

# 4th E-TASC Scientific Board

## Monitoring of ENR-MOD 2021 activities

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Energetic particle optimization of stellarator devices using  
near-axis magnetic fields

*ENR-MOD.01.IST-T001*

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November 26, 2021

# Outline



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- ❑ Motivation
- ❑ Objectives
- ❑ Tasks
- ❑ 2022 Plan
- ❑ Preliminary Results
- ❑ Deliverables for 2022
- ❑ Possible Revisions for 2022

# Motivation

PAPER

## Modeling of energetic particle transport in optimized stellarators

A. Bader<sup>7,1</sup> , D.T. Anderson<sup>1</sup>, M. Drevlak<sup>2</sup>, B.J. Faber<sup>1</sup> , C.C. Hegna<sup>1</sup>, S. Henneberg<sup>2</sup> ,  
M. Landreman<sup>3</sup> , J.C. Schmitt<sup>4</sup> , Y. Suzuki<sup>5</sup>  and A. Ware<sup>6</sup>

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[Nuclear Fusion, Volume 61, Number 11](#)

Citation A. Bader et al 2021 *Nucl. Fusion* **61** 116060

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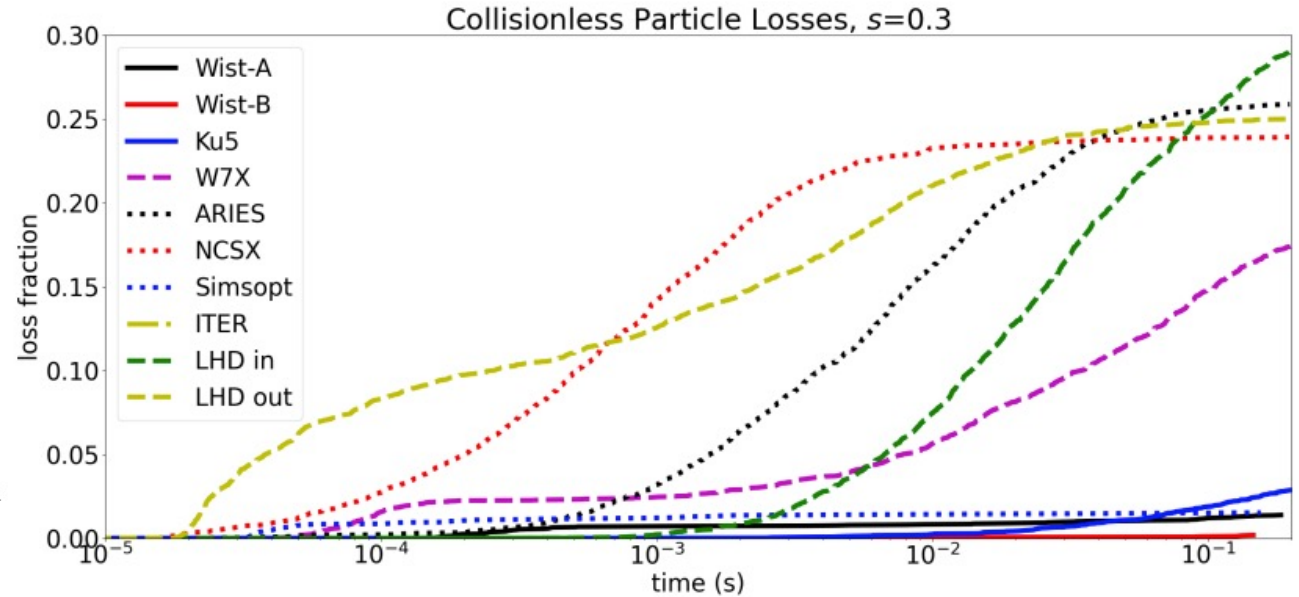
<sup>3</sup> University of Maryland, College Park, MD, United States of America

<sup>4</sup> Auburn University, Auburn, AL, United States of America

<sup>5</sup> Hiroshima University, Higashi-Hiroshima, Japan

<sup>6</sup> University of Montana, Missoula, MT, United states of America

Figure 4: Particle loss as a function of time for all configurations for particles born on the  $s=0.3$  surface.



Why some stellarators confine particles better than others?

How to increase the performance of optimized stellarators?

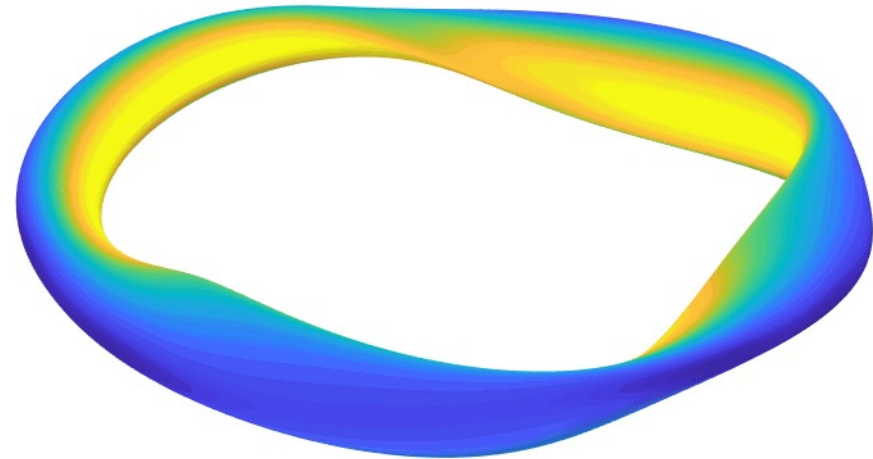
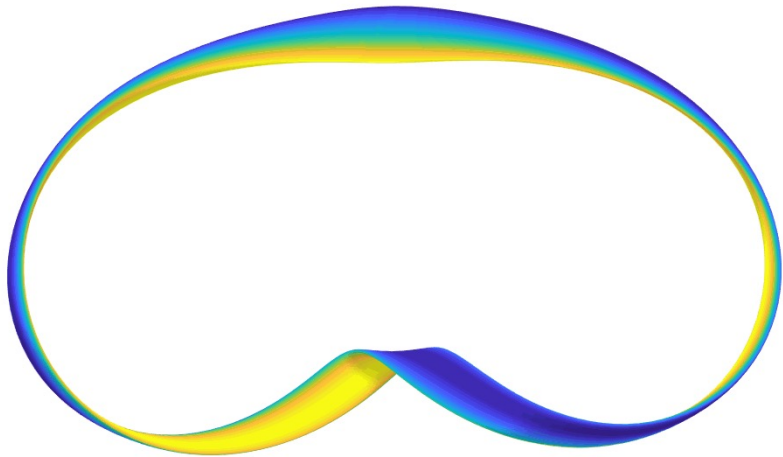
Can we survey the phase-space of possible stellarator shapes?

