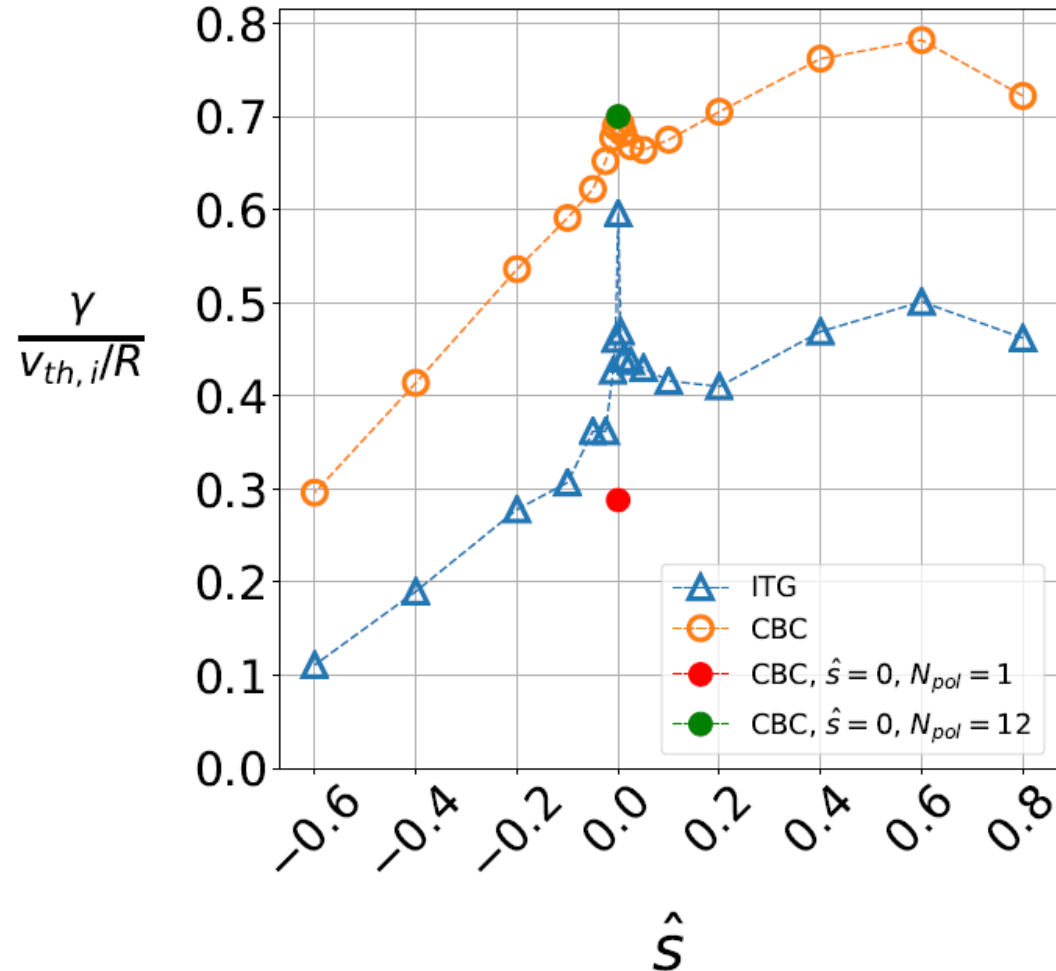
- Low magnetic shear important for ITB formation
- Integer (or low order rational) surfaces important for ITB formation
- Turbulent self-interaction strongest around rational surfaces
- Turbulent self-interaction seems to be stabilizing



Low magnetic shear + self-interaction = ITB ?

# EPFL Preliminary study

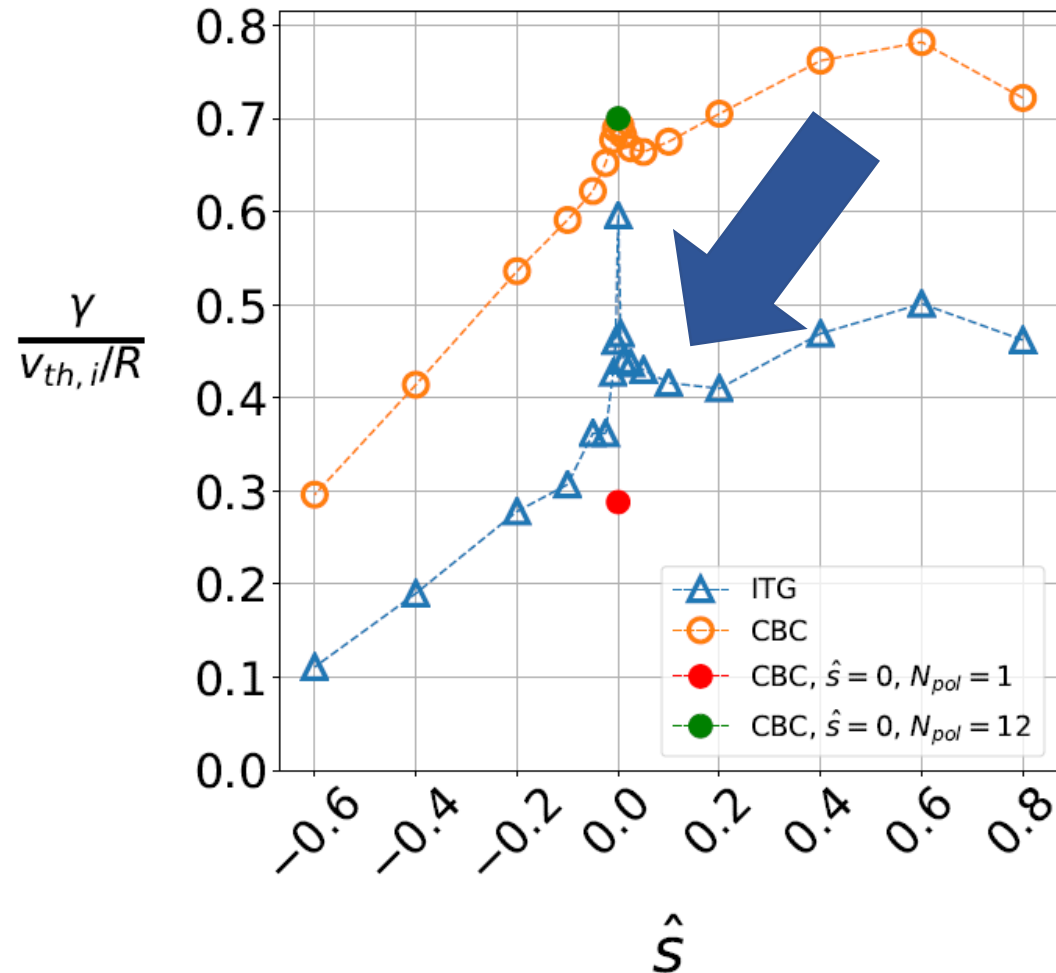
- Linear low and zero shear simulations
- Nonlinear low and zero shear simulations
- Simulations with **kinetic electrons**
- Starting point – Cyclone Base Case (CBC) parameters



We believe we observe **transition from toroidal to slab ITG mode** as magnetic shear approaches zero.

