



# Tackling turbulence from pedestal top to foot with global and local GENE simulations



Leonhard Leppin<sup>1</sup>, Tobias Görler<sup>1</sup>, Marco Cavedon<sup>1</sup>, Mike Dunne<sup>1</sup>, Elisabeth Wolfrum<sup>1</sup>, Frank Jenko<sup>1</sup>, ASDEX Upgrade Team<sup>2</sup>

<sup>1</sup>Max Planck Institute for Plasma Physics, Garching b. München, Germany

<sup>2</sup>See author list of U. Stroth et al. 2022 Nucl. Fusion **62** 042006



This work has been carried out within the framework of the EUROfusion Consortium, funded by the European Union via the Euratom Research and Training Programme (Grant Agreement No 101052200 – EUROfusion). Views and opinions expressed are however those of the author(s) only and do not necessarily reflect those of the European Union or the European Commission. Neither the European Union nor the European Commission can be held responsible for them.

# Intro

## Relates to Key deliverable 1: Gyrokinetic turbulence characterization in H-mode pedestal



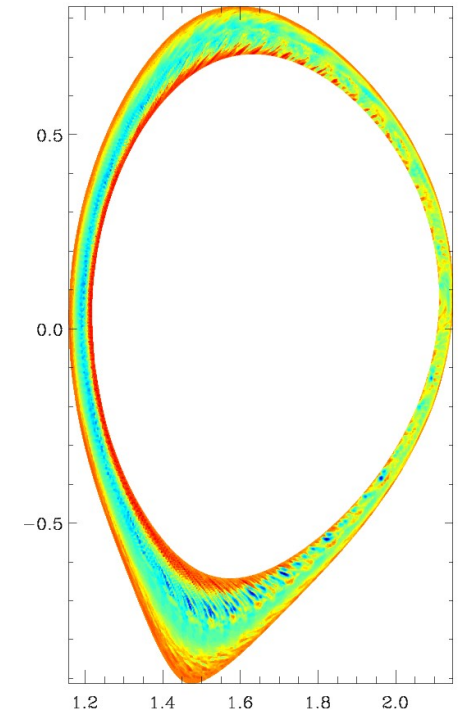
**GENE**

Gyrokinetic code to study small-scale turbulence  
(nonlinear, electromagnetic, Eulerian,  $\delta f$  Vlasov code)

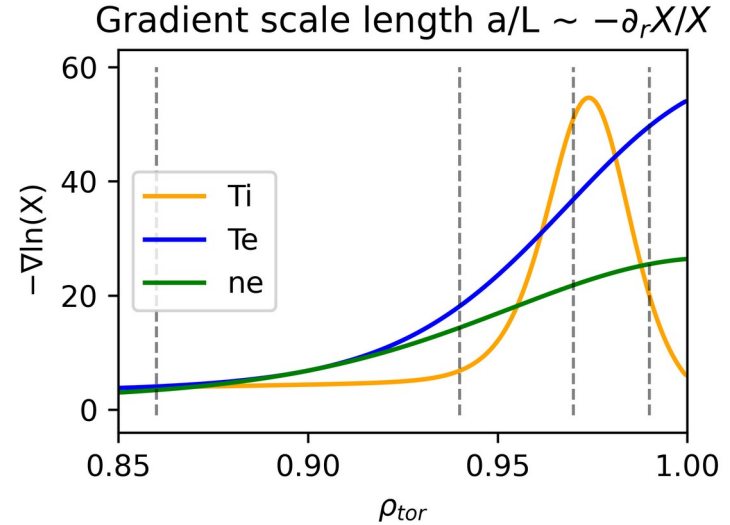
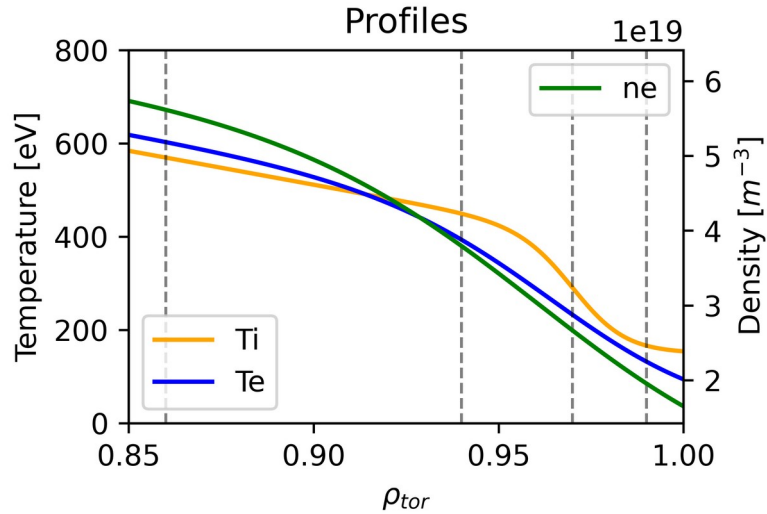


**with upgrade for global,  
nonlinear, experimental  $\beta$  simulations**

**Highlight: Heat flux structure of a pedestal – in radius and scale**



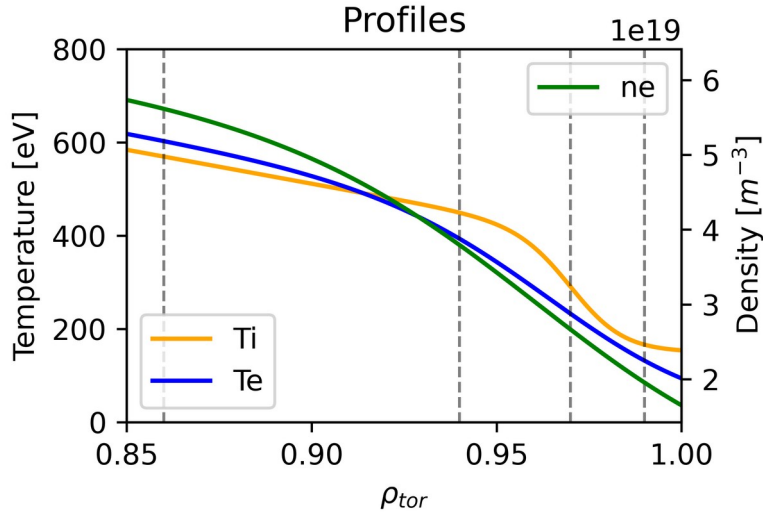
# Scenario: ELMy H-mode pedestal from AUG



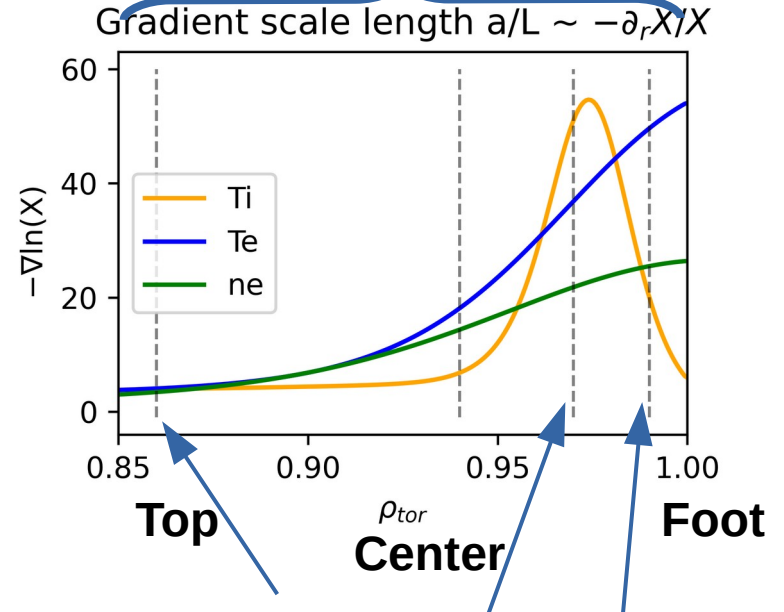
- Asdex Upgrade #31529 [1]
- NBI + ECRH heating,  $P_{tot} \sim 8.7\text{MW}$
- On-axis B-field -2.5 T, plasma current 1MA
- ELM- synchronized profiles (6ms after ELM, almost pre-ELM)
- pressure-constrained magnetic equilibrium