# Objectives

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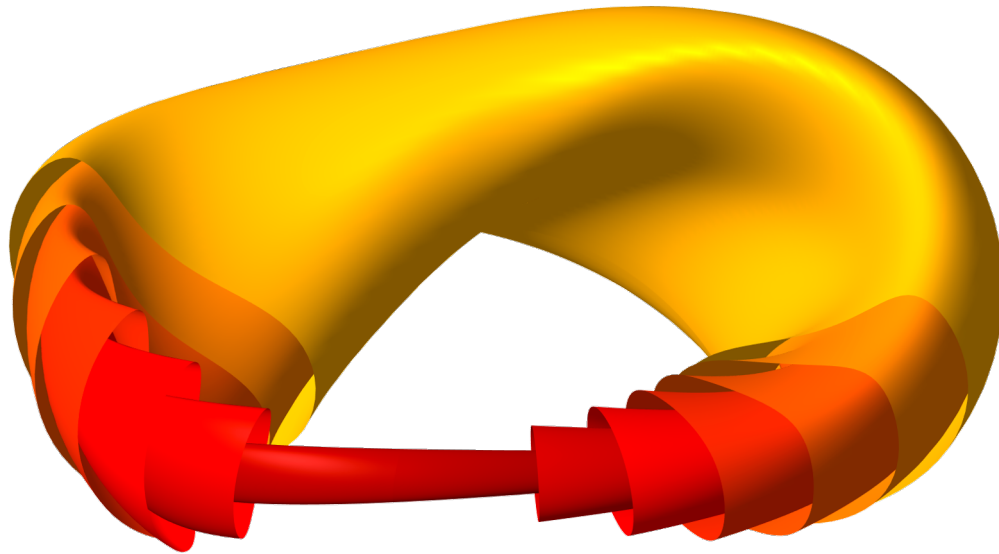
Find reactor relevant stellarator shapes in a reliable and efficient manner

- Use the near-axis expansion to get a stellarator shape ( $\sim 1$  ms)
- Obtain particle confinement using particle tracer codes ( $\sim 10$  s)
- Optimize in the space of near-axis configurations ( $\sim 10$  min)



# The near-axis expansion

Solve ideal MHD equations for small  $\epsilon = a/R$

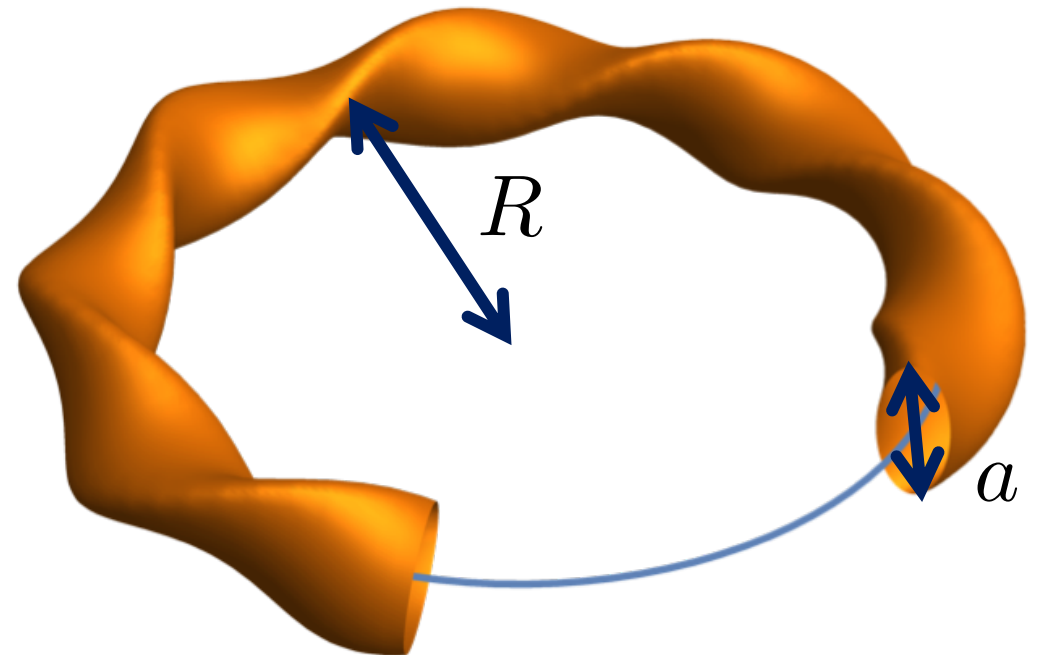


Direct approach

[R. Jorge et al, JPP 86 (2020)]

Inverse approach

[M. Landreman et al, JPP 86 (2018)]