$$k_y \rho_i = 0.45$$

# Linear study with kinetic electrons

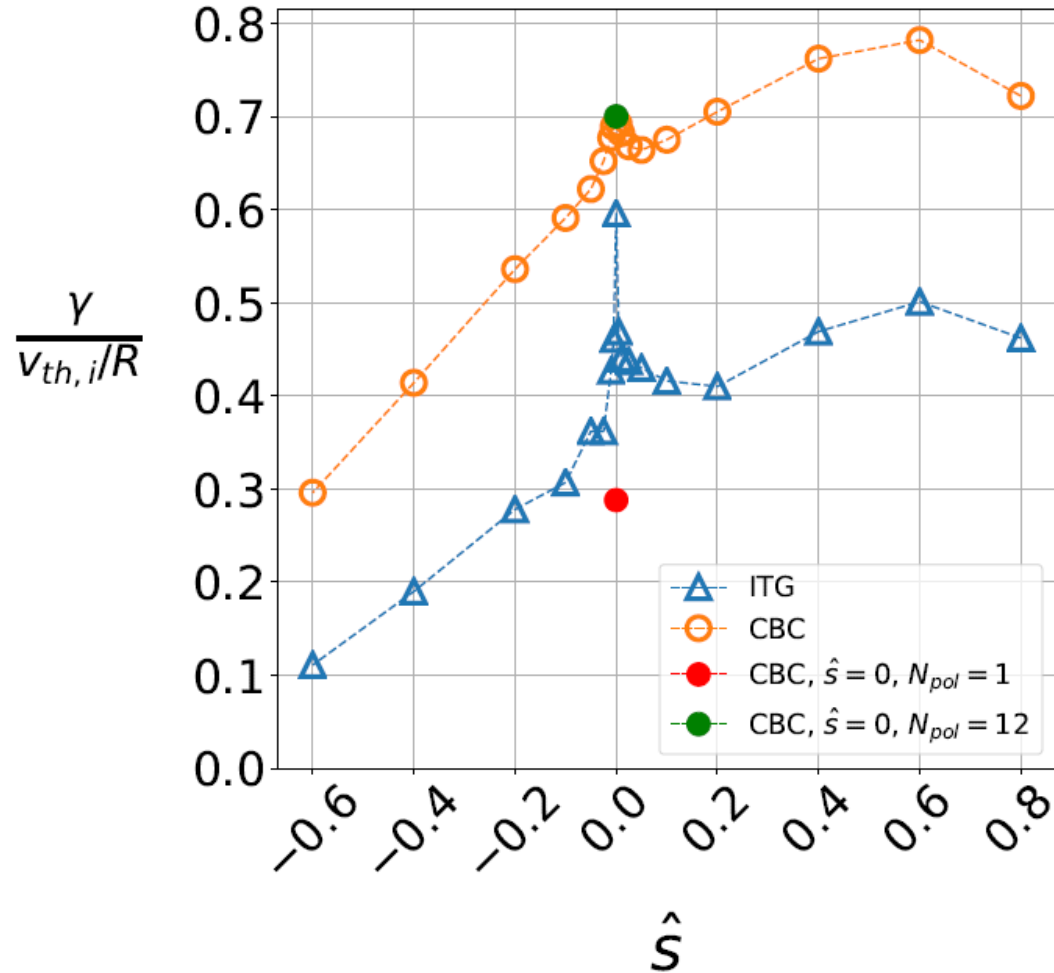


Slab ITG to toroidal ITG for purely ITG drive

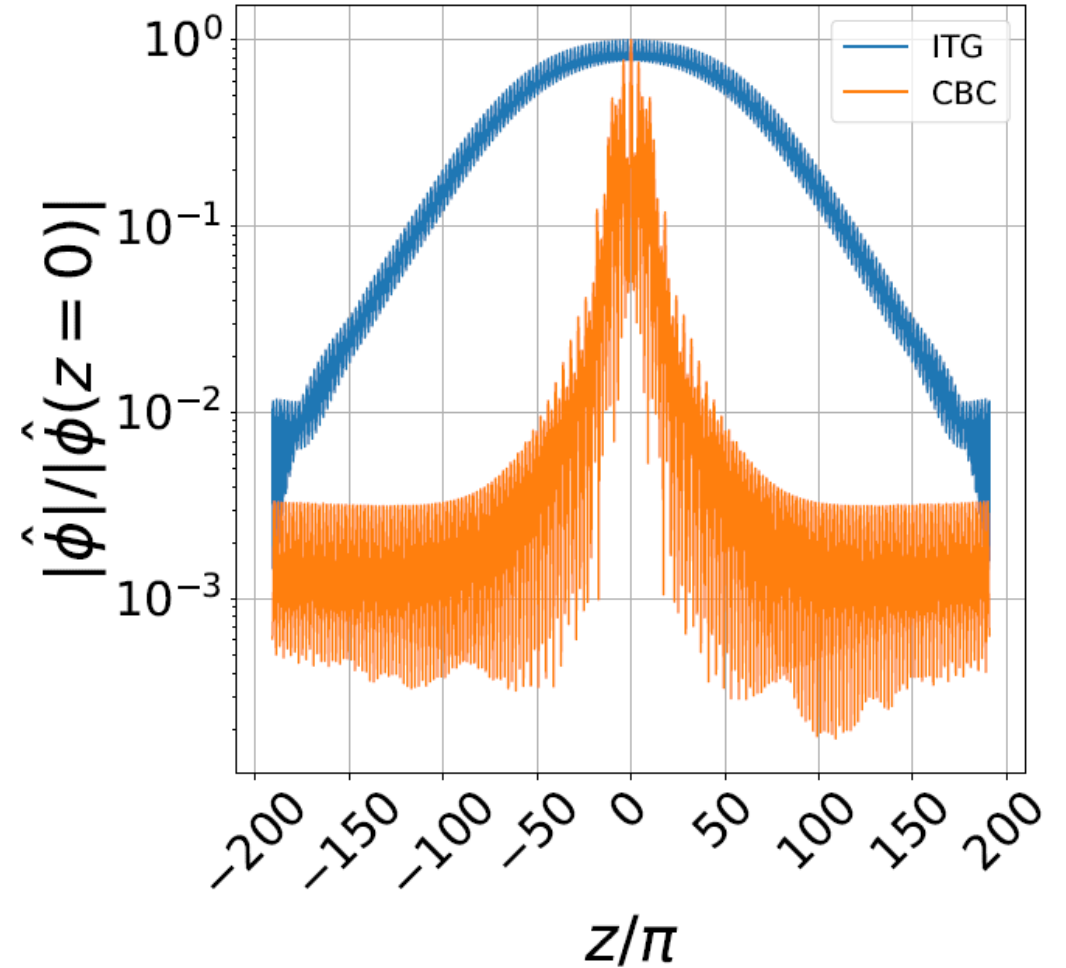
$k_y \rho_i = 0.45$



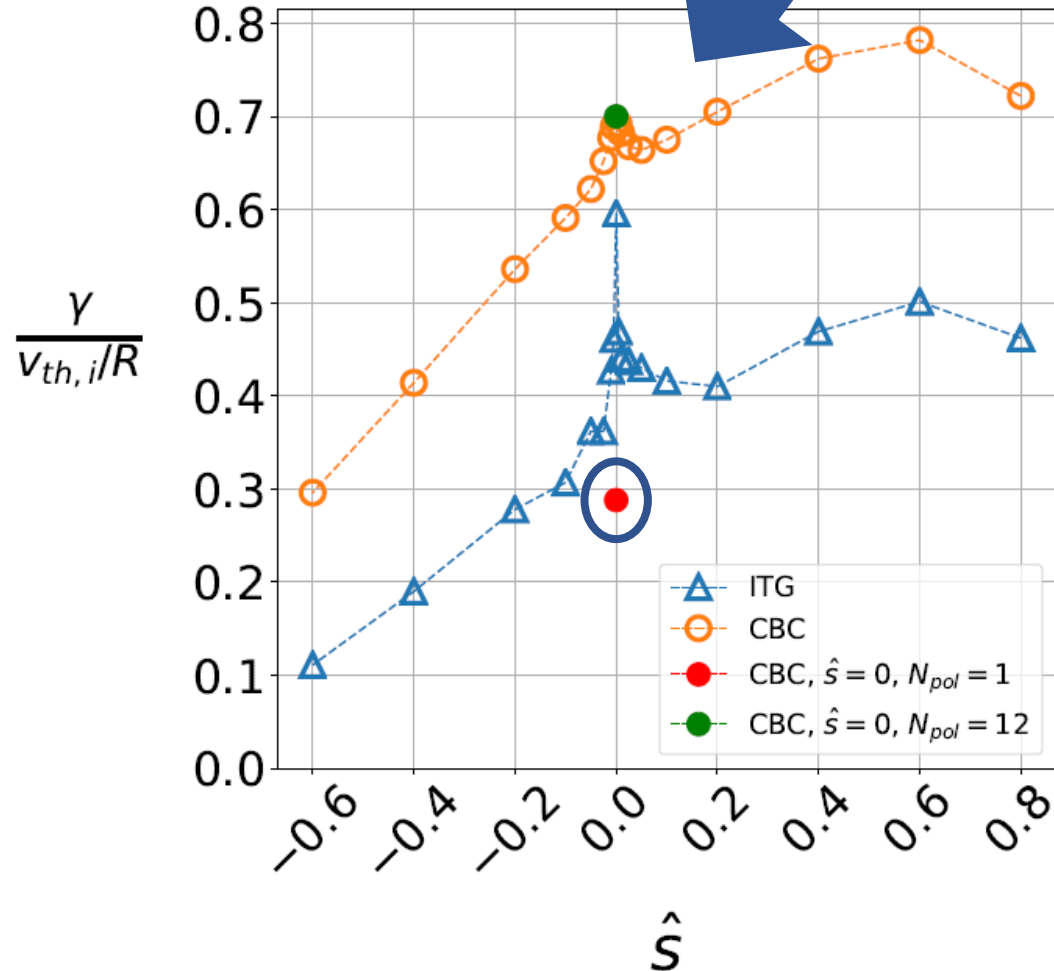
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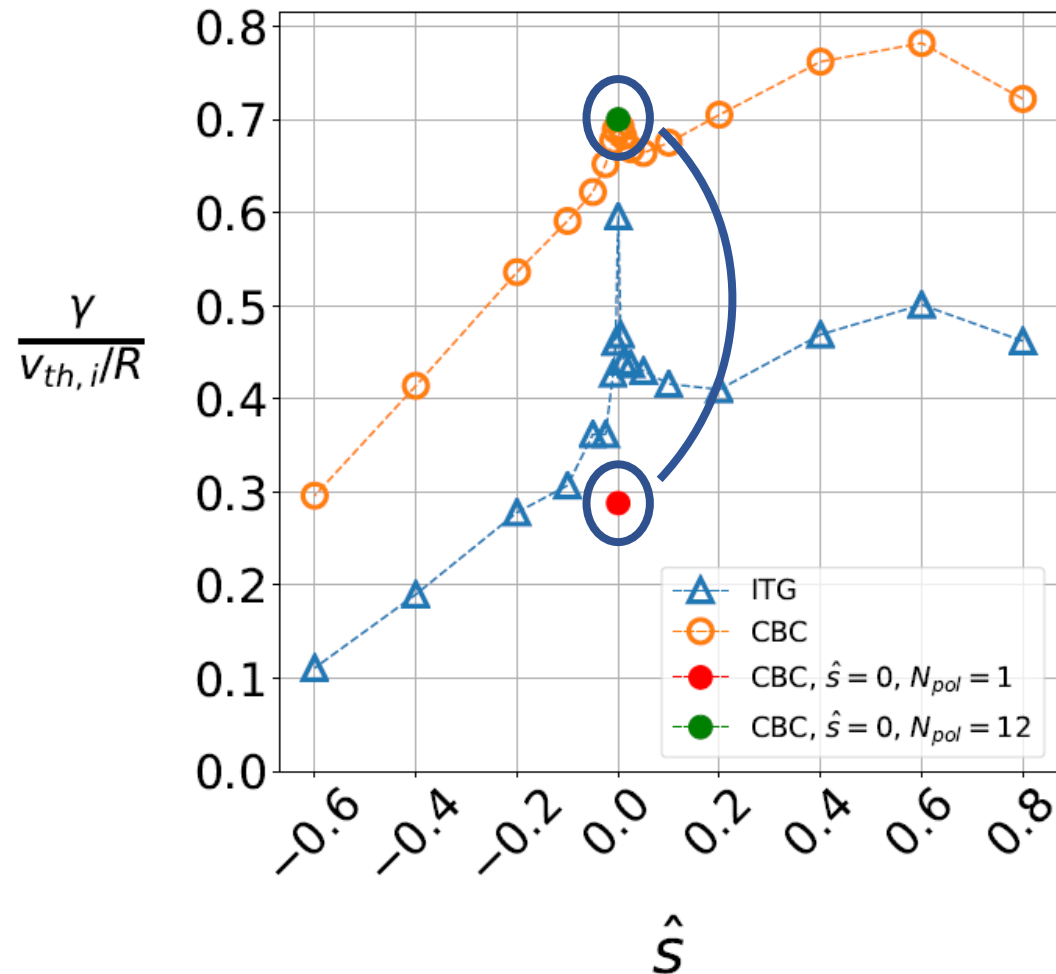
# Linear study with kinetic electrons