[1] Cavedon et al, PPCF, 2017

# Scenario: ELMy H-mode pedestal from AUG



## Nonlinear, ion scale simulations

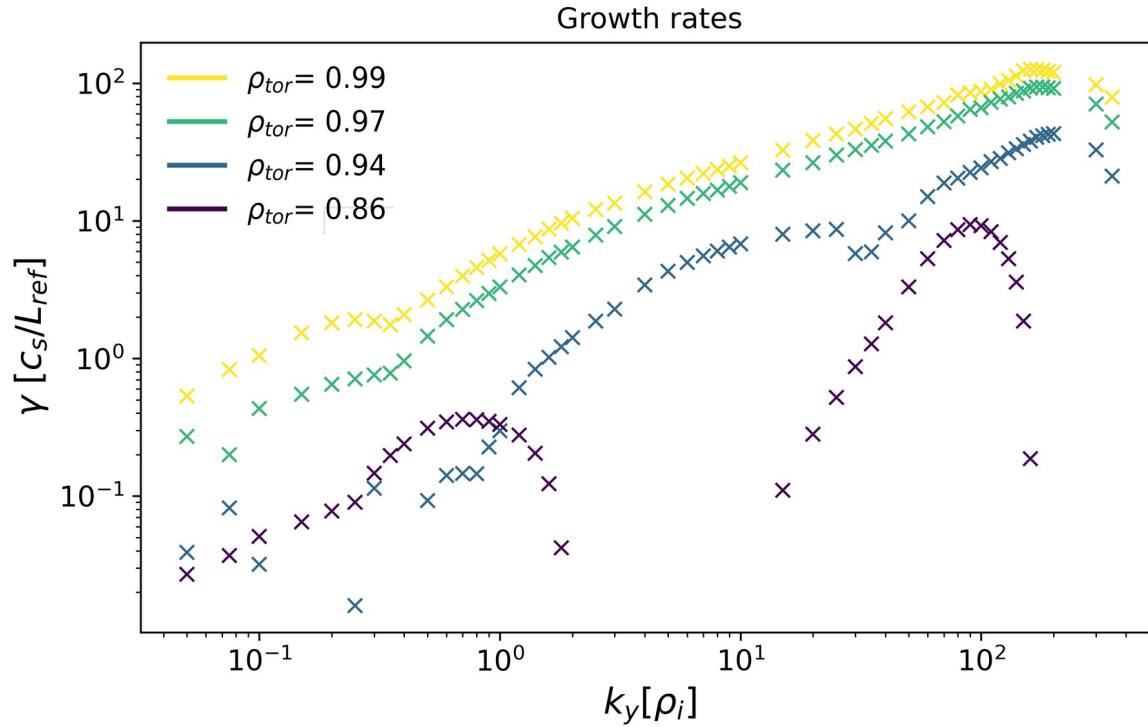


- Asdex Upgrade #31529 [1]
- NBI + ECRH heating,  $P_{\text{tot}} \sim 8.7\text{MW}$
- On-axis B-field -2.5 T, plasma current 1MA
- ELM- synchronized profiles (6ms after ELM, almost pre-ELM)
- pressure-constrained magnetic equilibrium

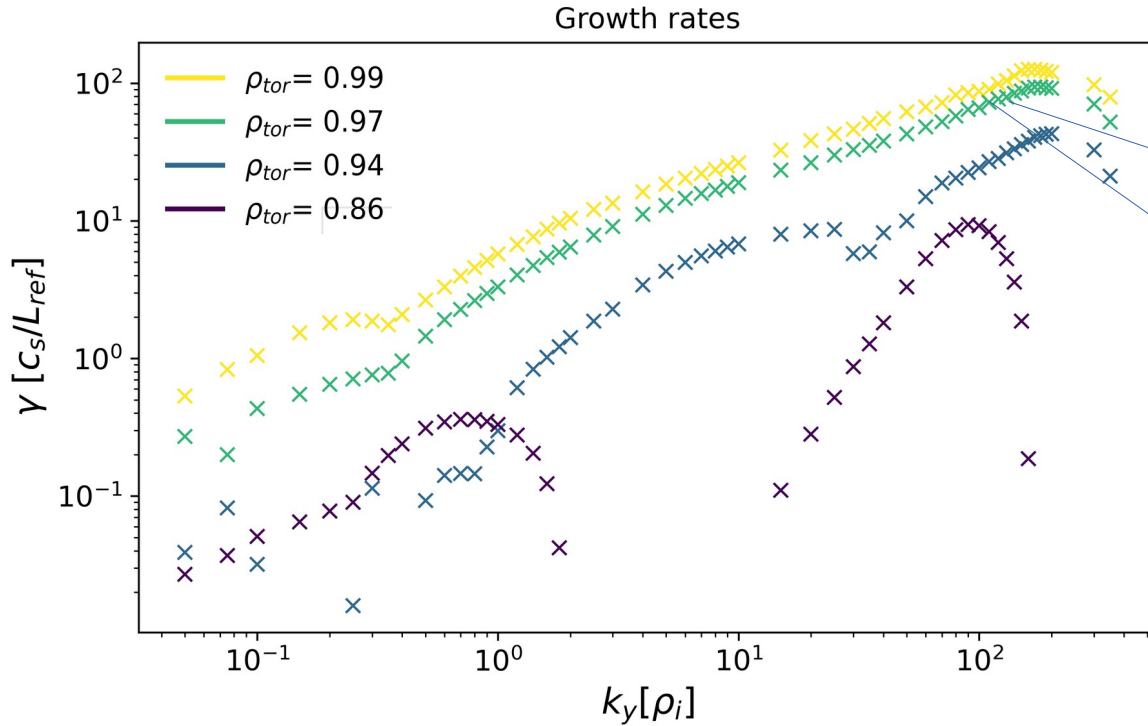
**Instabilities via linear, local sim.**  
**ETG heat flux via nonlinear, local sim.**