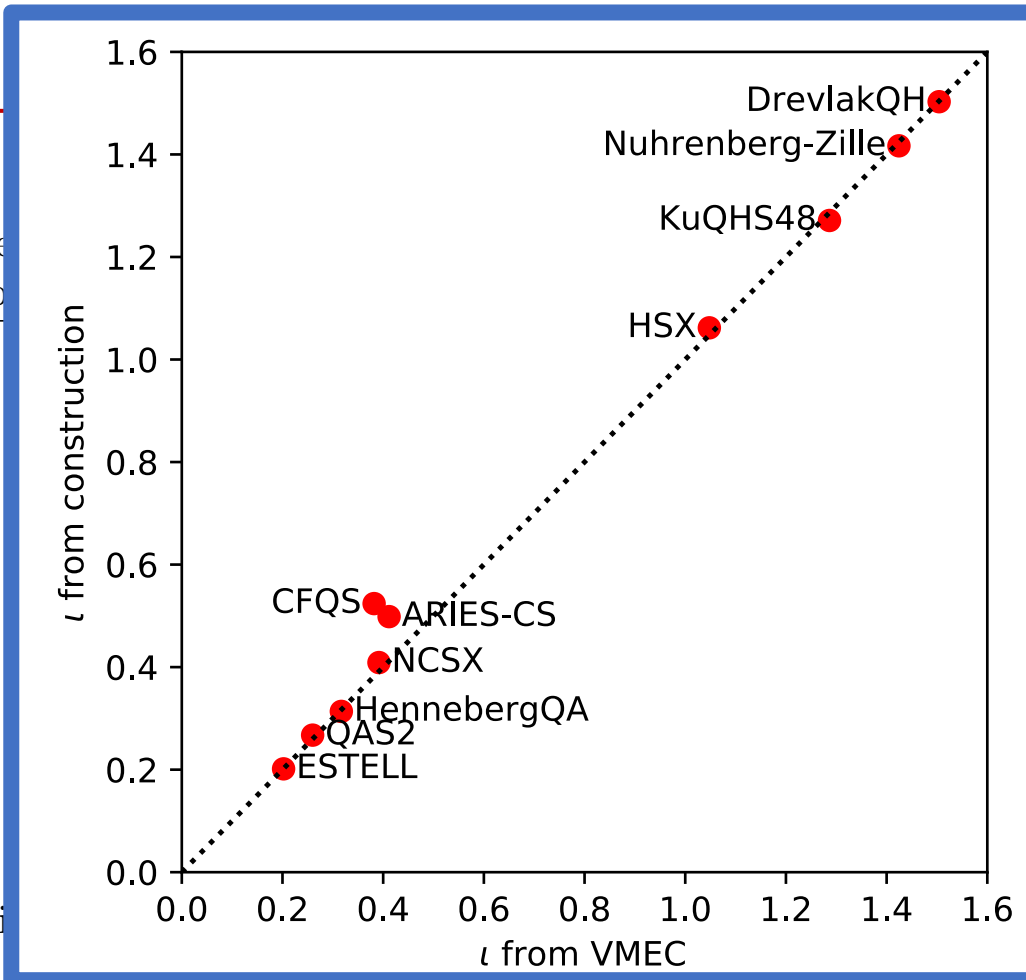


# Every quasisymmetric design available

- Construct a near-axis representation for each design
- Compare VM

Configuration	Aspect ratio
NZ1988	12
HSX	10
KuQHS48	8.1
WISTELL-A	6.7
Drevlak	8.6
NCSX	4.4
ARIES-CS	4.5
QAS2	2.6
ESTELL	5.3
CFQS	4.3
Henneberg	3.4

Table 1: Quasi



Best-fit	$B_0$ [T]
.157	0.205
.28	1.00
.147	1.20
.791	2.54
.0899	3.97
.408	1.55
.0740	5.69
.347	1.79
.570	1.00
.586	0.933
.302	2.41

Jorge & Landreman, *Plasma Phys. Control. Fusion* **63** (2021)

## **WPI – Particle Tracer Code for Arbitrary Geometries**

Solve guiding-center equations in several geometries (analytical/numerical):

- Near-axis direct approach (Mercier)
- Near-axis indirect approach (Garren-Boozer)
- VMEC
- Helena (benchmarking)

## **WP2 – Measure and Optimize Confinement**

- Numerical and analytical metrics for fast-particle confinement
- Integration with SIMSOPT

## **WP3 – Optimized Stellarator Equilibria**

Find optimized designs (not necessarily quasisymmetric or quasi-isodynamic) using a SIMSOPT based optimization

## **WP4 – Physics Study of Nemov's $\Gamma_c$ criterion**

Study the viability of using the parameter  $\Gamma_c$  as a cost function in fast-particle confinement stellarator studies

## **WP5 – Orbits in Free-Boundary Magnetic Fields**

Using SPEC or VMEC, study particle orbits in coil-based optimized stellarators



# 2022 plan – Particle Tracing

**::gyronimo:: - gyromotion for the people, by the people -**

*An object-oriented library for gyromotion applications in plasma physics.*

## Philosophy and purpose:

Have you ever had a bright and promising idea about gyromotion in plasmas that just faded away the moment you realised the amount of non-trivial, tedious, unrewarding, non-physics details you would have to implement before you could get a simple glimpse over the results?

P. Rodrigues, [github.com/prodrigs/gyronimo](https://github.com/prodrigs/gyronimo)

Solve guiding-center equations of motion in a general coordinate system

$$\left[1 + \frac{\tilde{v}_{\parallel}}{\tilde{\Omega}} (\mathbf{b} \cdot \tilde{\nabla} \times \mathbf{b})\right] \frac{d\mathbf{X}}{d\tau} = \tilde{v}_{\parallel} \mathbf{b} + \frac{1}{\tilde{\Omega}} \left[ \tilde{v}_{\parallel}^2 \tilde{\nabla} \times \mathbf{b} + \mathbf{b} \times (\tilde{v}_{\parallel} \partial_{\tau} \mathbf{b} + \frac{1}{2} \tilde{\mu} \tilde{\nabla} \tilde{B} - \tilde{\mathbf{E}}) \right]$$
$$\left[1 + \frac{\tilde{v}_{\parallel}}{\tilde{\Omega}} (\mathbf{b} \cdot \tilde{\nabla} \times \mathbf{b})\right] \frac{d\tilde{v}_{\parallel}}{d\tau} = - \left( \mathbf{b} + \frac{\tilde{v}_{\parallel}}{\tilde{\Omega}} \tilde{\nabla} \times \mathbf{b} \right) \cdot \left( \frac{1}{2} \tilde{\mu} \tilde{\nabla} \tilde{B} + \tilde{v}_{\parallel} \partial_{\tau} \mathbf{b} - \tilde{\mathbf{E}} \right).$$

Test near axis direct approach, indirect, VMEC, SPEC, Bio—Savart, etc...

## **WPI**

C++ Library with documentation on GitHub

Python interface with near-axis expansion coordinates

C++ interface with VMEC equilibria

Benchmark with other particle tracer codes such as BEAMS3D

## **WP2**

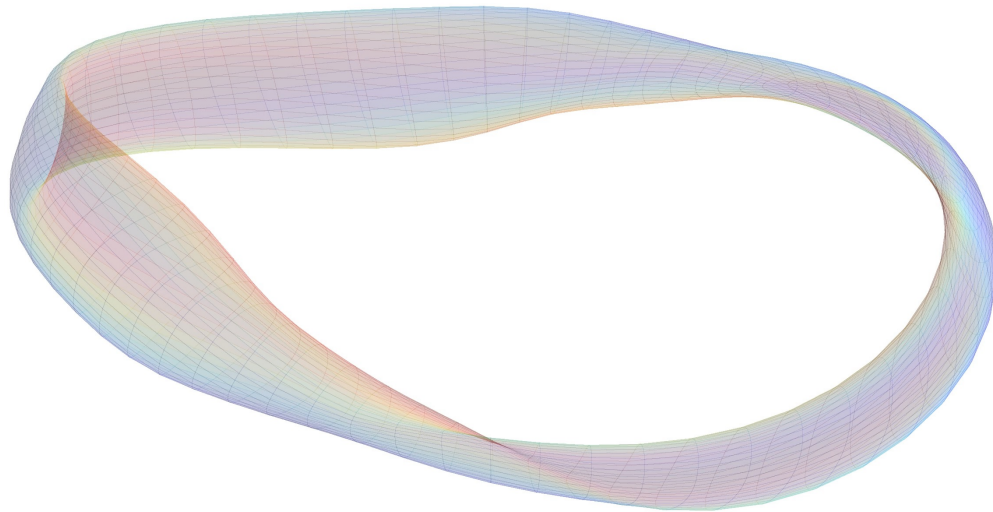
Analytical/numerical objective function to measure fast particle confinement