With CBC drive discontinuity at  $s=0$

Toroidal ITG growth rate reduced by self-interaction at  $s=0$

$k_y \rho_i = 0.45$



Continuity recovered if number of poloidal turns is increased

Number of poloidal turns does not affect purely ITG driven  $s=0$  simulation

$$k_y \rho_i = 0.45$$

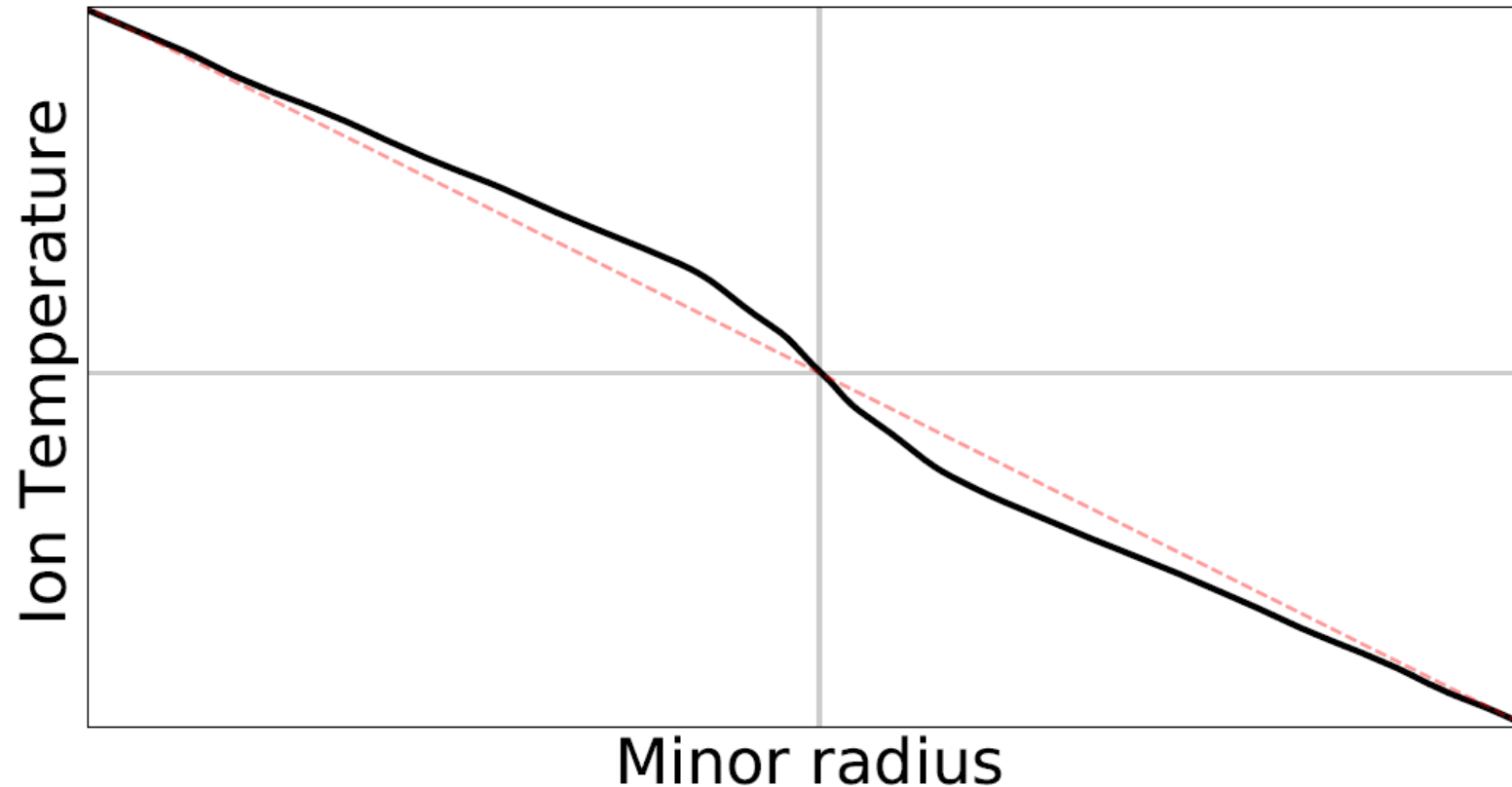
# Key linear results

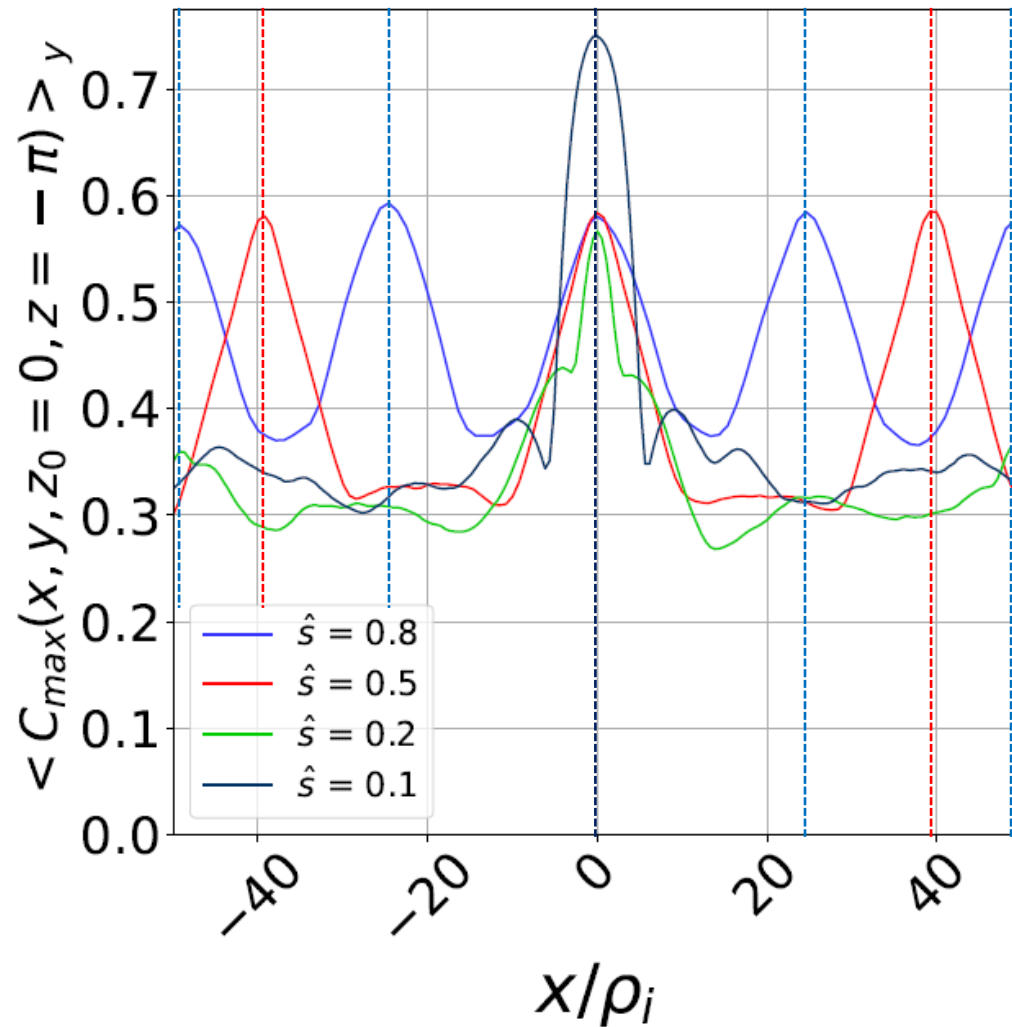
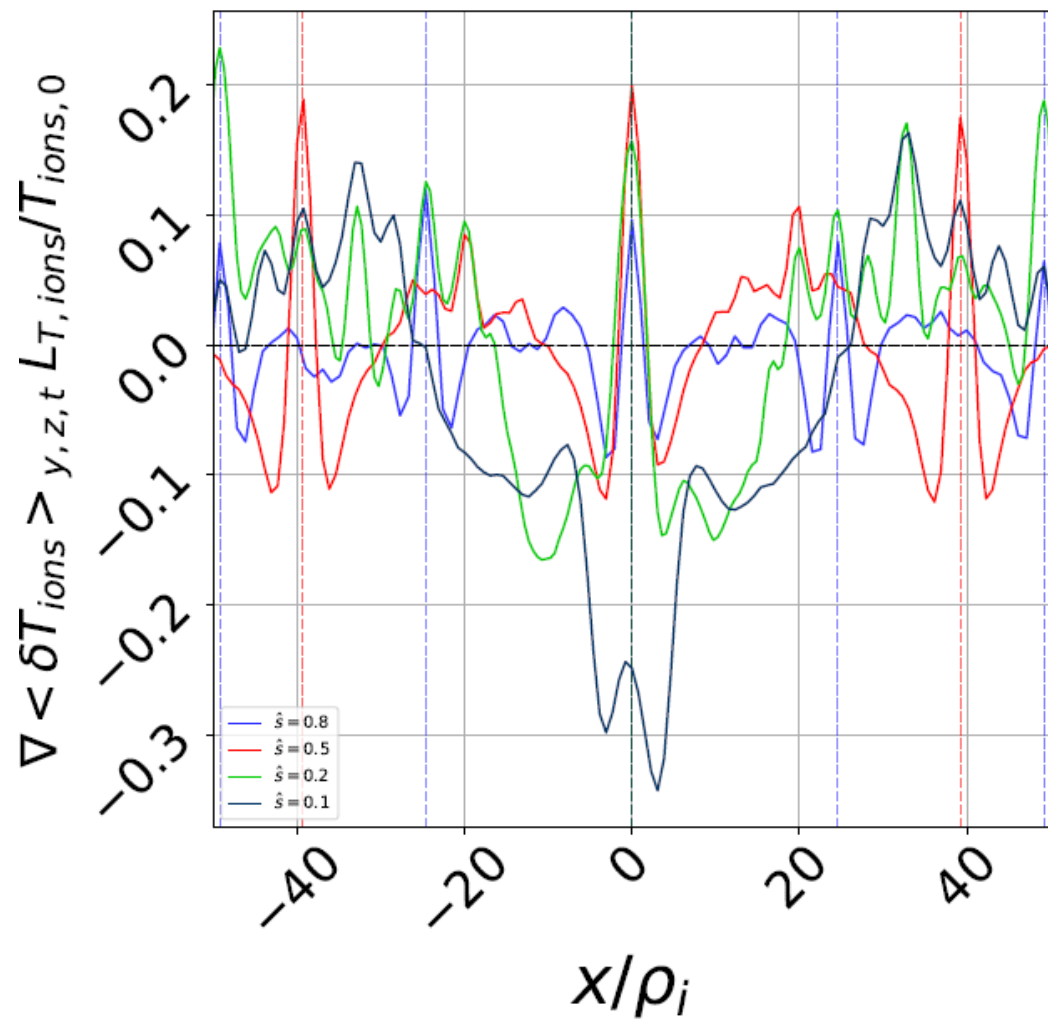
**So far the main takeaway from linear studies is that at low magnetic shear there can be a transition to slab ITG.**

- The linear growth rate of toroidal ITG was strongly reduced by self-interaction unlike slab ITG;
- Slab ITG extends further along magnetic field lines;
- This transition could be very important for turbulent self-interaction.

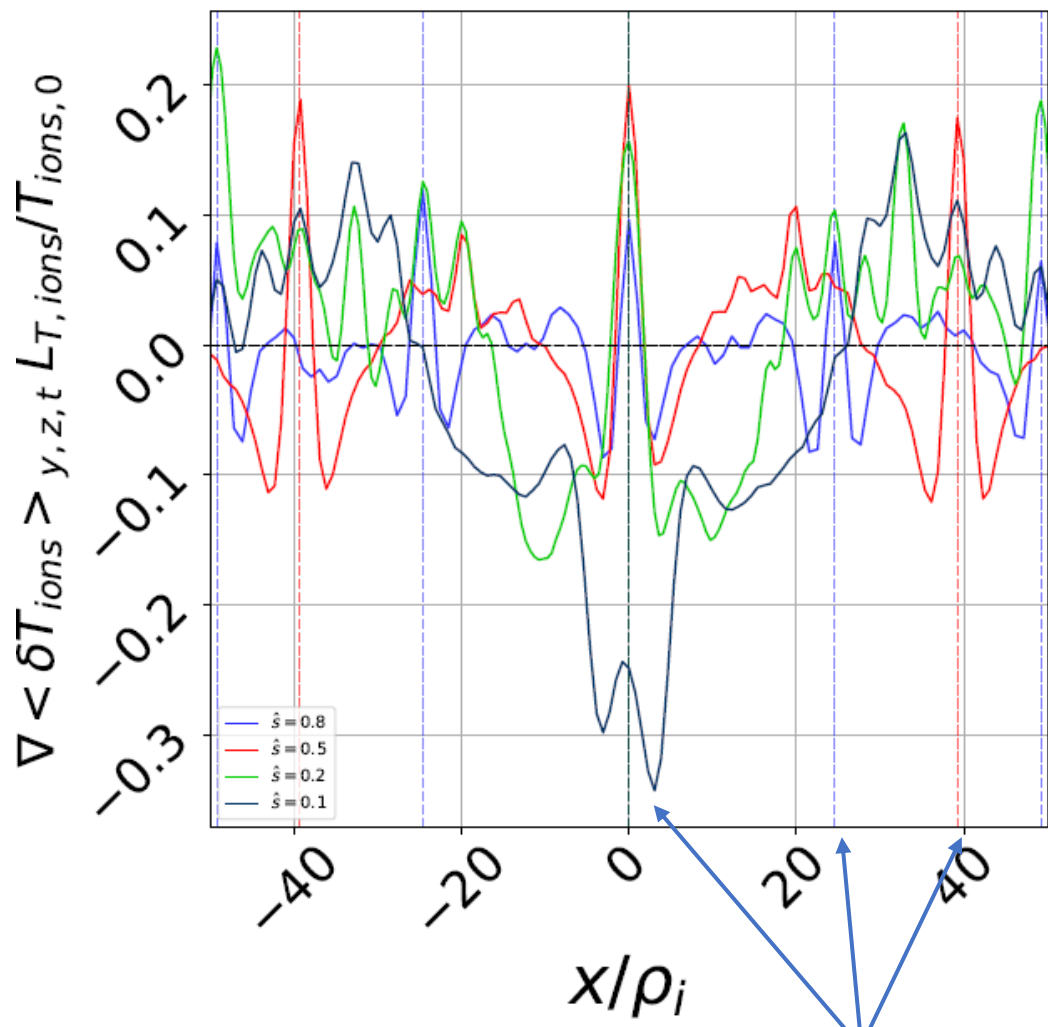
# EPFL Nonlinear study

For  $s=0.1$  we see a **strong corrugation in the plasma profiles** when compared to background gradients