[1] Cavedon et al, PPCF, 2017

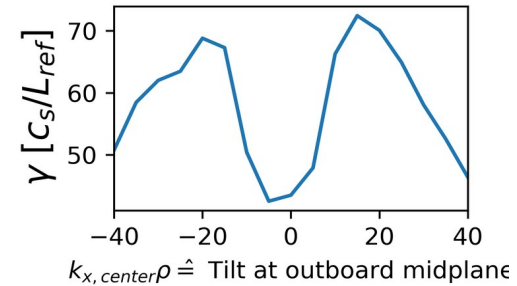
# Linear instabilities



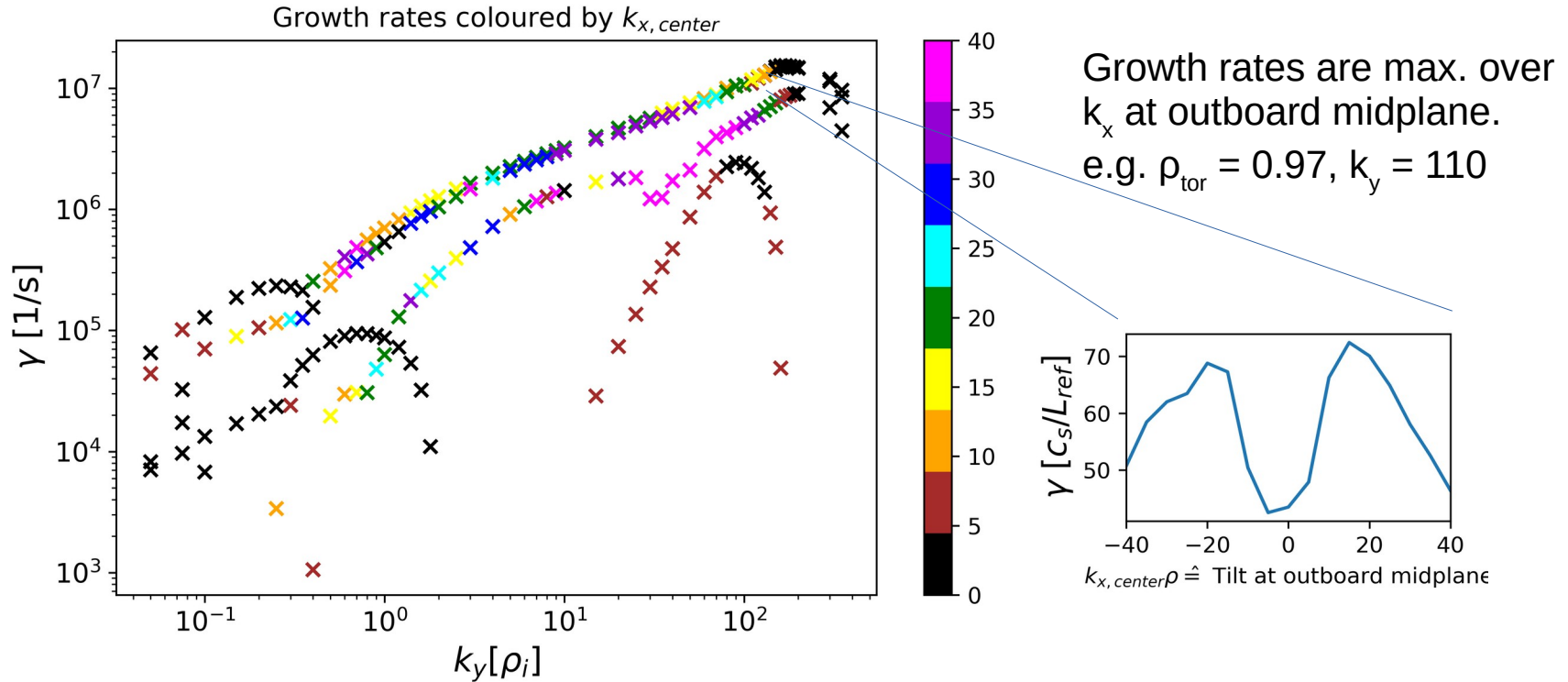
# Linear instabilities



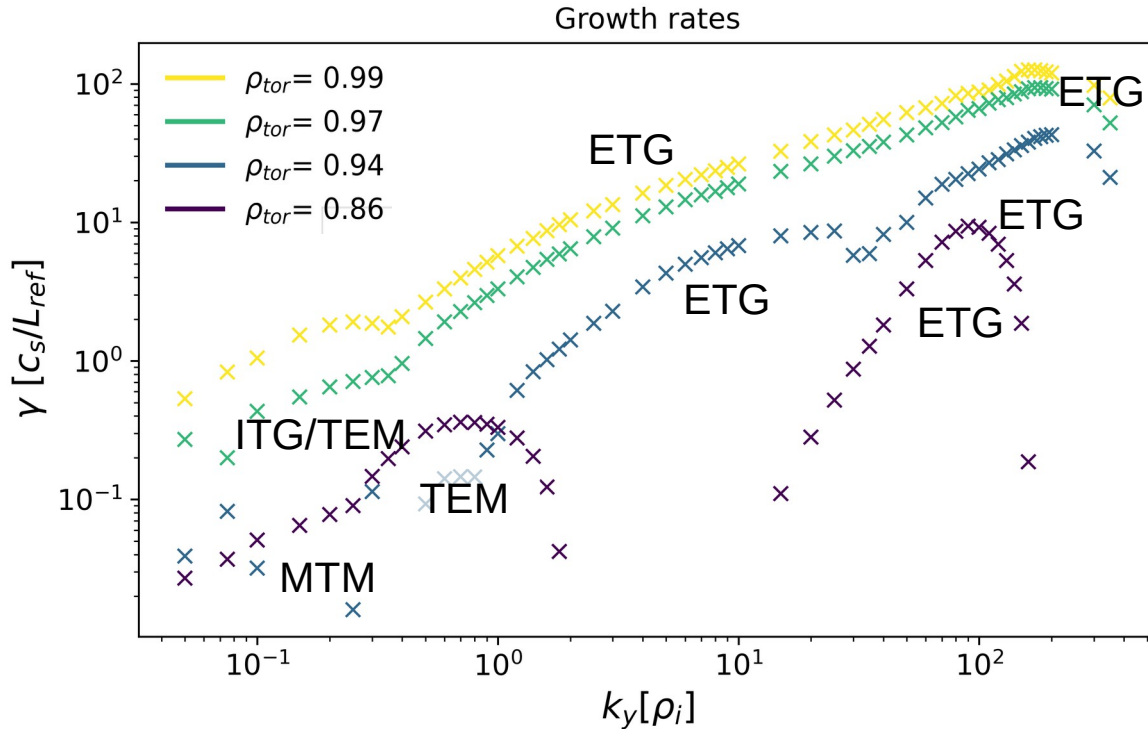
Growth rates are max. over  $k_x$  at outboard midplane.  
 e.g.  $\rho_{tor} = 0.97, k_y = 110$



# Linear instabilities



# Linear instabilities



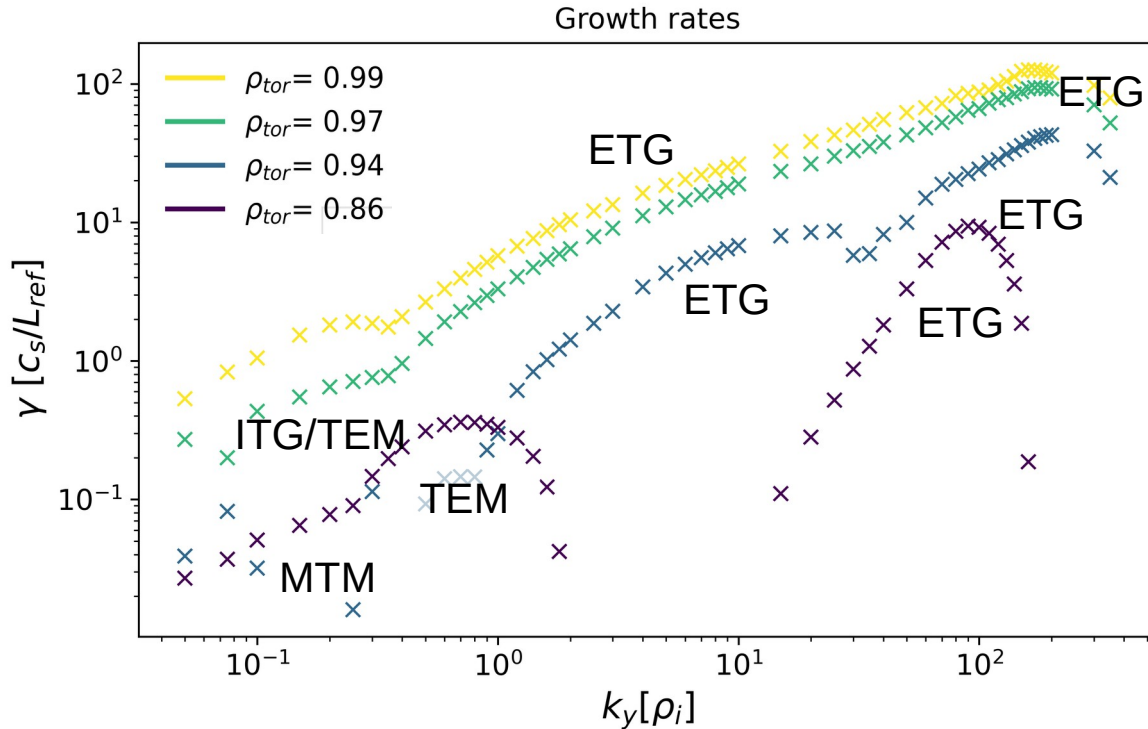
## Identification based on:

- Scale / wavenumber range
- Frequency (drift direction)
- Sensitivity to gradients ( $T_i$ ,  $T_e$ ,  $n$ )
- Sensitivity to plasma  $\beta$  and coll.
- Diffusivity ratios („Fingerprints“ [3])
- Parallel mode structure
- Velocity space structure
- Cross-phases