Python interface with SIMSOPT/STELLOPT

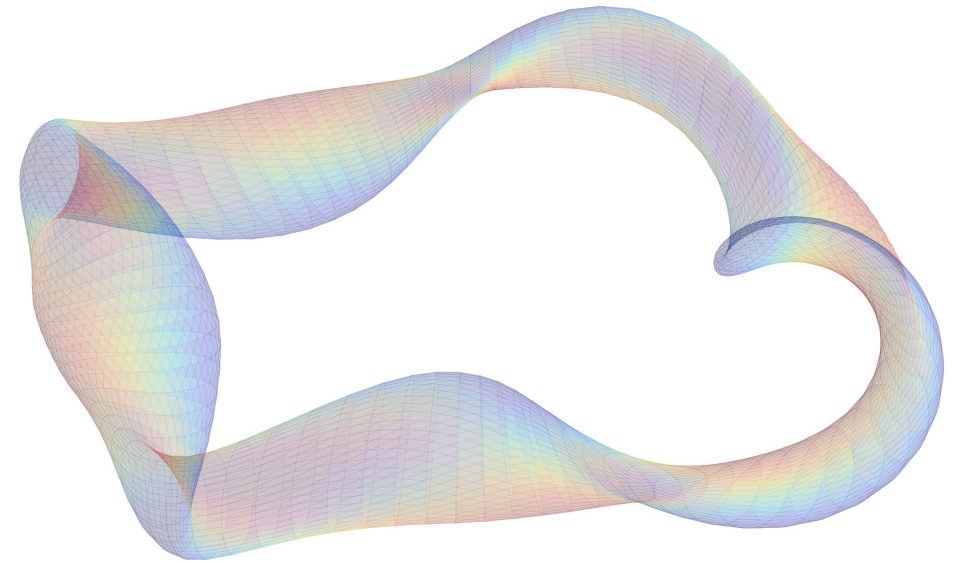
# Preliminary Results

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## Second order near-axis geometry



Quasi-axisymmetric design  
of [1], section 5.2



Quasi-helically symmetric  
design of [1], section 5.4

[1] Landreman & Sengupta (2019), 85(6), JPP

# Preliminary Results

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## WPI

- C++ Library with documentation on GitHub ✓
- Python interface with near-axis expansion coordinates ✓
- C++ interface with VMEC equilibria
- Benchmark with other particle tracer codes such as BEAMS3D ✓

## WP2

- Analytical/numerical objective function to measure fast particle confinement
- Python interface with SIMSOPT/STELLOPT ✓

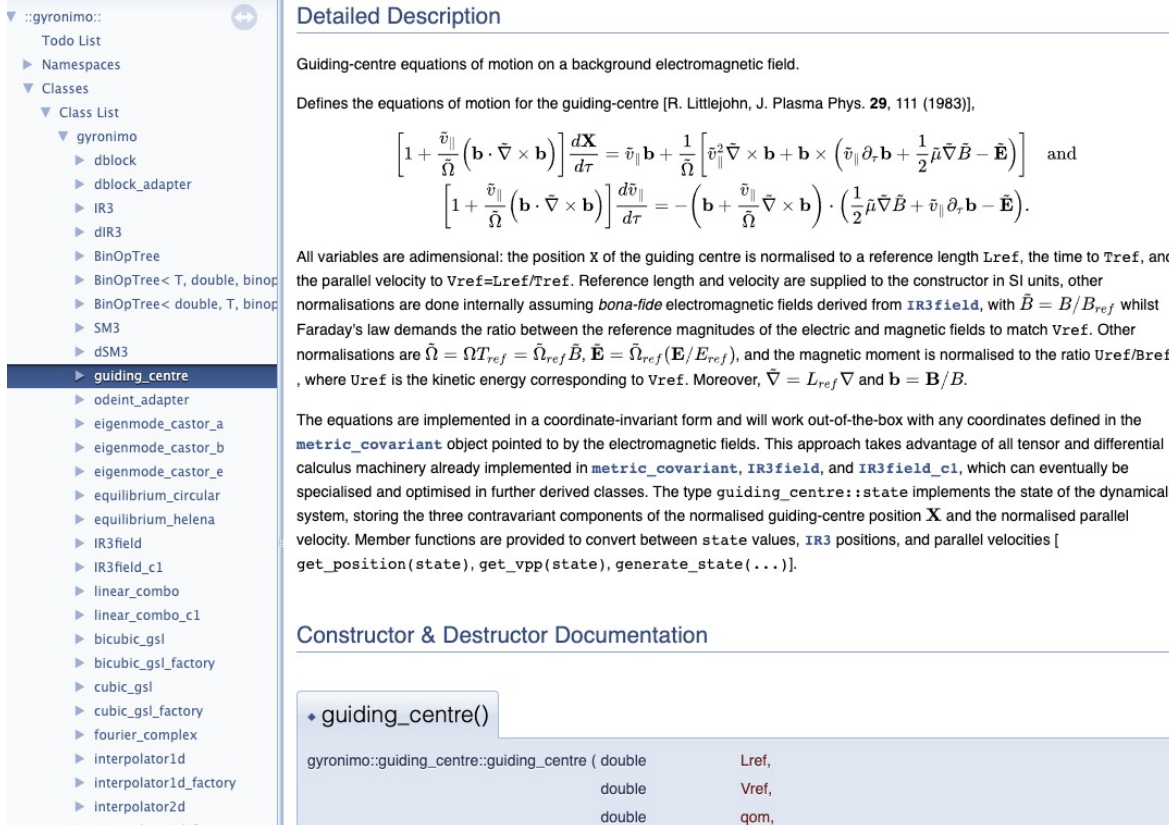
# Preliminary Results

## WPI

### C++ Library with documentation on GitHub

**::gyronimo::** 0.0

An object-oriented library for gyromotion applications in plasma physics.



**Detailed Description**

Guiding-centre equations of motion on a background electromagnetic field.

Defines the equations of motion for the guiding-centre [R. Littlejohn, J. Plasma Phys. **29**, 111 (1983)],

$$\left[1 + \frac{\tilde{v}_{\parallel}}{\tilde{\Omega}} (\mathbf{b} \cdot \tilde{\nabla} \times \mathbf{b})\right] \frac{d\mathbf{X}}{d\tau} = \tilde{v}_{\parallel} \mathbf{b} + \frac{1}{\tilde{\Omega}} \left[ \tilde{v}_{\parallel}^2 \tilde{\nabla} \times \mathbf{b} + \mathbf{b} \times (\tilde{v}_{\parallel} \partial_{\tau} \mathbf{b} + \frac{1}{2} \tilde{\mu} \tilde{\nabla} \tilde{B} - \tilde{\mathbf{E}}) \right] \quad \text{and}$$
$$\left[1 + \frac{\tilde{v}_{\parallel}}{\tilde{\Omega}} (\mathbf{b} \cdot \tilde{\nabla} \times \mathbf{b})\right] \frac{d\tilde{v}_{\parallel}}{d\tau} = - \left( \mathbf{b} + \frac{\tilde{v}_{\parallel}}{\tilde{\Omega}} \tilde{\nabla} \times \mathbf{b} \right) \cdot \left( \frac{1}{2} \tilde{\mu} \tilde{\nabla} \tilde{B} + \tilde{v}_{\parallel} \partial_{\tau} \mathbf{b} - \tilde{\mathbf{E}} \right).$$