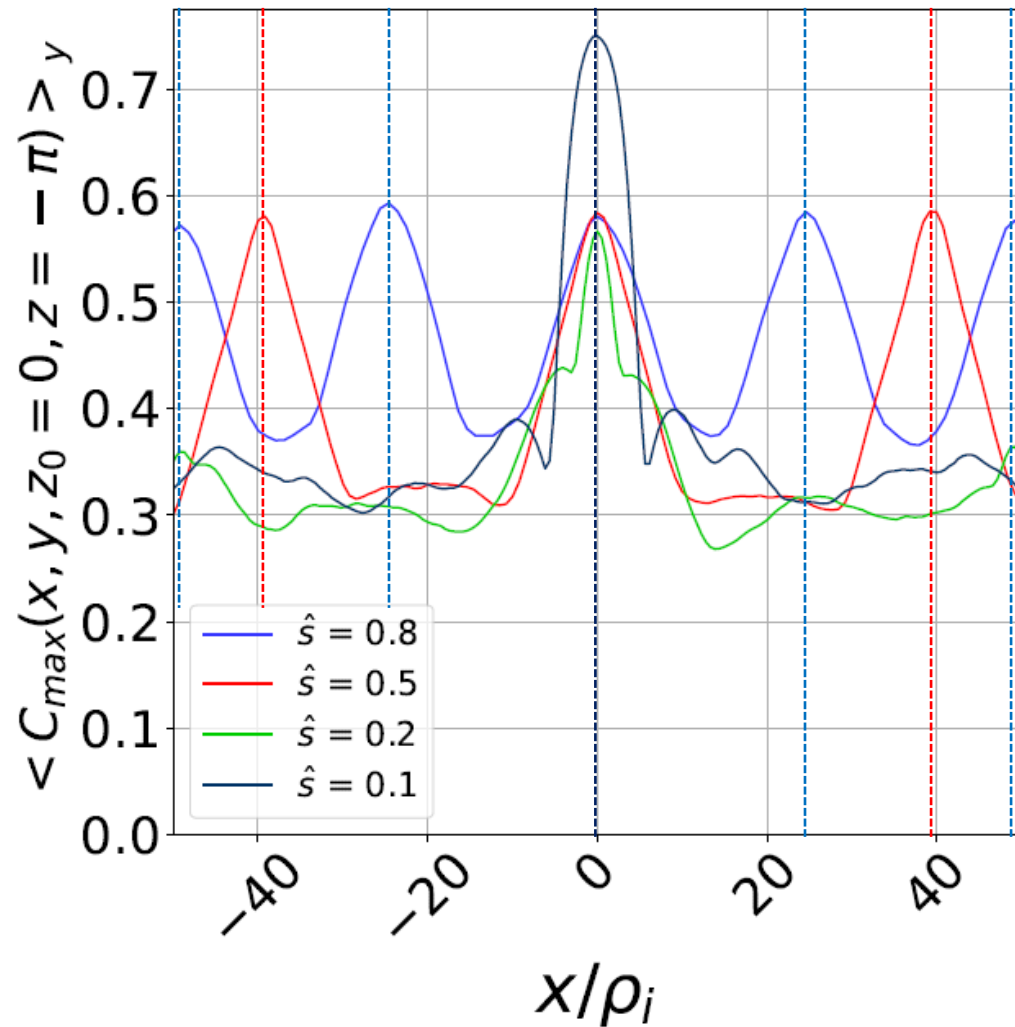




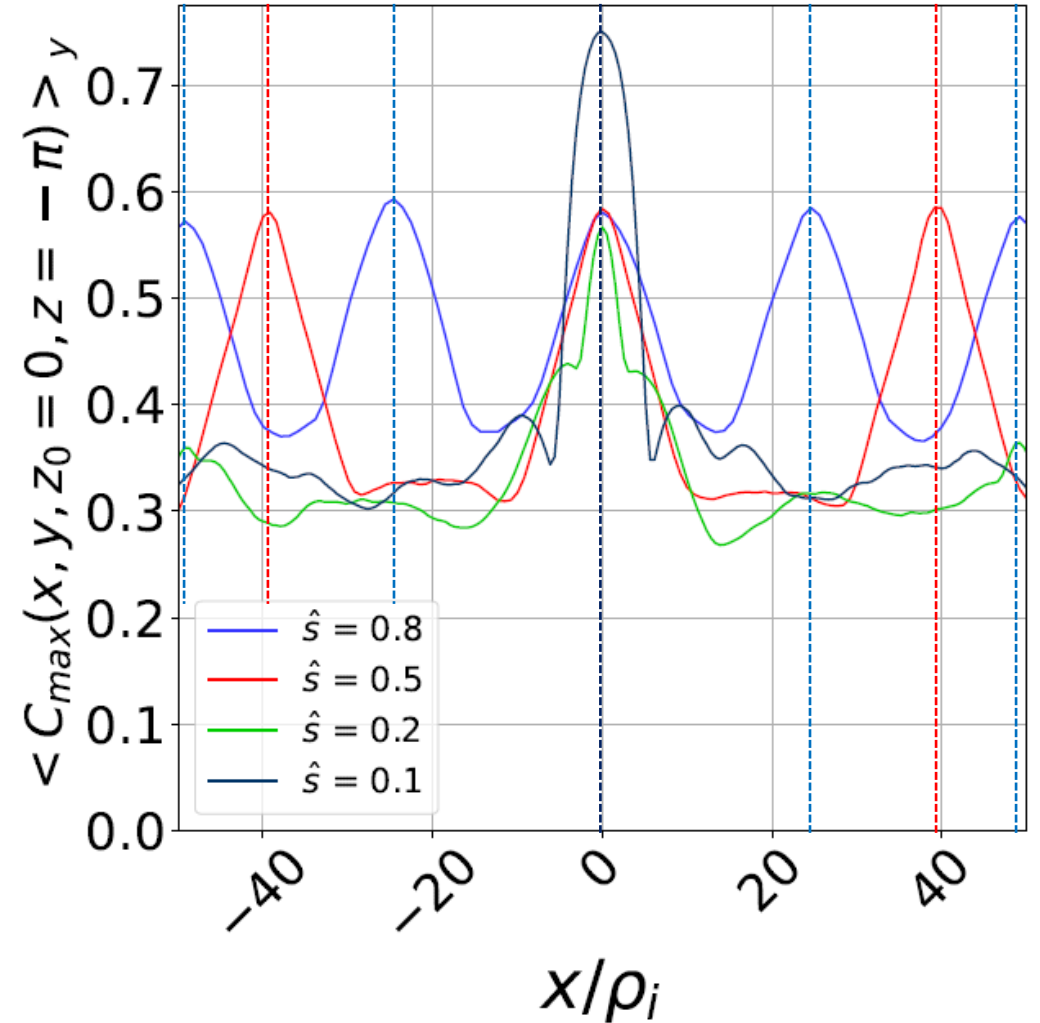
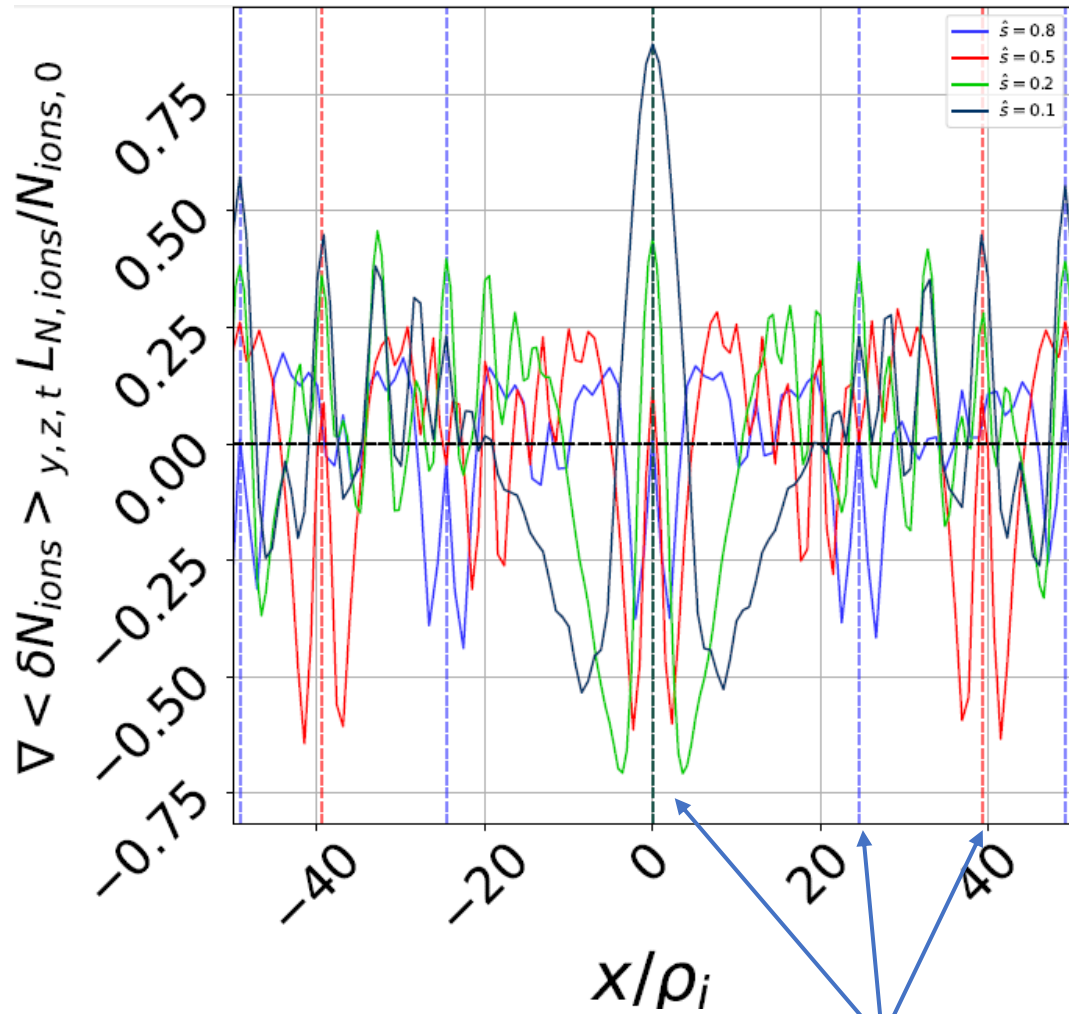
# Temperature gradients and auto-correlation



Rational surface location

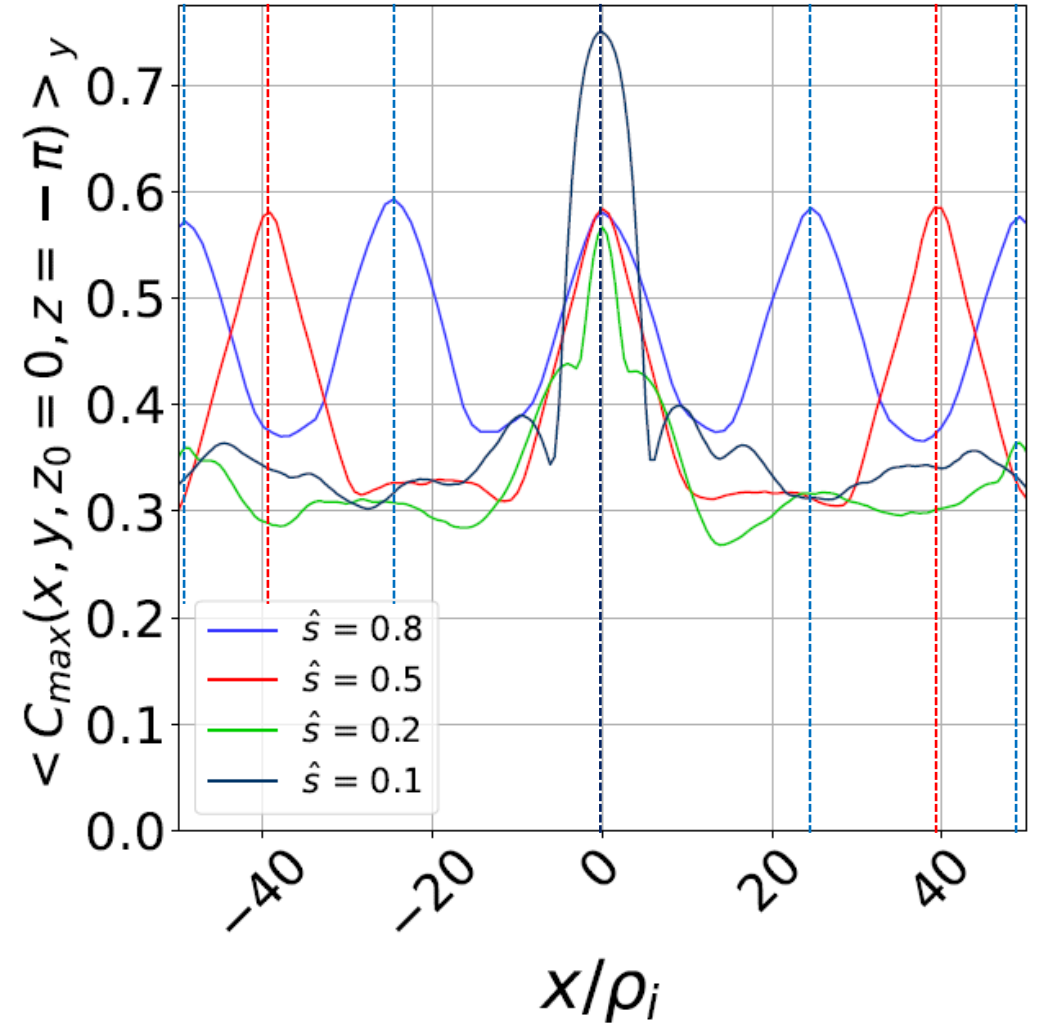
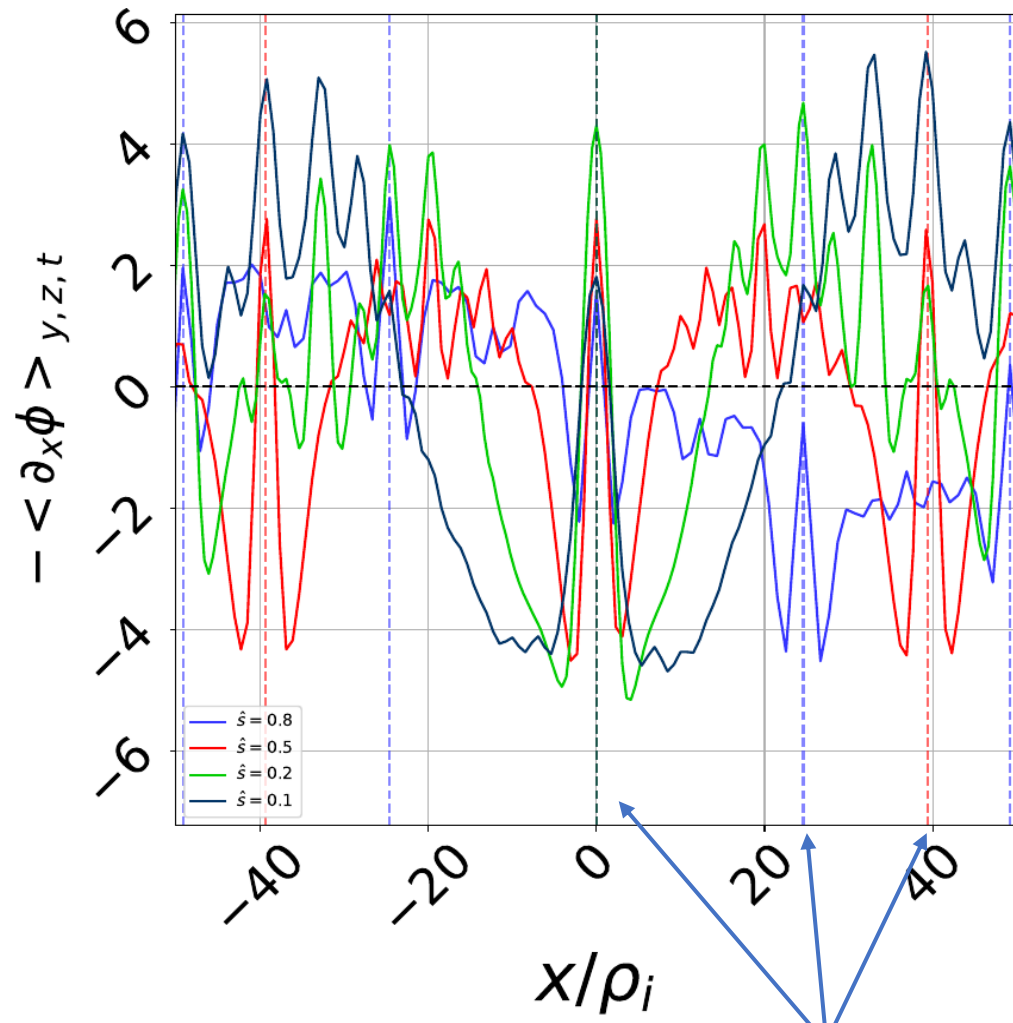