[3] Kotschenreuther et al, Nucl. Fus., 2019



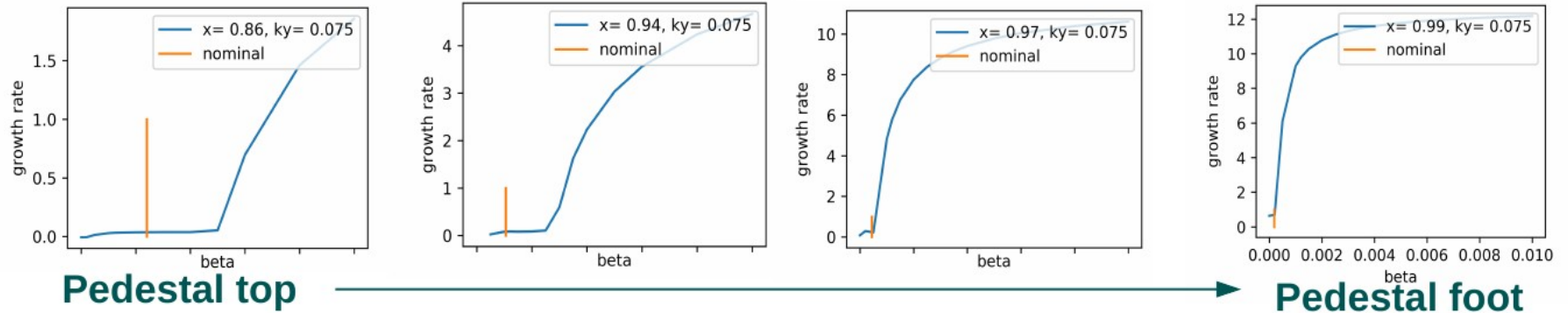
# Linear instabilities



- **Ion scales:**  
Top: TEM/MTM → Center: ITG/TEM  
Growth rate gap at  $\rho_{tor} = 0.94$  (blue)
- **Electron scales: ETG**  
with additional intermediate  $k_y$   
ETG instabilities towards pedestal center
- Overall growth rates increase towards pedestal center/ foot

# Close to linear KBM threshold

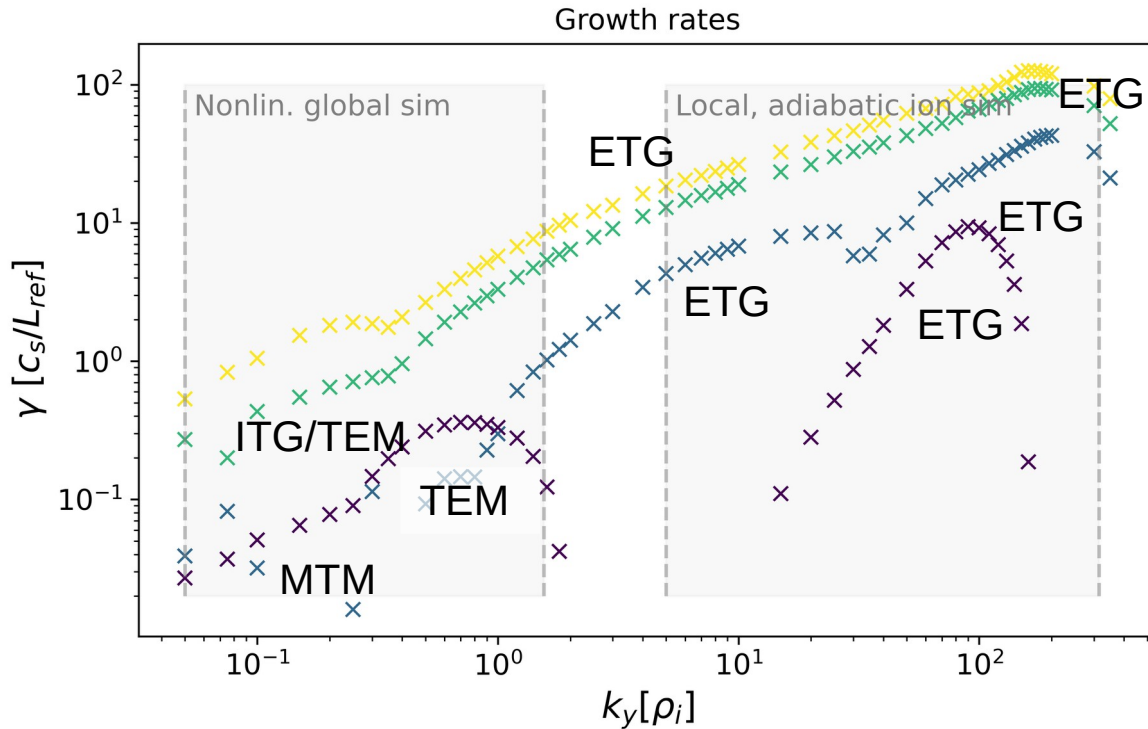
The pedestal is close to a linear KBM threshold. (In agreement with [4])  
Distance decreases towards pedestal foot.



[4] Hatch et al, Nucl. Fus., 2015

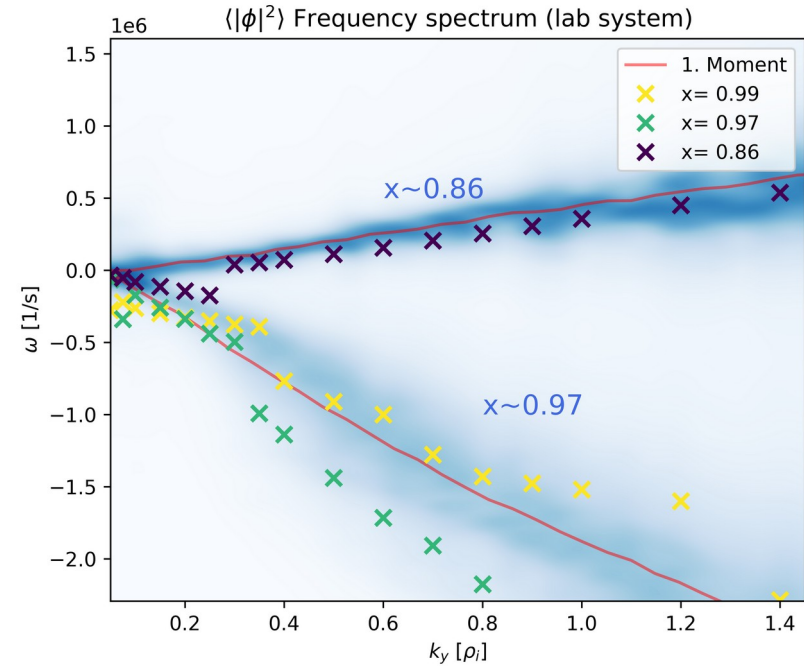
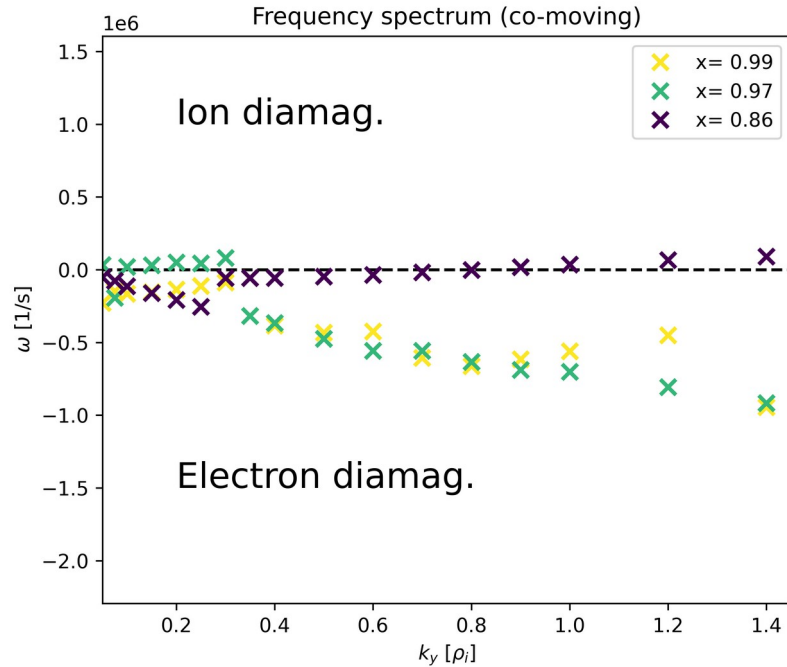
# Linear instabilities