All variables are adimensional: the position  $\mathbf{X}$  of the guiding centre is normalised to a reference length  $L_{ref}$ , the time to  $T_{ref}$ , and the parallel velocity to  $V_{ref}=L_{ref}/T_{ref}$ . Reference length and velocity are supplied to the constructor in SI units, other normalisations are done internally assuming *bona-fide* electromagnetic fields derived from `IR3field`, with  $\tilde{B} = B/B_{ref}$  whilst Faraday's law demands the ratio between the reference magnitudes of the electric and magnetic fields to match  $V_{ref}$ . Other normalisations are  $\tilde{\Omega} = \Omega T_{ref} = \tilde{\Omega}_{ref} \tilde{B}$ ,  $\tilde{\mathbf{E}} = \tilde{\Omega}_{ref} (\mathbf{E}/E_{ref})$ , and the magnetic moment is normalised to the ratio  $U_{ref}/B_{ref}$ , where  $U_{ref}$  is the kinetic energy corresponding to  $V_{ref}$ . Moreover,  $\tilde{\nabla} = L_{ref} \nabla$  and  $\mathbf{b} = \mathbf{B}/B$ .

The equations are implemented in a coordinate-invariant form and will work out-of-the-box with any coordinates defined in the `metric_covariant` object pointed to by the electromagnetic fields. This approach takes advantage of all tensor and differential calculus machinery already implemented in `metric_covariant`, `IR3field`, and `IR3field_c1`, which can eventually be specialised and optimised in further derived classes. The type `guiding_centre::state` implements the state of the dynamical system, storing the three contravariant components of the normalised guiding-centre position  $\mathbf{X}$  and the normalised parallel velocity. Member functions are provided to convert between state values, `IR3` positions, and parallel velocities [`get_position(state)`, `get_vpp(state)`, `generate_state(...)`].

**Constructor & Destructor Documentation**

◆ `guiding_centre()`

```
gyronimo::guiding_centre::guiding_centre ( double Lref,
                                           double Vref,
                                           double qom,
```

Work by Paulo Rodrigues, IST

# Preliminary Results

## WPI

C++ Library with documentation on GitHub ✓

Python interface with near-axis expansion coordinates ✓

The screenshot displays the official documentation for pybind11. The header includes the project name 'pybind11' and the version 'stable'. A search bar is present. The navigation menu on the left is organized into sections: 'THE BASICS' (including installation and first steps) and 'ADVANCED TOPICS' (including functions, classes, and exceptions). The main content area features the 'pybind11' logo, a tagline, and several status badges indicating that documentation, CI, and build processes are passing. It also provides links to various examples and lists supported Python versions (2.7 to 3.10).

**PYBIND11:** Seamless creation of Python bindings to C++ code

- Straight interfacing of gyronimo C++ libraries in Python
- Easy to connect with other stellarator codes
- Orchestrating multiple particle tracing
- Post-processing and plotting
- Lightweight

### In other words...

Python – Slower, easy to read, write and plot data

C++ – Faster, core of the code, solves the equations

# Preliminary Results

## WPI

C++ Library with documentation on GitHub ✓

Python interface with near-axis expansion coordinates ✓

### C++

```
#include <../include/pybind11/pybind11.h>
#include <../include/pybind11/stl.h>
#include <gyronimo/interpolators/cubic_gsl.hh>
#include <gyronimo/metrics/metric_stellna.hh>
#include <gyronimo/core/dblock.hh>

#include <gyronimo/dynamics/odeint_adapter.hh>
#include <gyronimo/core/codata.hh>
namespace py = pybind11;
using namespace gyronimo;

std::vector< std::vector<double>> gc_solver(
    int field_periods,

// Integrate for t in [0,Tfinal], with dt=Tfinal/nsamples, using RK4.
boost::numeric::odeint::runge_kutta4<gyronimo::guiding_centre::state> integration_algorithm;
boost::numeric::odeint::integrate_const(
    integration_algorithm, odeint_adapter(&gc),
    initial_state, 0.0, Tfinal, Tfinal/nsamples, push_back_state_and_time(x_vec,&qsc,&gc) );

return x_vec;
}

// Python wrapper functions
PYBIND11_MODULE(NEAT, m) {
    m.doc() = "Gyronimo Wrapper for the Stellarator Near-Axis Expansion (STELLNA)";
    m.def("gc_solver",&gc_solver);
}
```

### Python

```
# Quasi-helically symmetric stellarator
r0=0.15
Lambda = [0.86]
Tfinal = 400
stel = Qsc.from_paper(4, nphi=nphi)
theta0 = [1.15]
phi0 = [0.0]
```

```
# Call Gyronimo
sol = np.array(NEAT.gc_solver(int(stel.nfp),
```

```
import mayavi.mlab as mlab
fig = mlab.figure(bgcolor=(1,1,1), size=(430,720))
[mlab.plot3d(rpos_cartesian[i][0], rpos_cartesian[i][1], rpos_cartesian[i][2],
```

# Preliminary Results

## WPI

C++ Library with documentation on GitHub ✓

Python interface with near-axis expansion coordinates ✓

## Python

Able to model QA (NCSX), QH (HSX) and QI (W7-X) stellarators

- QA and QH stellarators conserve canonical angular momentum

$$p_\chi = \left( \frac{\partial L}{\partial \dot{\zeta}} \right)_{\chi, \dot{\chi}, \zeta} = q \frac{N}{M} \psi - q \psi_p + \frac{m v_{\parallel}}{B} \left( G + \frac{N}{M} I \right)$$

orbit width  $\propto 1/|\iota - N|$

- QI stellarators do not conserve canonical angular momentum, but estimates are possible

$$\Delta\psi = -\frac{\partial}{\partial \alpha} \int_{B(l_0)}^{B(l)} h \frac{\partial}{\partial B} \left( \frac{v_{\parallel}}{\Omega} \right) dB$$

```
# Quasi-helically symmetric stellarator
r0=0.15
Lambda = [0.86]
Tfinal = 400
stel = Qsc.from_paper(4, nphi=nphi)
theta0 = [1.15]
phi0 = [0.0]
```