# Density gradients and auto-correlation

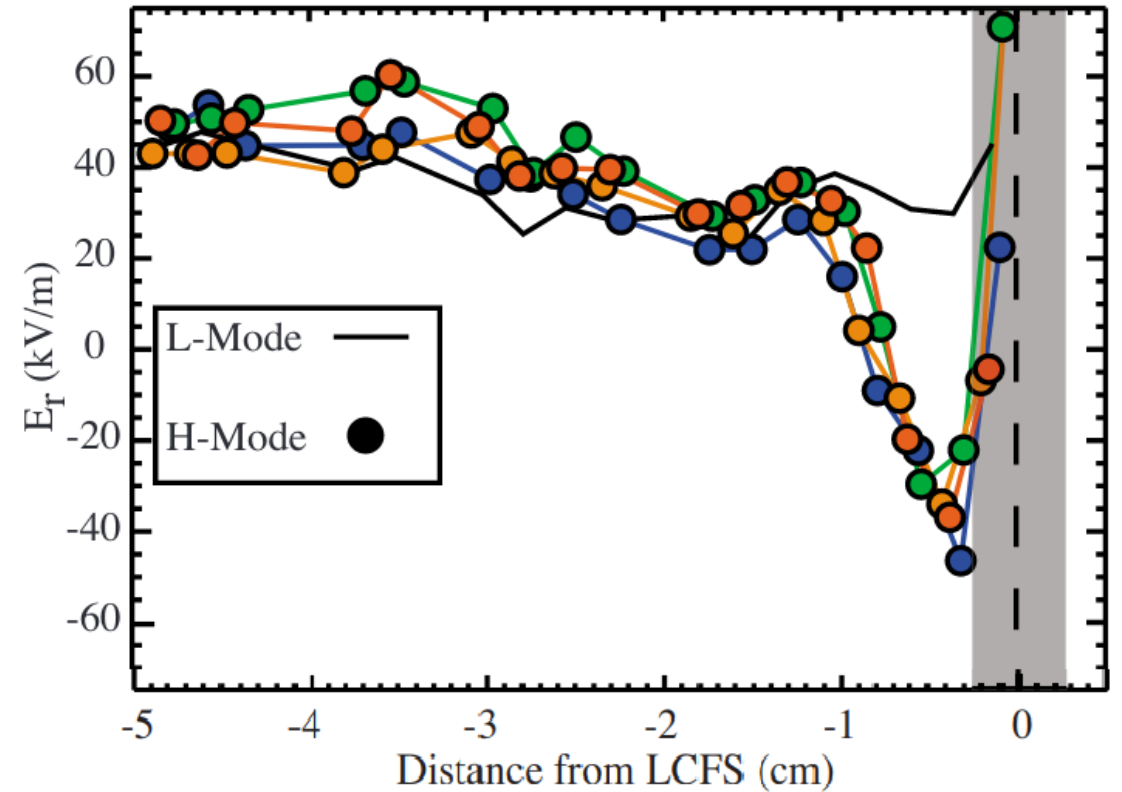
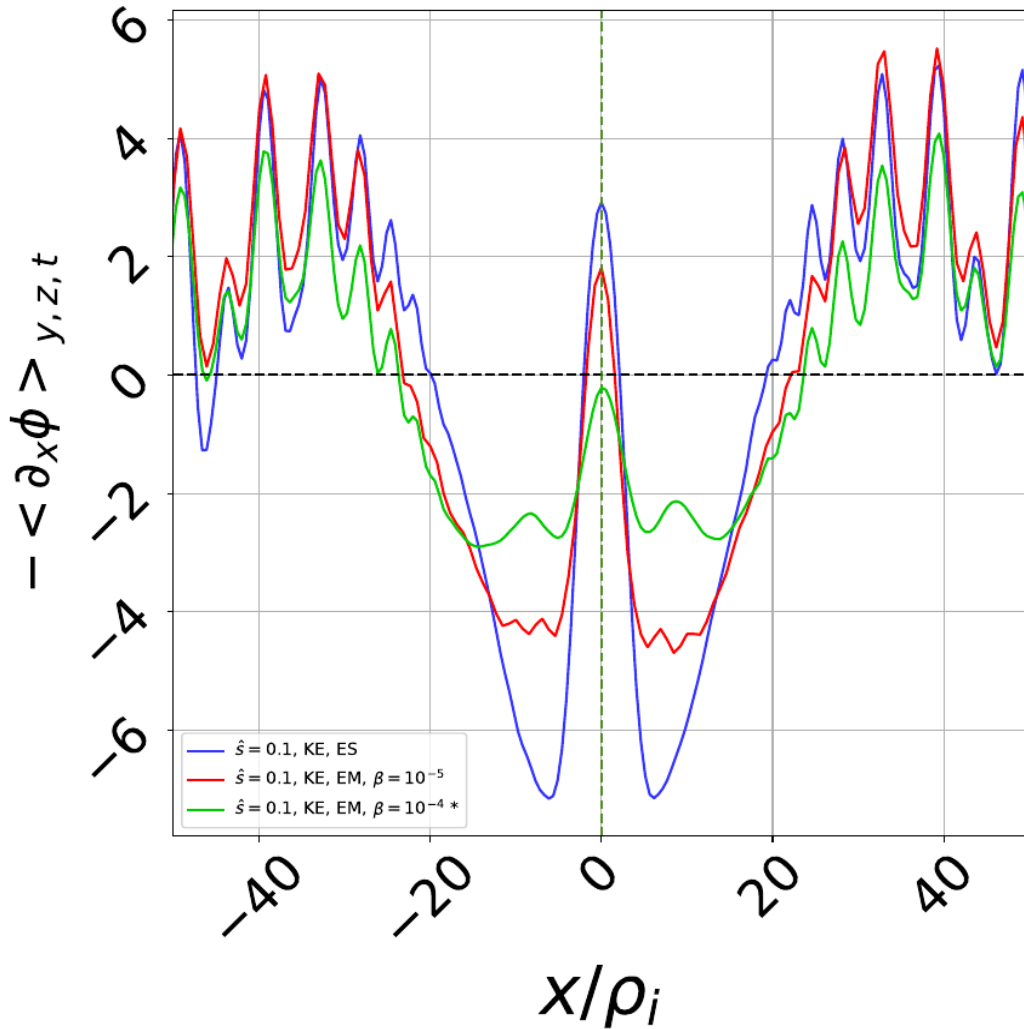


Rational surface location



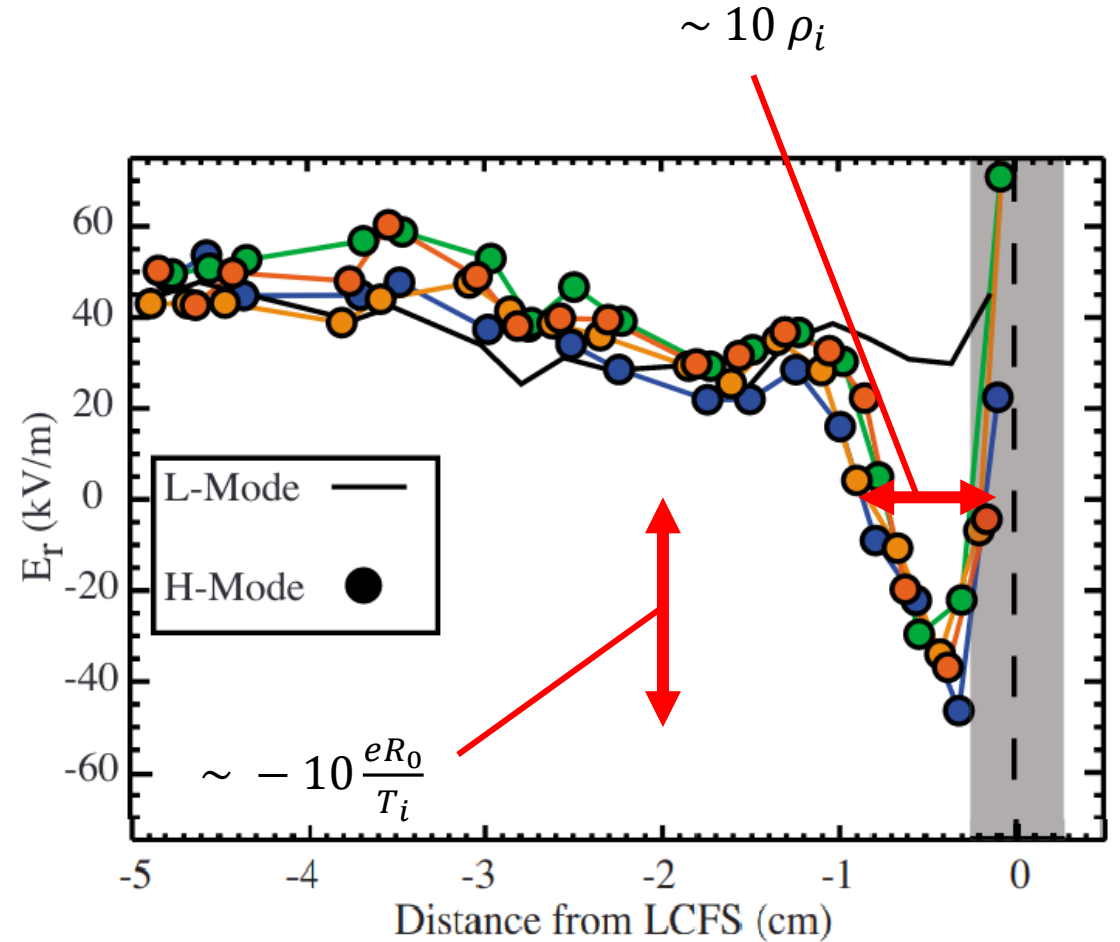
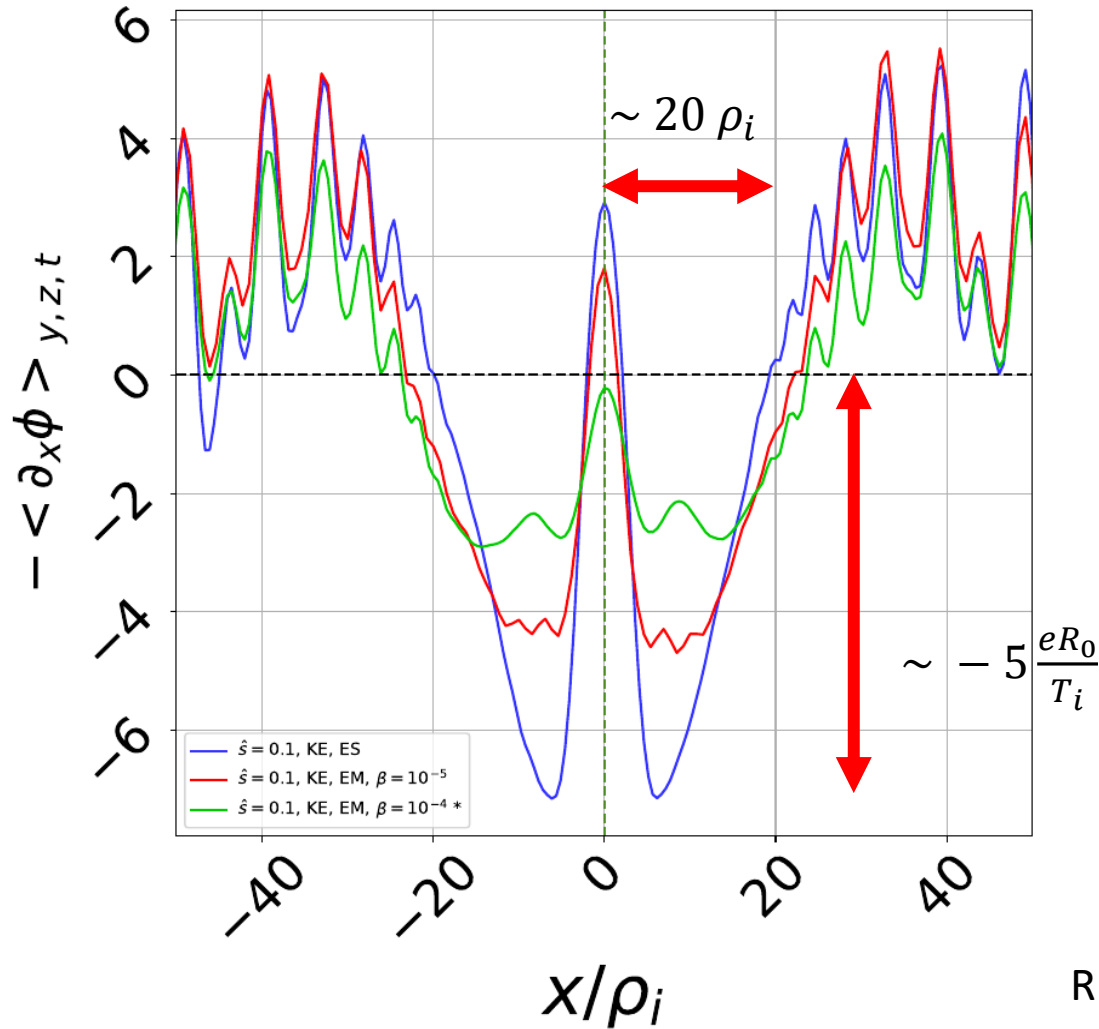