Nonlinear simulation domains in  $k_y$ :



- **Ion scales:**  
Top: TEM/MTM → Center: ITG/TEM  
Growth rate gap at  $\rho_{tor} = 0.94$  (blue)
- **Electron scales: ETG**  
with additional intermediate  $k_y$   
ETG instabilities towards pedestal center
- Overall growth rates increase towards pedestal center/ foot

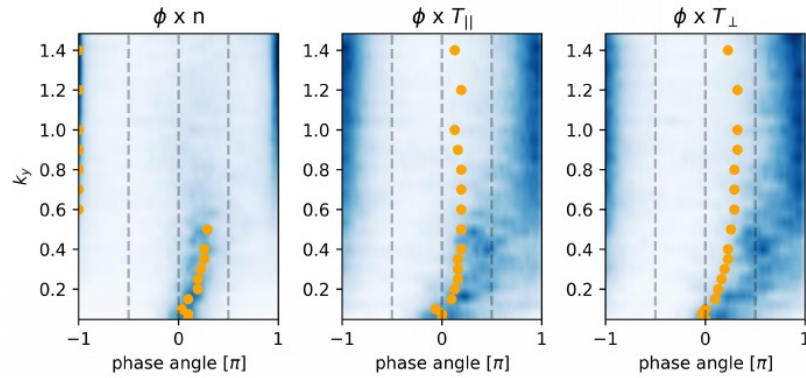
# Connecting linear instabilities and nonlinear modes: Frequencies



→ Linear frequencies remain present at pedestal top and center

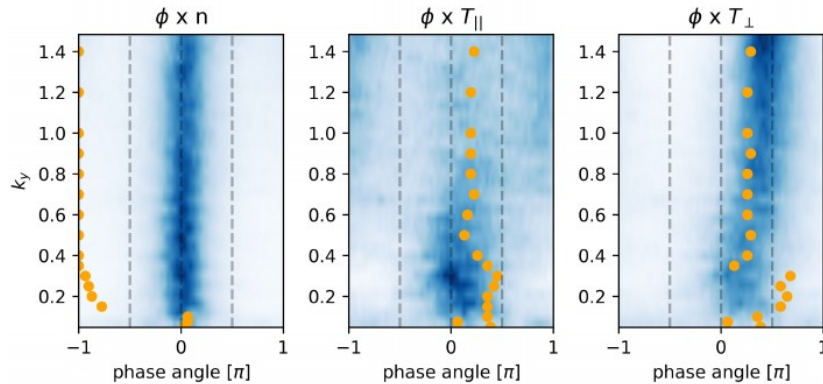
# Connecting linear instabilities and nonlinear modes: Cross phases

Cross phases Electrons (nonlin  $x=0.89$ ; lin  $x=0.88, k_x \text{center}=\text{max}$ )

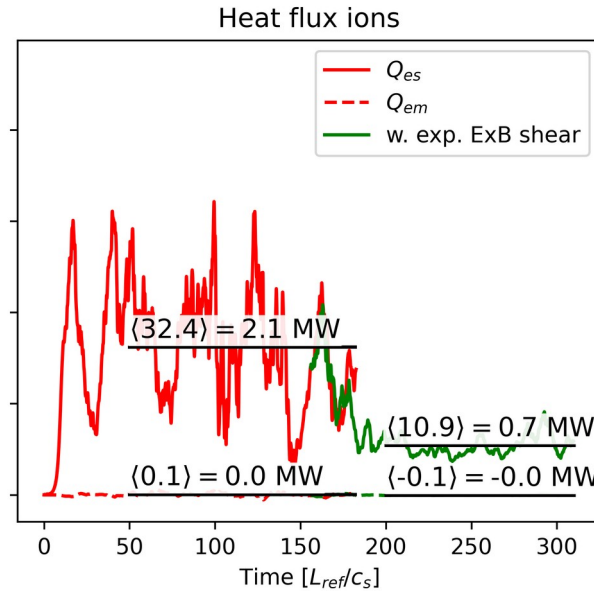
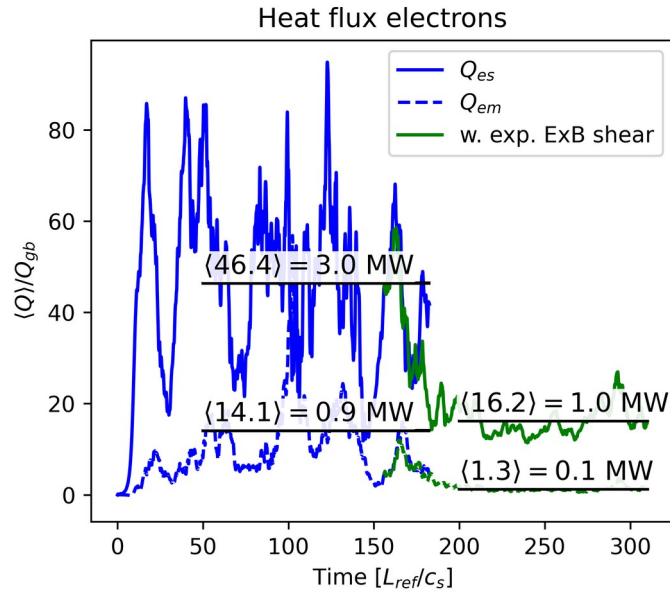


→ Cross phases support that some linear mode characteristics survive in particular at pedestal top

Cross phases Electrons (nonlin  $x=0.97$ ; lin  $x=0.97, k_x \text{center}=\text{max}$ )