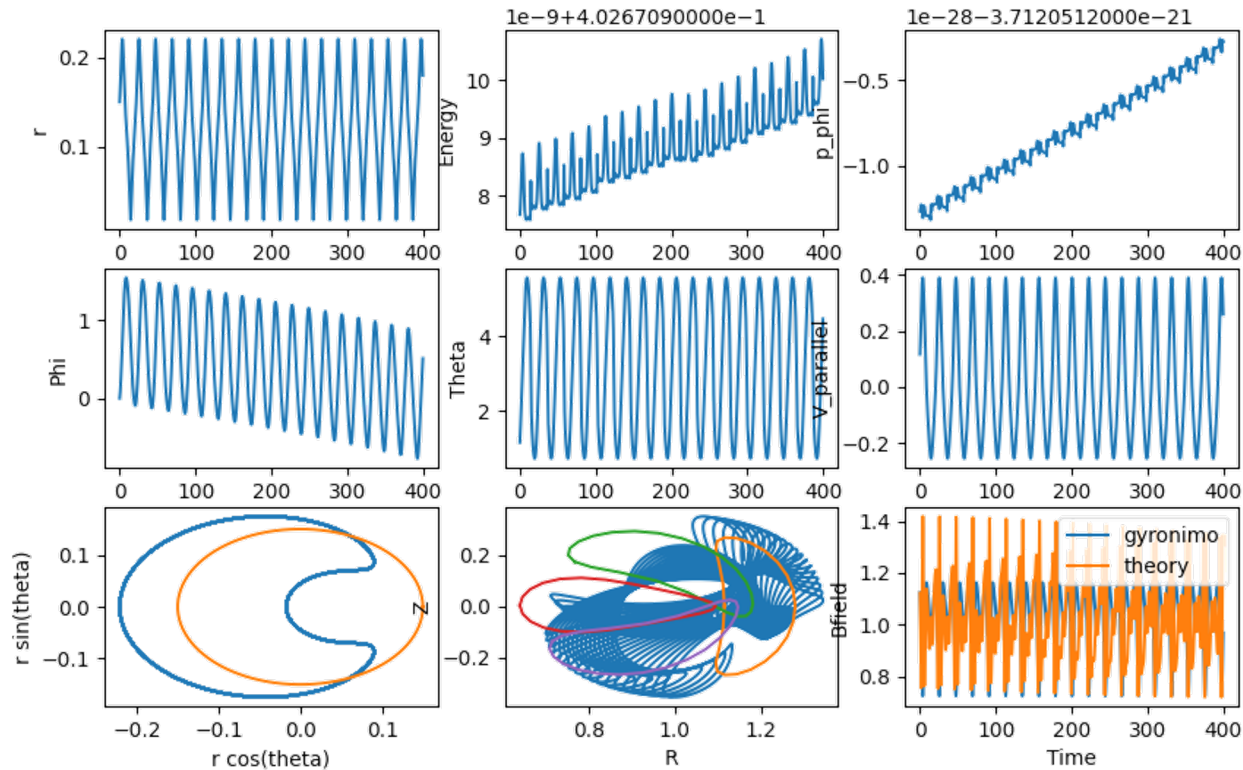


# Preliminary Results

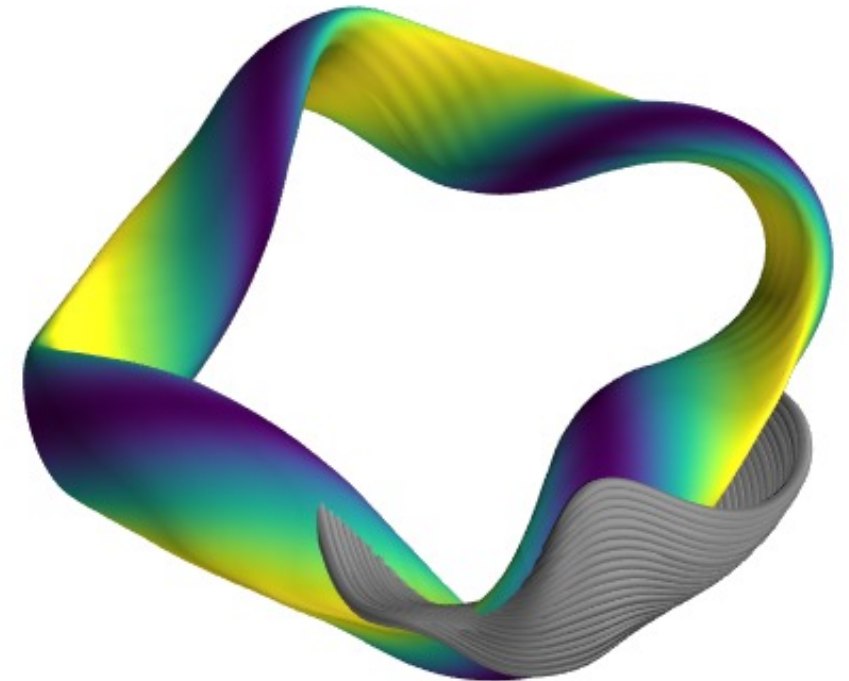
## WPI

C++ Library with documentation on GitHub ✓

Python interface with near-axis expansion coordinates ✓



QH Stellarator

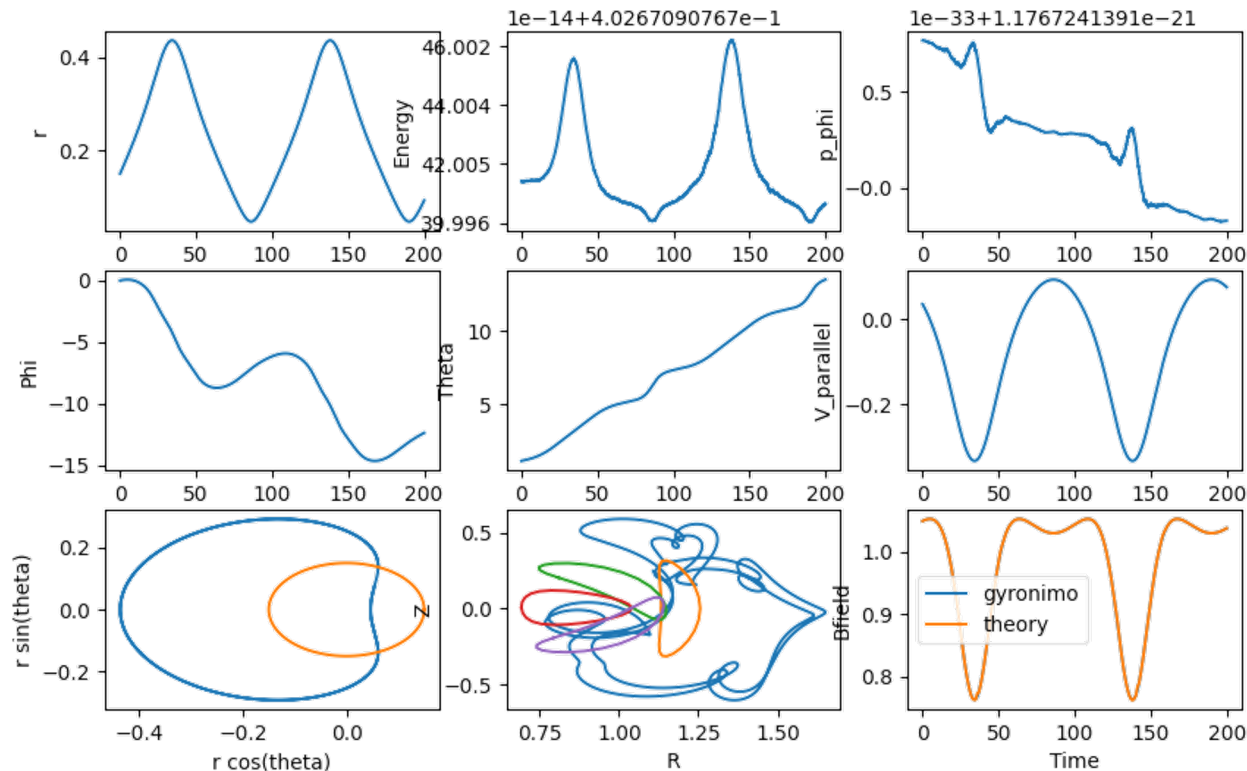


# Preliminary Results

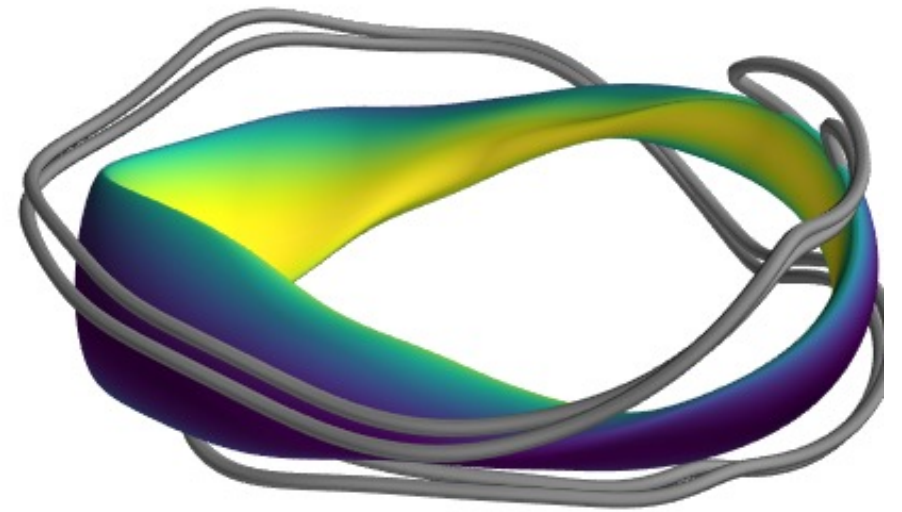
## WPI

C++ Library with documentation on GitHub ✓

Python interface with near-axis expansion coordinates ✓



QA Stellarator

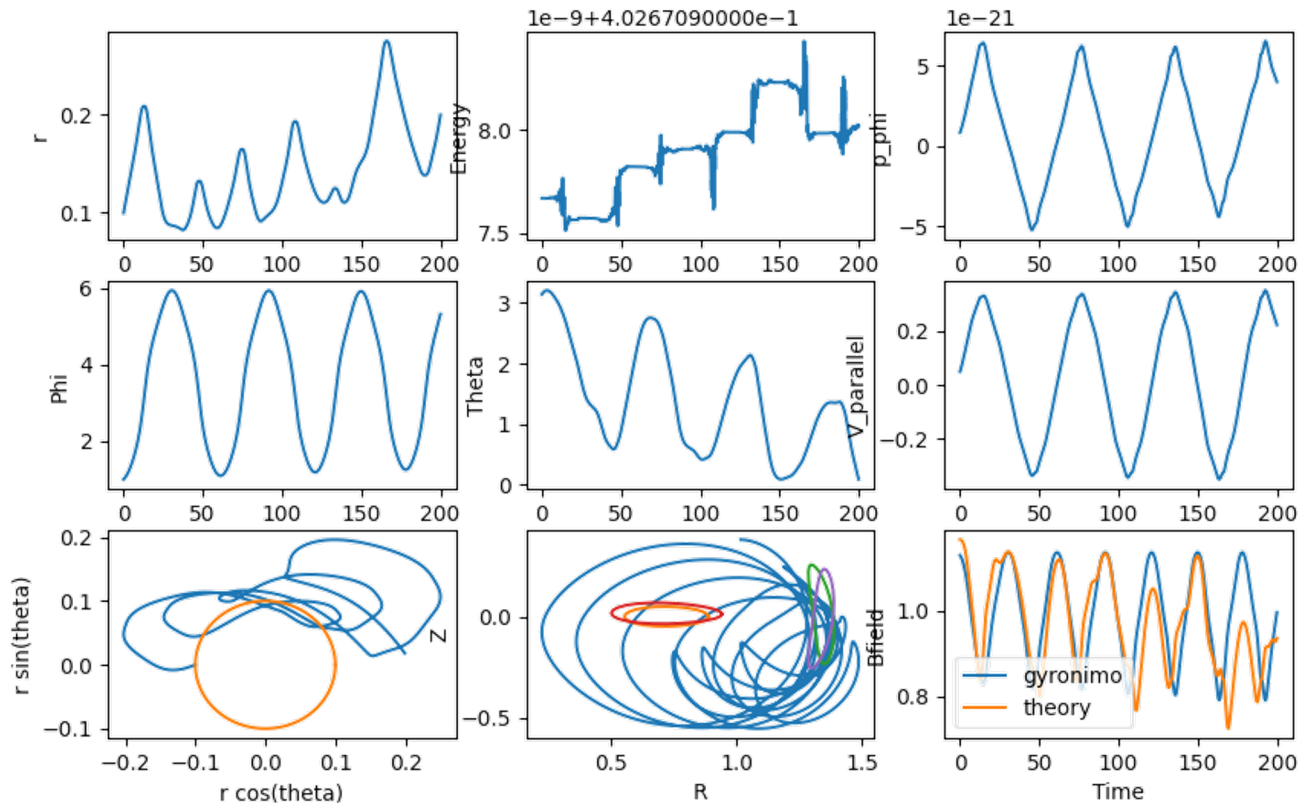


# Preliminary Results

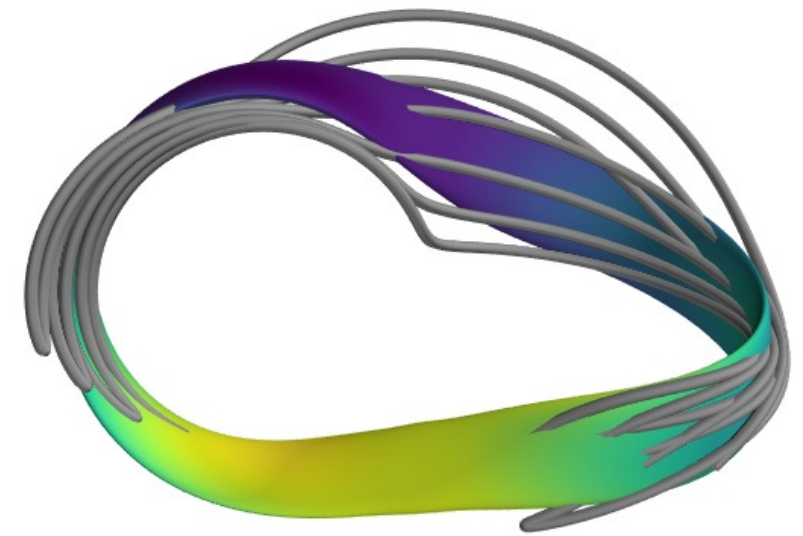
## WPI

C++ Library with documentation on GitHub ✓

Python interface with near-axis expansion coordinates ✓



QI Stellarator



# Preliminary Results

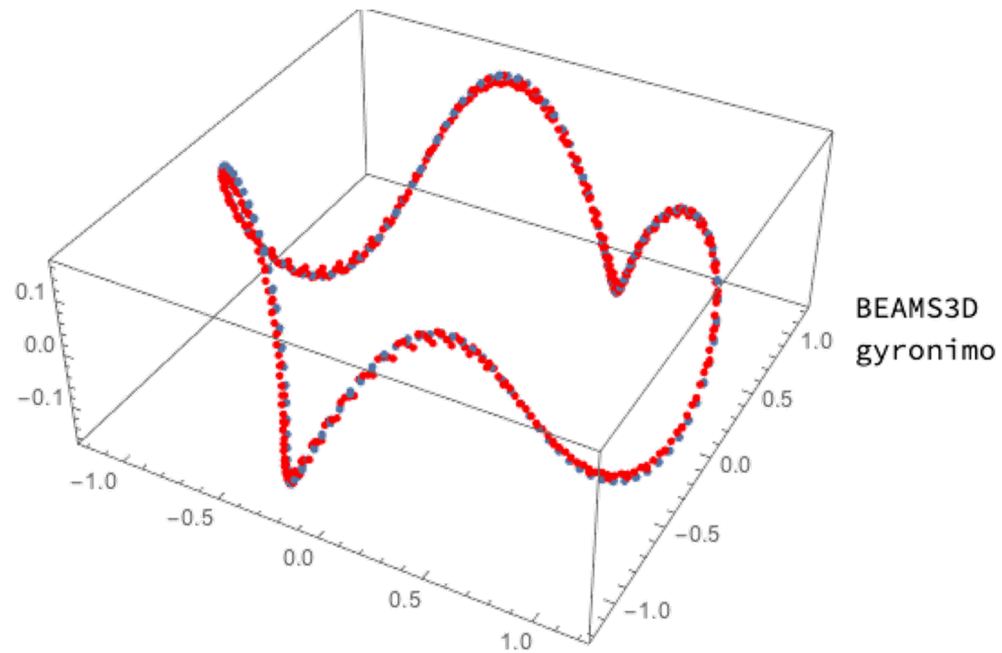
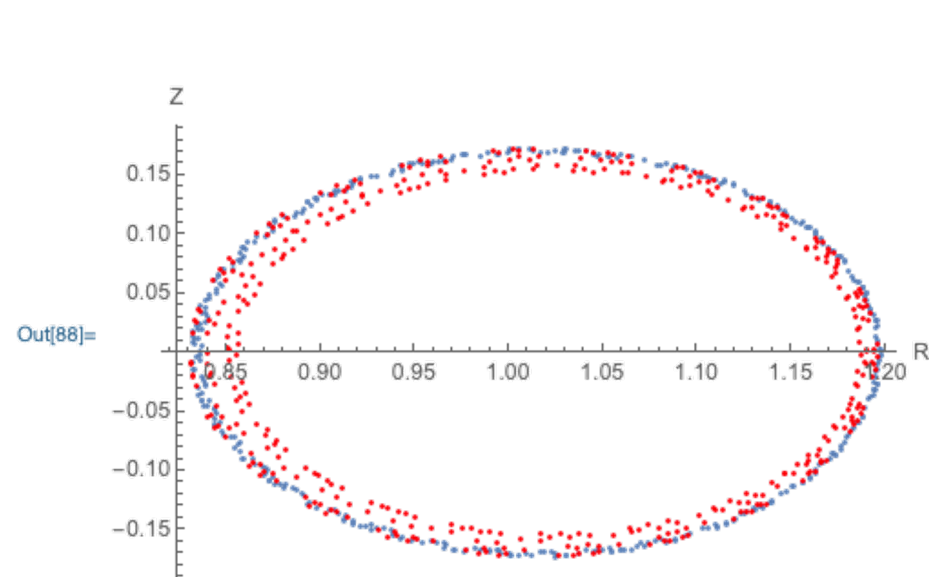
## WPI

C++ Library with documentation on GitHub ✓

Python interface with near-axis expansion coordinates ✓

C++ interface with VMEC equilibria

Benchmark with other particle tracer codes such as BEAMS3D ✓



Need to extend to  $10^5$  particles

# Preliminary Results

## WP2

Analytical/numerical objective function to measure fast particle confinement