# Radial Electric field well at the edge



R. M. McDermott *et al.*, Phys. Plasmas 16, 056103 (2009)

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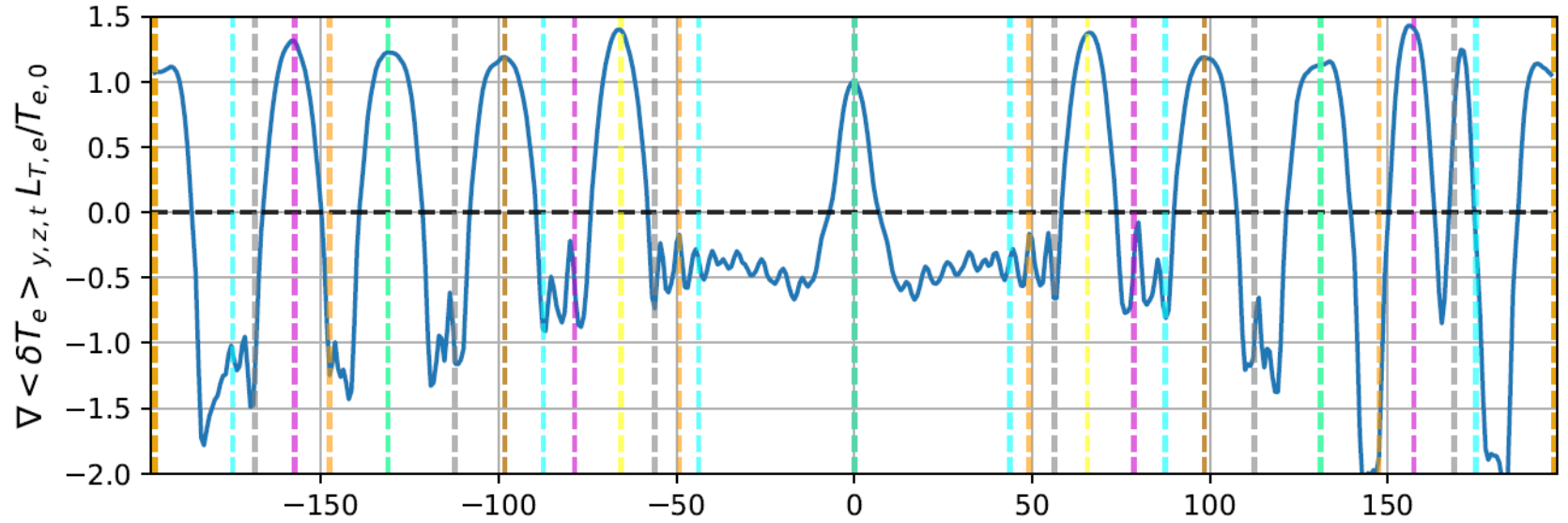
R. M. McDermott *et al.*, Phys. Plasmas 16, 056103 (2009)

# Effects of $N_{pol}$ on corrugations

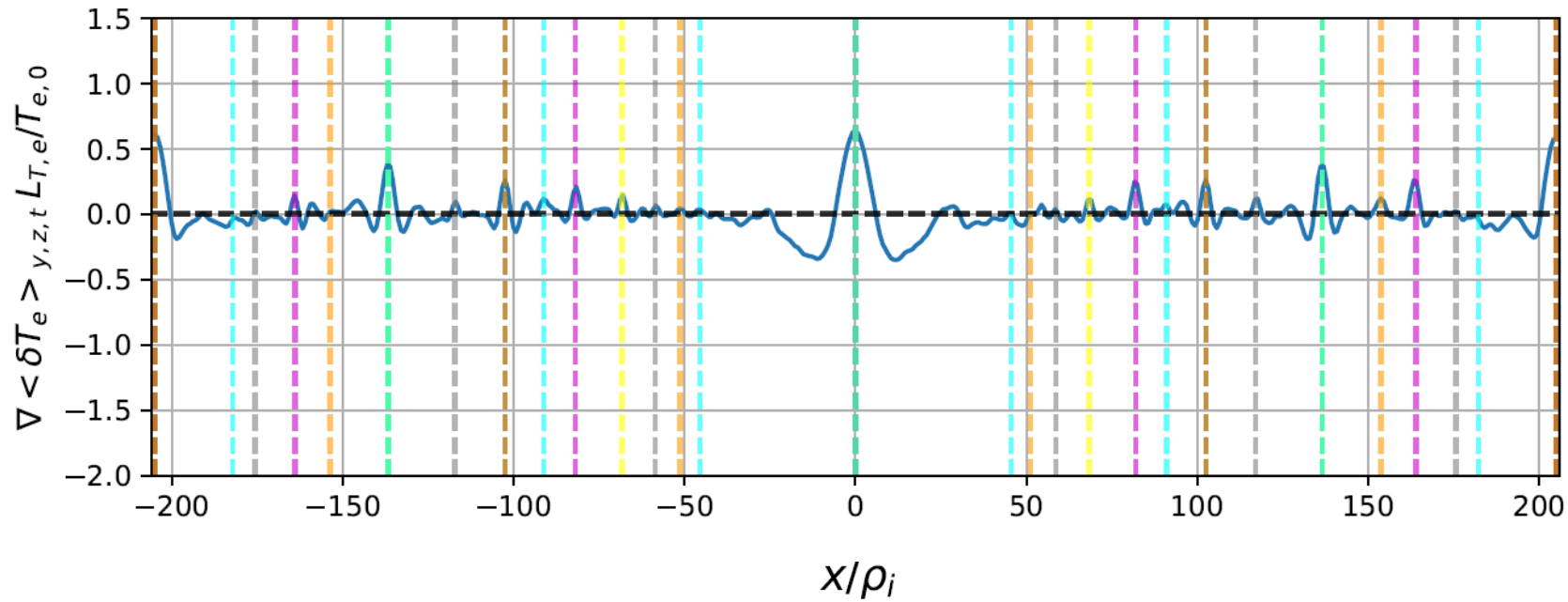
Number of poloidal turns  $N_{pol}$  has a large impact on profile corrugation due to reduction of self-interaction

$s=0.05$

$N_{pol} = 1$



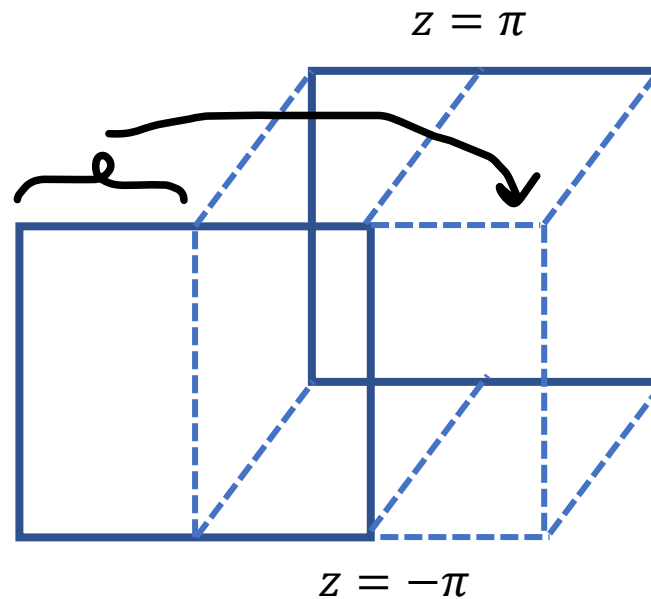
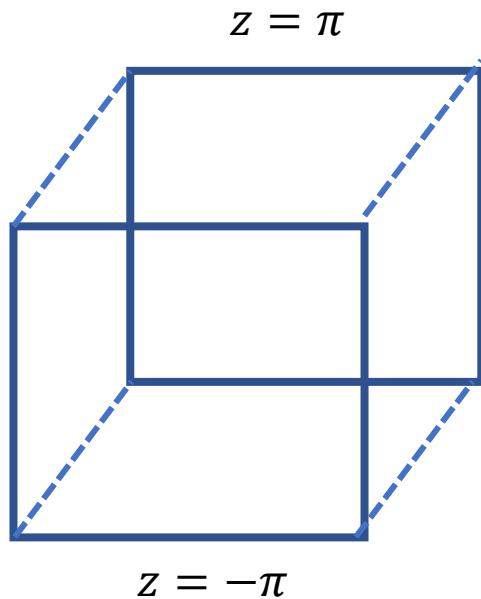
$N_{pol} = 3$