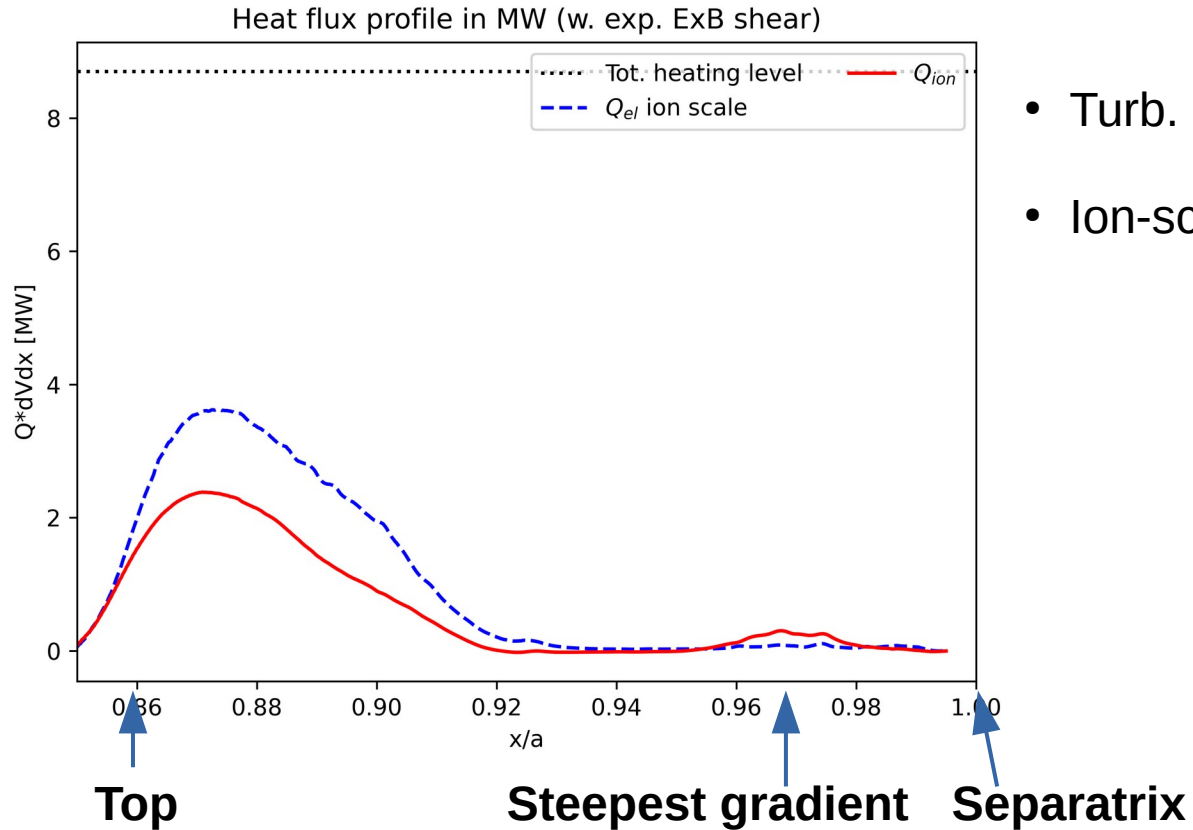


# Global, ion scale: Turbulent heat fluxes



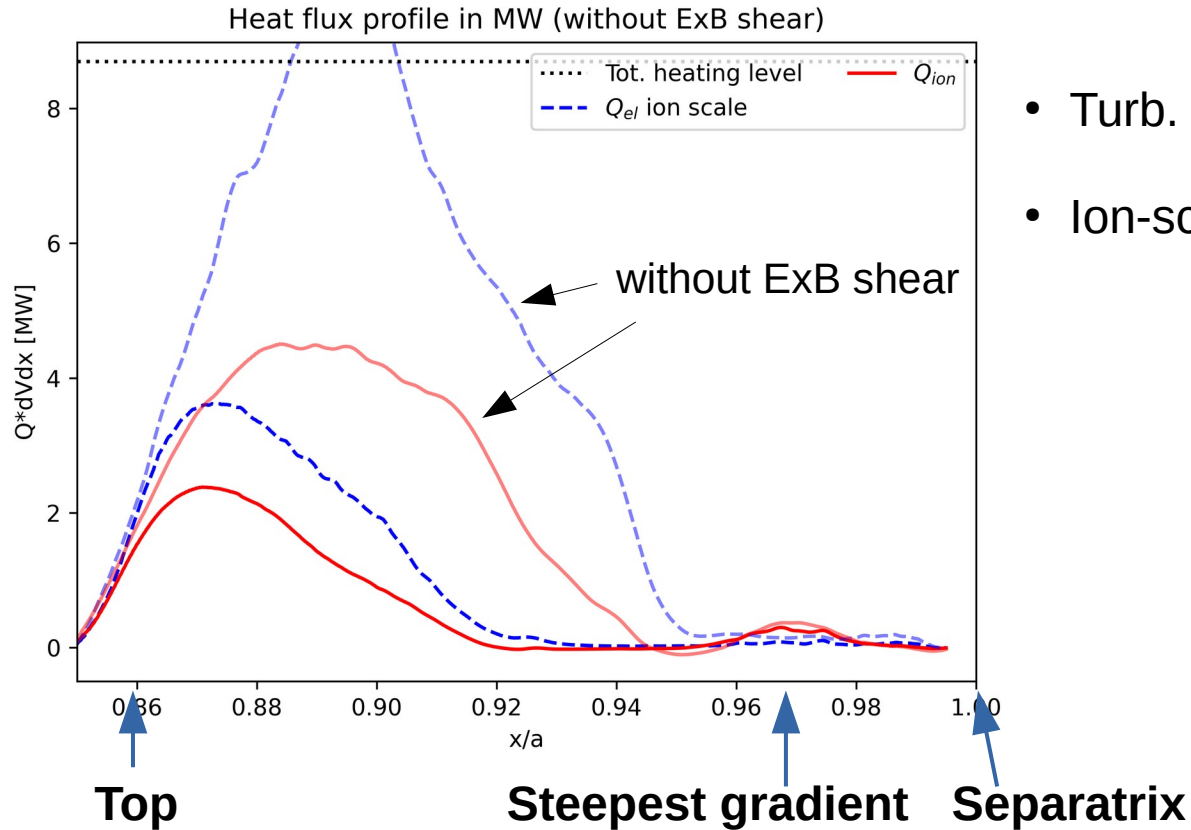
- Simulation is stable and quasi-stationary state is reached
- ExB shear reduces heat fluxes by  $\sim 3$

# Heat flux structure in pedestal



- Turb. ion heat flux vanishes in center
- Ion-scale electron heat flux vanishes as well

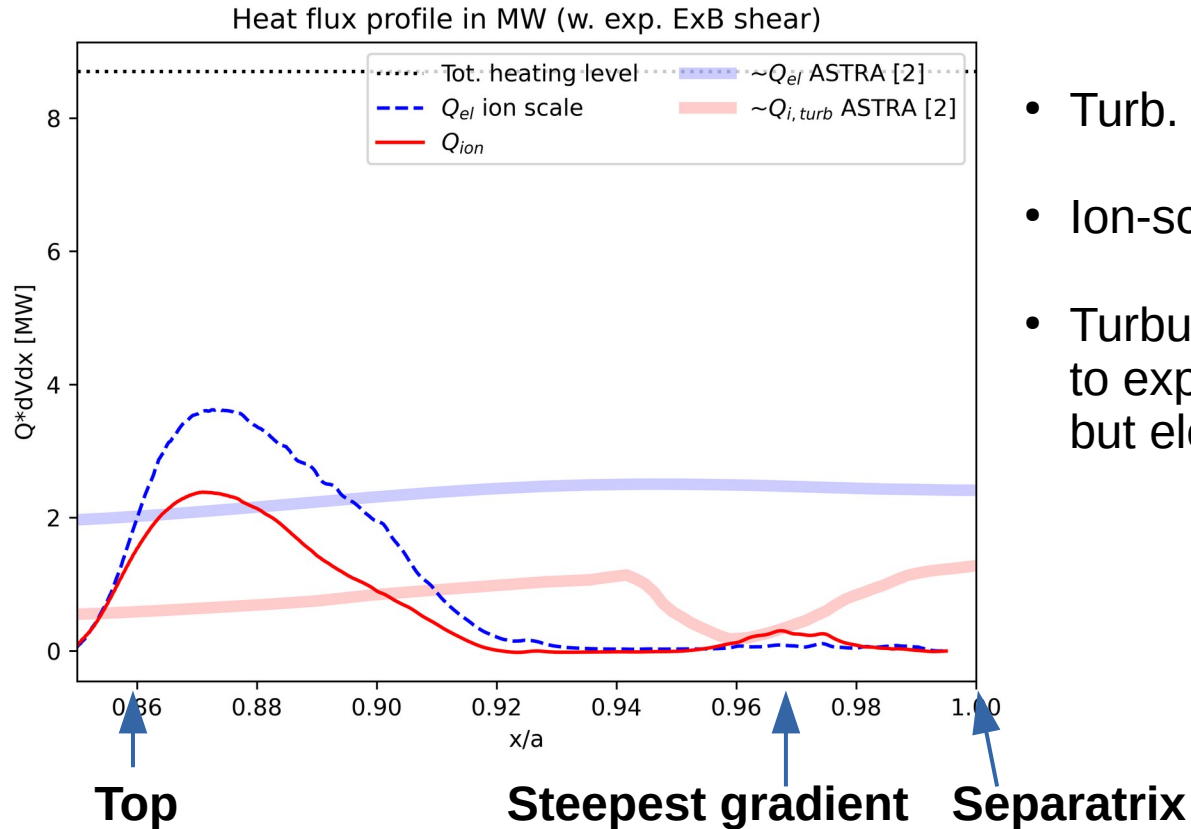
# Heat flux structure in pedestal



- Turb. ion heat flux vanishes in center
- Ion-scale electron heat flux vanishes as well