Python interface with SIMSOPT/STELLOPT ✓

SIMSOPT requires a `get_dofs` and a `set_dofs` function

In pyQSC

```
def get_dofs(self):  
    """  
    Return a 1D numpy vector of all possible optimizable  
    degrees of freedom, for simsopt.  
    """  
    return np.concatenate((self.rc, self.zs, self.rs, self.zc,  
                           np.array([self.etabar, self.sigma0, self.B2s, self.B2c,  
                                    self.p2, self.I2, self.c0, self.alpha0, self.delta]),  
                           self.B0_vals, self.d_cvals, self.d_svals, self.alpha_cvals, self.alpha_svals))
```

```
def set_dofs(self, x):  
    """  
    For interaction with simsopt, set the optimizable degrees of  
    freedom from a 1D numpy vector.  
    """  
    self.rc = x[self.nfourier * 0 : self.nfourier * 1]  
    self.zs = x[self.nfourier * 1 : self.nfourier * 2]  
    self.rs = x[self.nfourier * 2 : self.nfourier * 3]  
    self.zc = x[self.nfourier * 3 : self.nfourier * 4]  
    self.etabar = x[self.nfourier * 4 + 0]  
    self.sigma0 = x[self.nfourier * 4 + 1]  
    self.B2s = x[self.nfourier * 4 + 2]  
    self.B2c = x[self.nfourier * 4 + 3]  
    self.p2 = x[self.nfourier * 4 + 4]  
    self.I2 = x[self.nfourier * 4 + 5]
```

[https://github.com/rogeriojorge/NearAxis\\_Optimization](https://github.com/rogeriojorge/NearAxis_Optimization)

# Deliverables for 2022

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## **New open-source code**

- Gyronimo ([github.com/prodrigs/gyronimo](https://github.com/prodrigs/gyronimo))
- Extensive documentation and CI/CD

## **Scientific publications**

- Gyronimo oriented study – convergence and verification studies
- Orbit differences between different optimized stellarators

## **Conference presentations**

- EPS 2022
- Varenna 2022

# Possible Revisions for 2022

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## **Parallelization**

Original proposal accounted for both OpenMP and CUDA in the first year

In the first year, only OpenMP may be implemented

Additional code development in 2023

## **Orbits within coils**

Original proposal only had studies with coils at the end of the project

Magnetic fields from Biot-Savart projected to be added in mid-2022

However, realistic coil simulations only in 2023