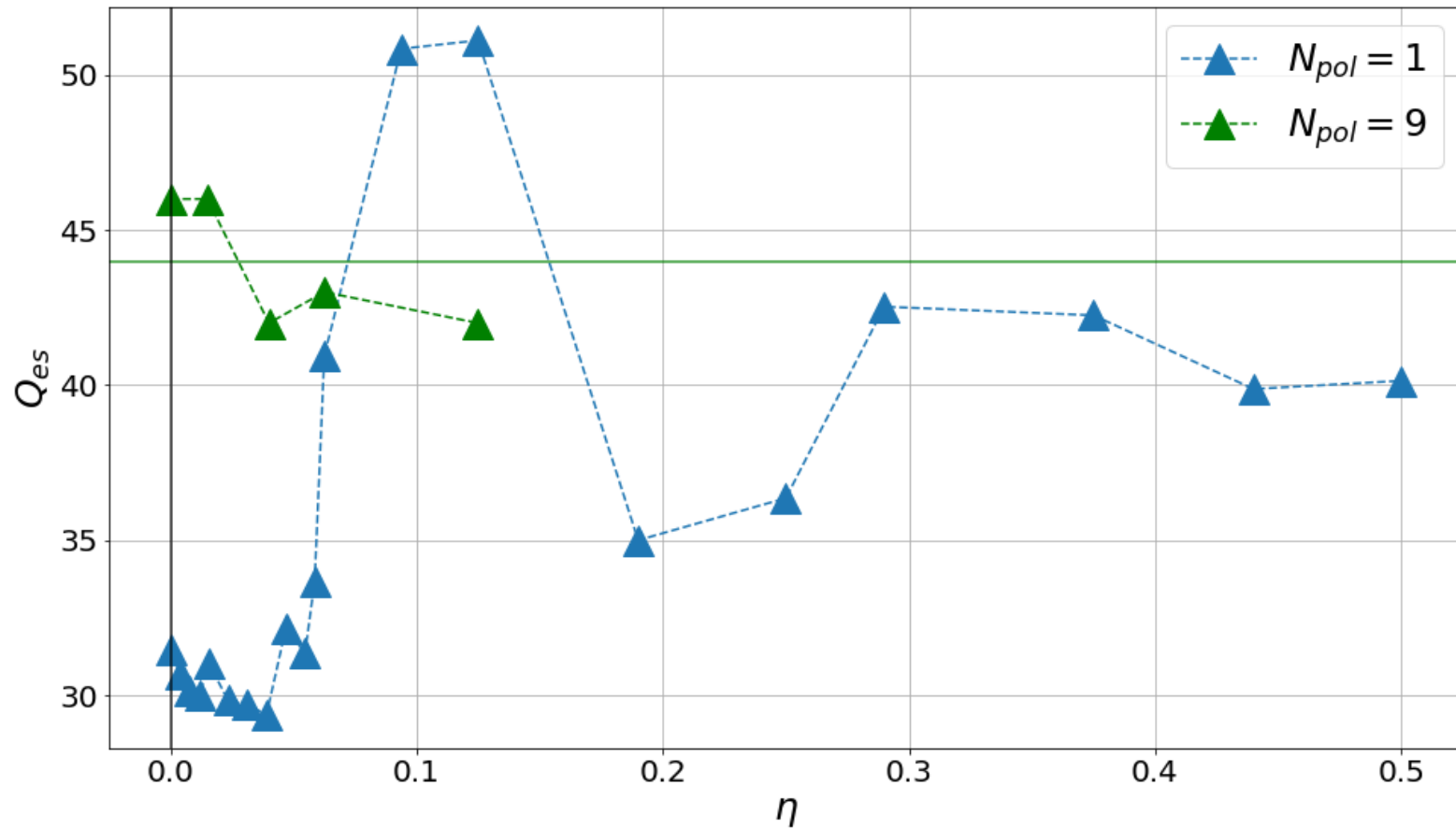


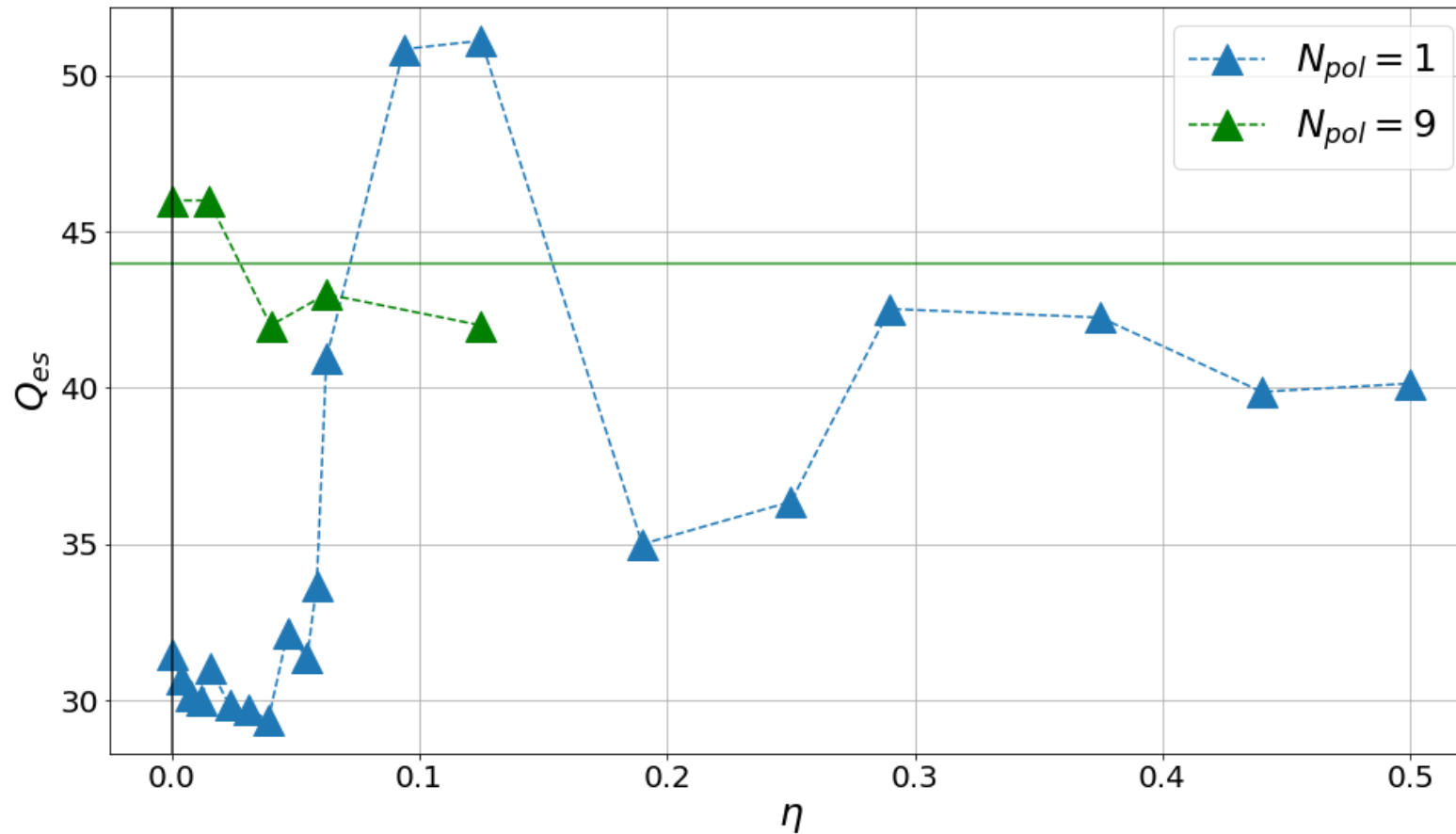
# $s=0$ boundary

In the  $s=0$  case

- Linear growth rates change slowly with  $q$
- Periodic parallel boundary with a shift







Preliminary simulations with **kinetic electrons** also show sharp change with varying  $\eta$  but require further study



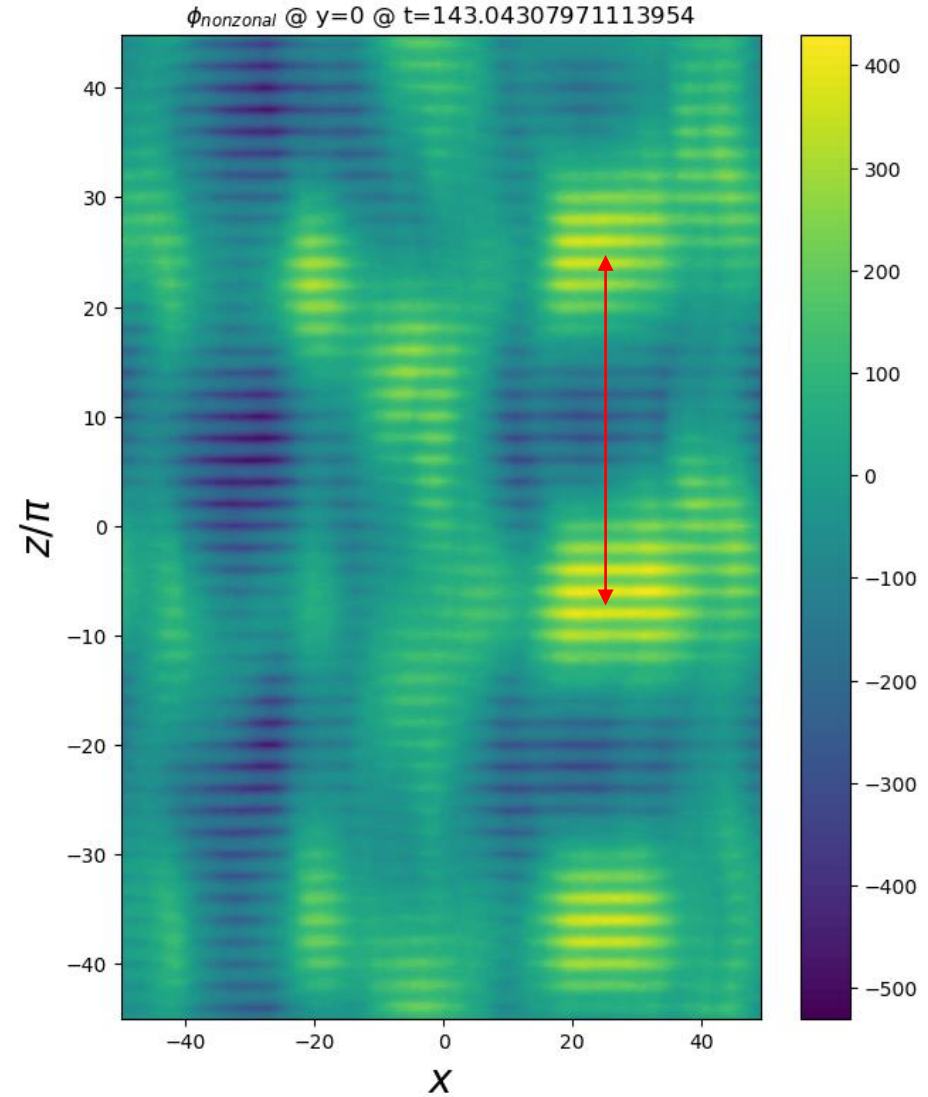
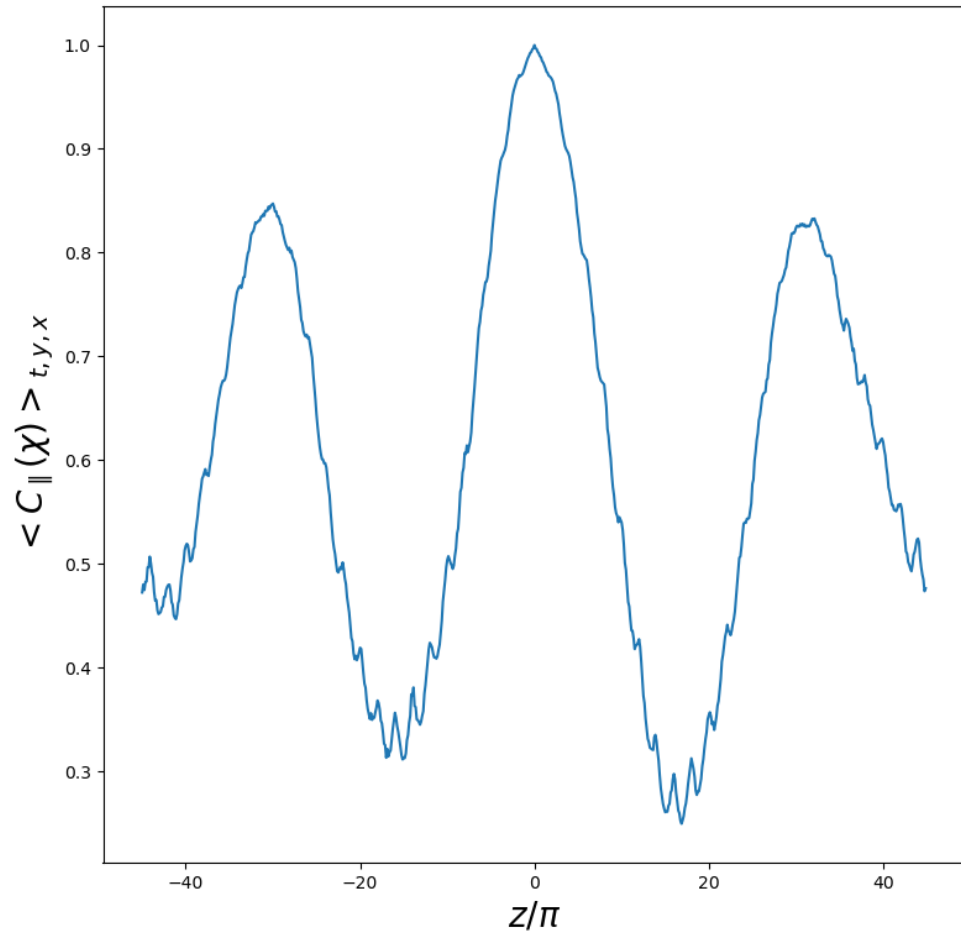
# Key nonlinear results