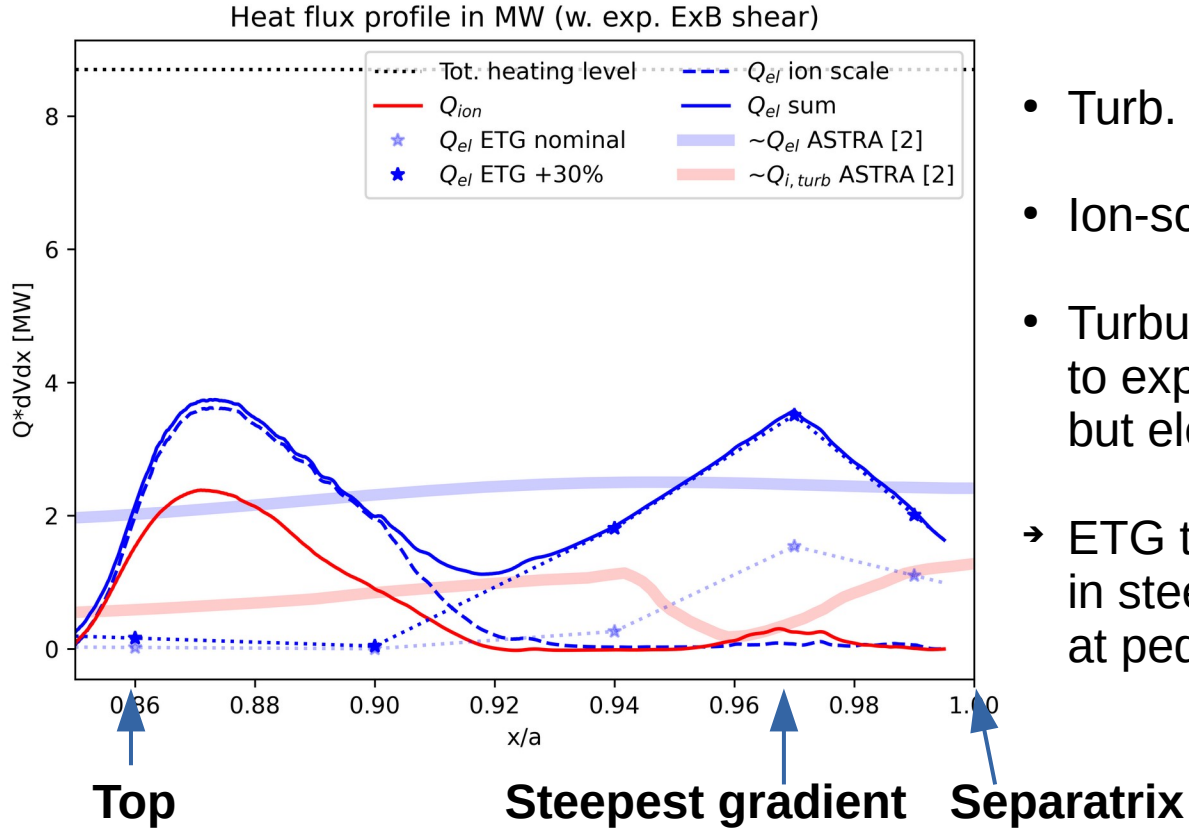
# Heat flux structure in pedestal



- Turb. ion heat flux vanishes in center
- Ion-scale electron heat flux vanishes as well
- Turbulent heat flux levels are comparable to experimental results [2], but electron heat flux in center missing?

[2] Viezzer et al, PPCF, 2020

# Heat flux structure in pedestal



- Turb. ion heat flux vanishes in center
- Ion-scale electron heat flux vanishes as well
- Turbulent heat flux levels are comparable to experimental results [2], but electron heat flux in center missing
- ETG takes over electron heat transport in steep gradient region from TEM at pedestal top

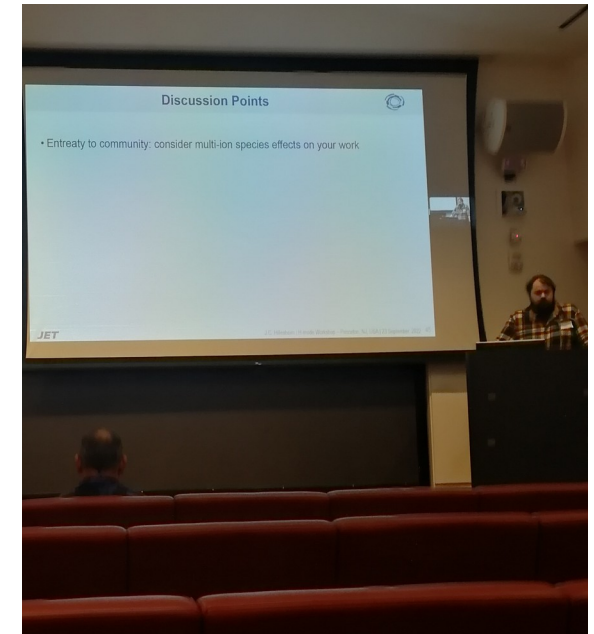
[2] Viezzer et al, PPCF, 2020

# Two subjective observations from



1) RMPs for ELM suppression (Resonant magnetic perturbations for suppression of edge localized modes)  
→ Magnetic non-axisymmetries possibly not only relevant for stellarators but also ITER operation

2) Discussion point by Jon Hillesheim (JET):  
Consider multi-ion species effects on your work  
(eventually we operate D-T plasmas;  
„there is nothing more powerful in science than a testable hypothesis“)





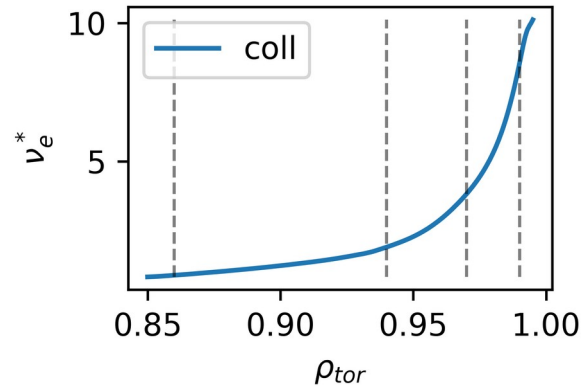
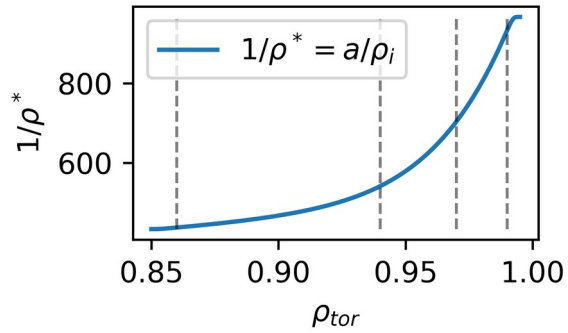
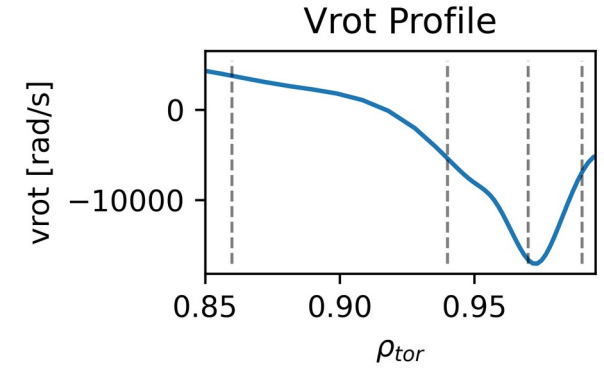
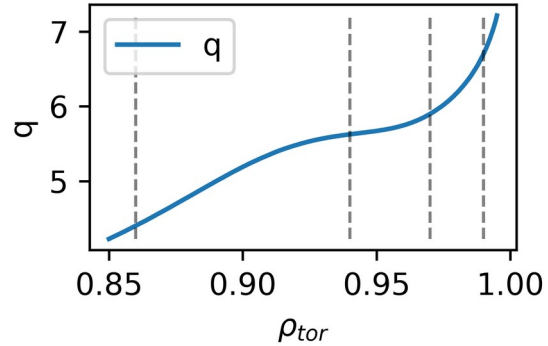
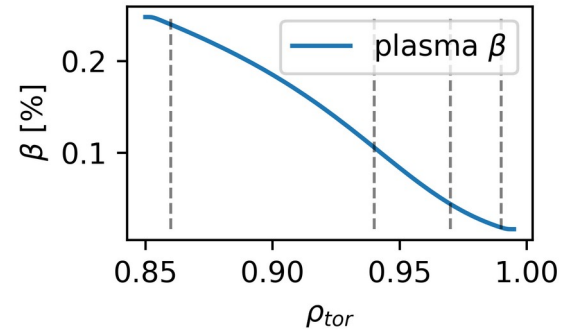
Thank you for your attention!