## Significant changes in the turbulent behaviour in simulations with low magnetic shear

- Strong stationary corrugations around low order rational surfaces that are comparable to the background profile gradients;
- We believe that this is a consequence of strong turbulent self-interaction in the parallel direction;
- Turbulent self-interaction seems to be stabilizing;
- Started comparisons between radial electric fields in the core with observed radial electric field wells at the plasma edge

# Currently outstanding questions

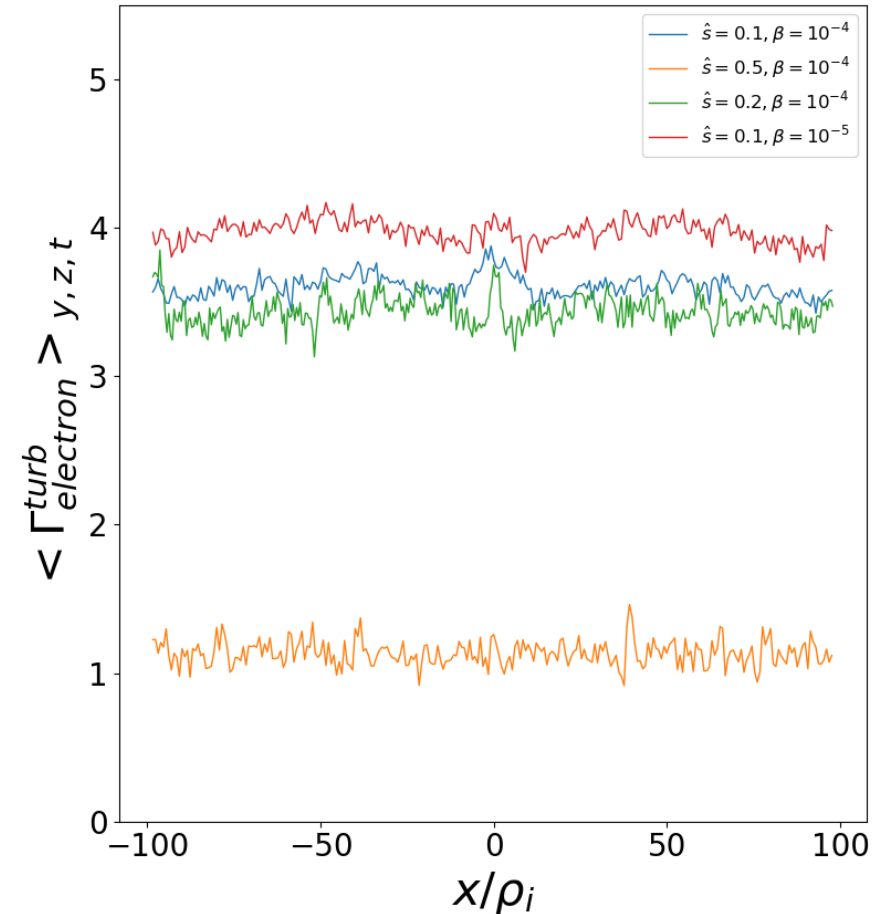
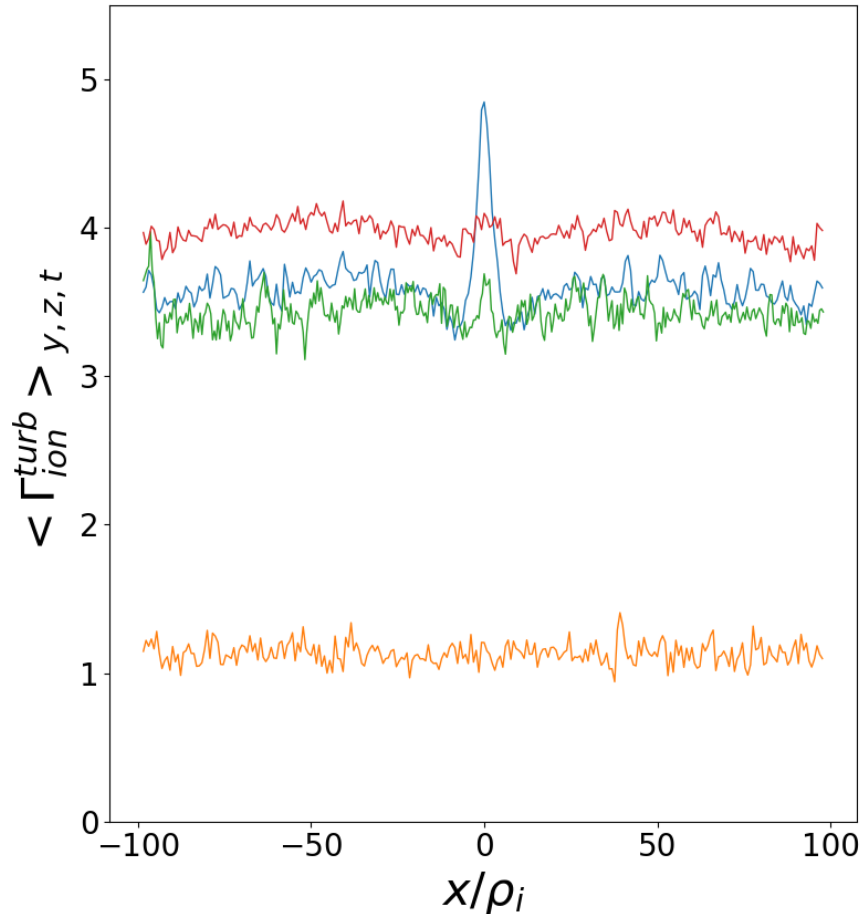
# s=0 long wavelength parallel wave



# Radial particle flux

$$\Gamma(x) = \langle n v_r \rangle_{FS}$$

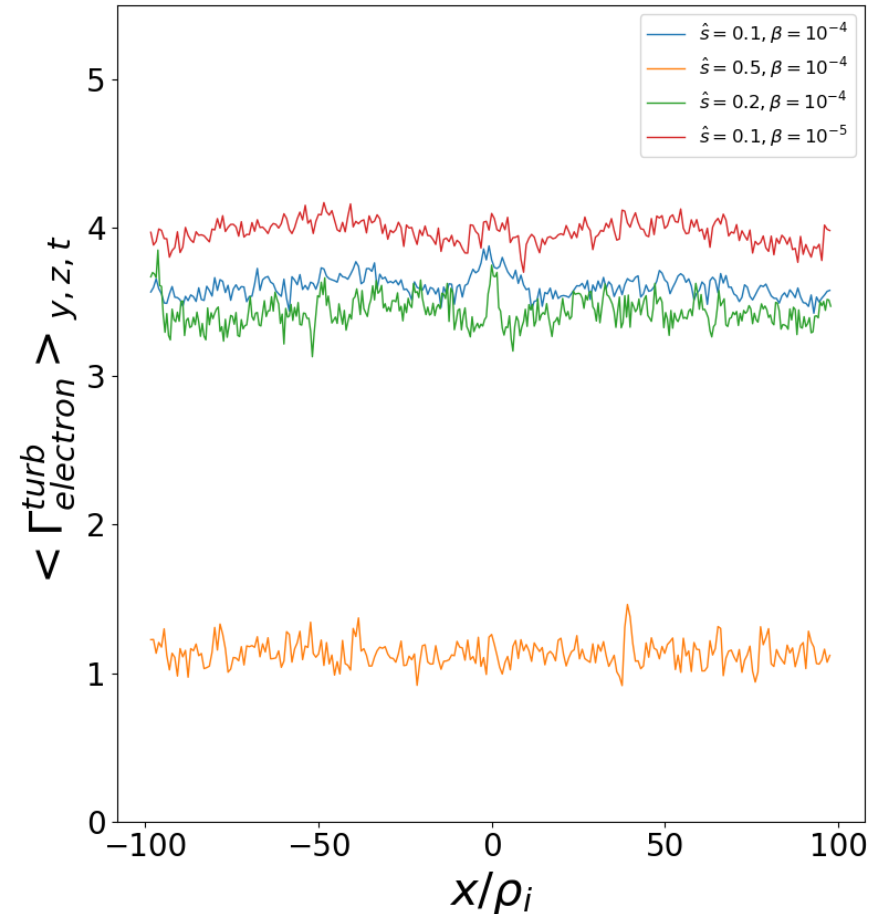
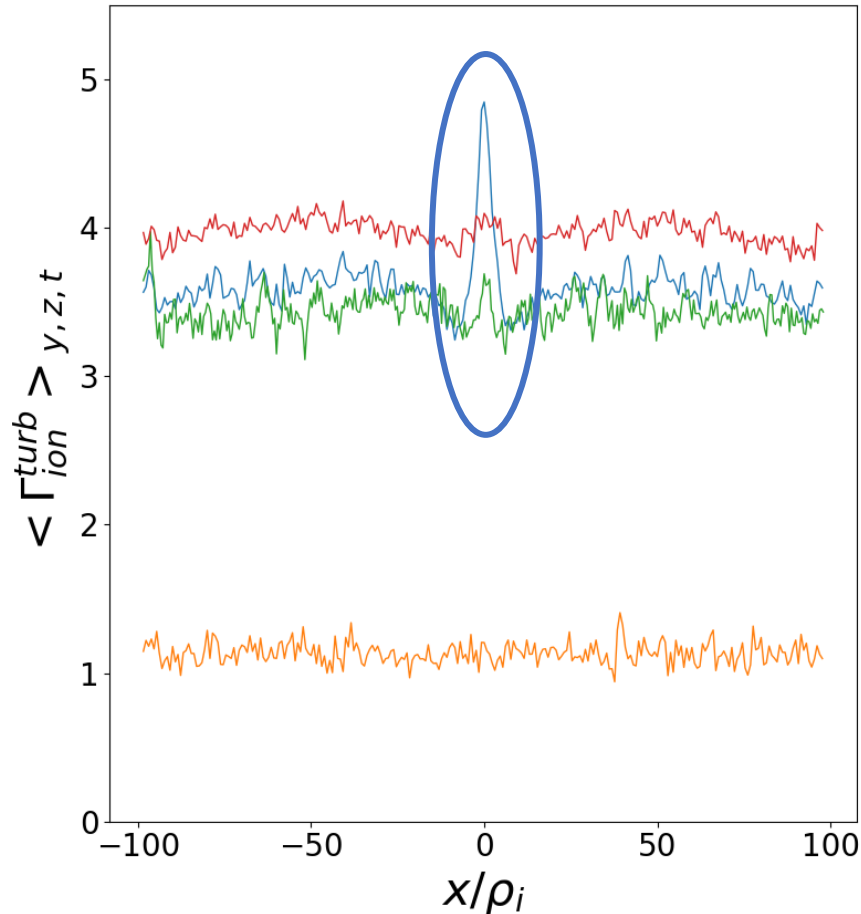
$$\Gamma(x) = - \left\langle \frac{c}{\tilde{C}} \frac{\partial \phi_1}{\partial y} M^{00}(\mathbf{x}) + \frac{1}{\tilde{C}} \frac{\partial A_{1,\parallel}}{\partial y} M^{01}(\mathbf{x}) \right\rangle_{FS}$$



# Radial particle flux

$$\Gamma(x) = \langle n v_r \rangle_{FS}$$

$$\Gamma(x) = - \left\langle \frac{c}{\tilde{C}} \frac{\partial \phi_1}{\partial y} M^{00}(\mathbf{x}) + \frac{1}{\tilde{C}} \frac{\partial A_{1,\parallel}}{\partial y} M^{01}(\mathbf{x}) \right\rangle_{FS}$$



References: Hauke Doerk, Gyrokinetic Simulation of Microtearing Turbulence, Universität Ulm, 2012

# EPFL Summary

- Transition between slab and toroidal ITG modes in linear study identified;
- M4.1: Found strong plasma profile (i.e.  $\phi_1, \nabla n_1, \nabla T_1$ ) corrugations around rational surfaces for  $\hat{s} \leq 0.1$ ;
- M4.1: Preliminary ITB simulations display a normalized  $E_r$  well that is comparable to experimental measurements in pedestal (pedestal  $E_r$  well is roughly 2 times narrower and 2 deeper);
- M4.1: Parallel correlation of turbulent eddies become much longer when  $\hat{s} \leq 0.1$  and very long wavelength modes appear for  $\hat{s} = 0$

Thank you for your attention