E-mail: [leonhard.leppin@ipp.mpg.de](mailto:leonhard.leppin@ipp.mpg.de)

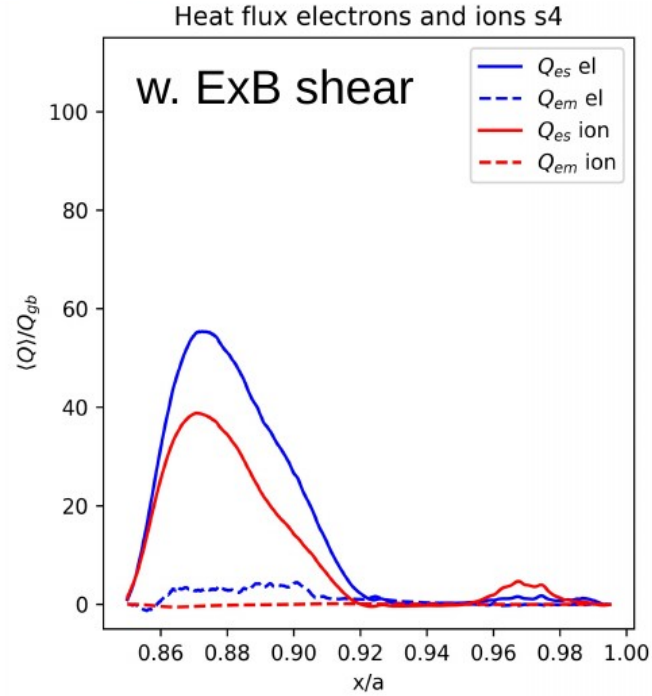
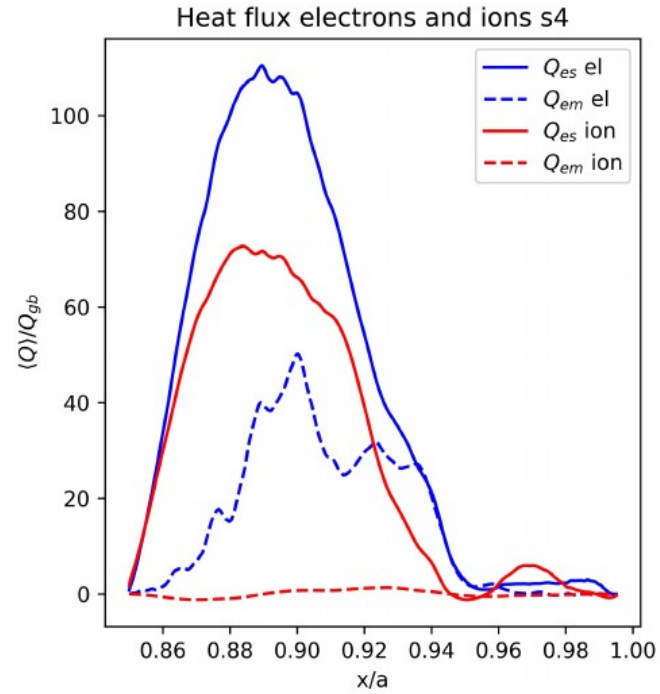


This work has been carried out within the framework of the EUROfusion Consortium, funded by the European Union via the Euratom Research and Training Programme (Grant Agreement No 101052200 — EUROfusion). Views and opinions expressed are however those of the author(s) only and do not necessarily reflect those of the European Union or the European Commission. Neither the European Union nor the European Commission can be held responsible for them.

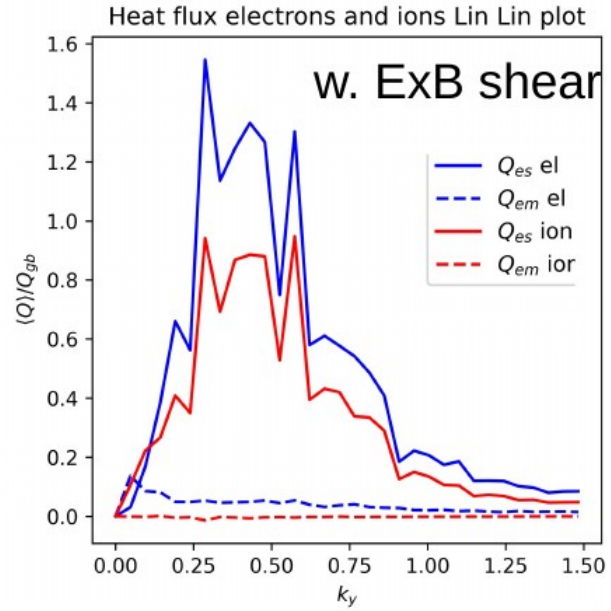
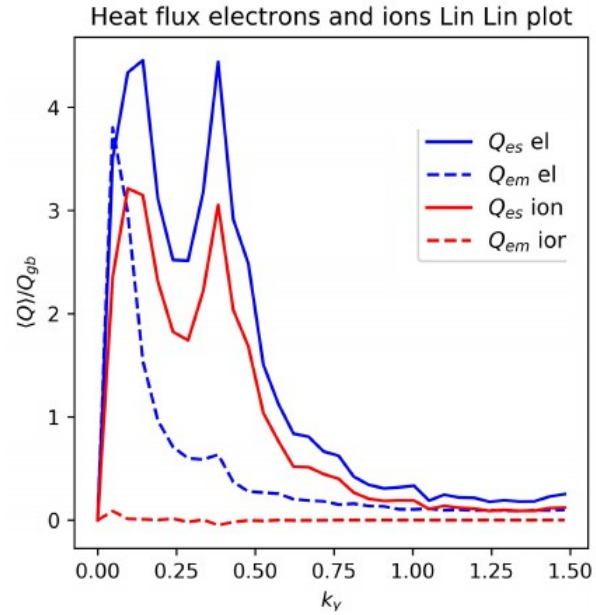
# Other profiles



# Heat flux profile without ExB shear



# Heat flux spectra



# Relation to TSVV Workplan



## Status of macroscopic (MHD-like) instabilities from GENE side



### – Status of implementation

- **General background distribution functions available in GENE due to Alessandro di Siena**
- **Specific implementation of shifted Maxwellian in progress / done by Petch Jitsuk (PhD @ Wisconsin Madison)**



# Recent GENE upgrade

- **In standard GENE:** Collect all temporal derivatives on one side of equation and introduce modified distribution function  $g$ :

$$\frac{\partial f_1}{\partial t} - \frac{q}{mc} \frac{\partial \bar{A}_{1\parallel}}{\partial t} \frac{\partial F_0}{\partial v_{\parallel}} = \dots \quad \& \quad g_1 := f_1 - \frac{q}{mc} \bar{A}_{1\parallel} \frac{\partial F_0}{\partial v_{\parallel}} \quad \longrightarrow \quad \frac{\partial g_1}{\partial t} = \dots$$

- **Problem:** global, nonlinear, electromagnetic simulations with experimental plasma  $\beta$  values tend to be unstable.
- **Solution:** Keep unmodified distribution  $f$  and use Ampere's law  $\nabla_{\perp}^2 A_{\parallel} = -\frac{4\pi}{c} j$  to derive field equation for  $E_{\parallel}^{\text{ind}} = -\frac{1}{c} \frac{\partial A_{\parallel}}{\partial t}$  [5] which can be solved numerically.

$$\left( \nabla_{\perp}^2 + \frac{4\pi}{c^2} \sum_b \frac{q_b^2}{m_b} \int d^3 v \mathcal{G}^{\dagger} v_{\parallel} \frac{\partial F_b}{\partial v_{\parallel}} \mathcal{G} \right) E_{\parallel}^{\text{ind}} = \frac{4\pi}{c^2} \sum_b q_b \int d^3 v \mathcal{G}^{\dagger} \{v_{\parallel} R_b\}$$

- **Implementation:** Fully integrated into GENE master branch and compatible with block-structured velocity space grids

[5] Crandall, PhD Thesis, 2019