



# Evolution of $v_{\perp}$ shear along slow power ramp towards L-H transition

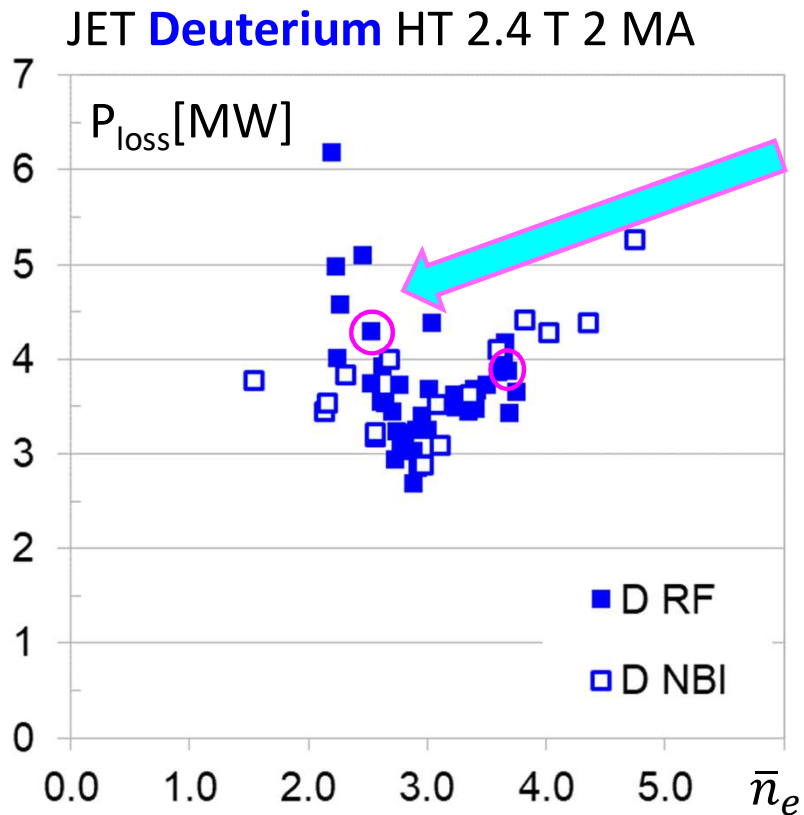
Emilia R. Solano

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# Typical results of $P_{L-H}$ density scan in Deuterium



Focus on L-H transition with very detailed Doppler reflectometry measurements

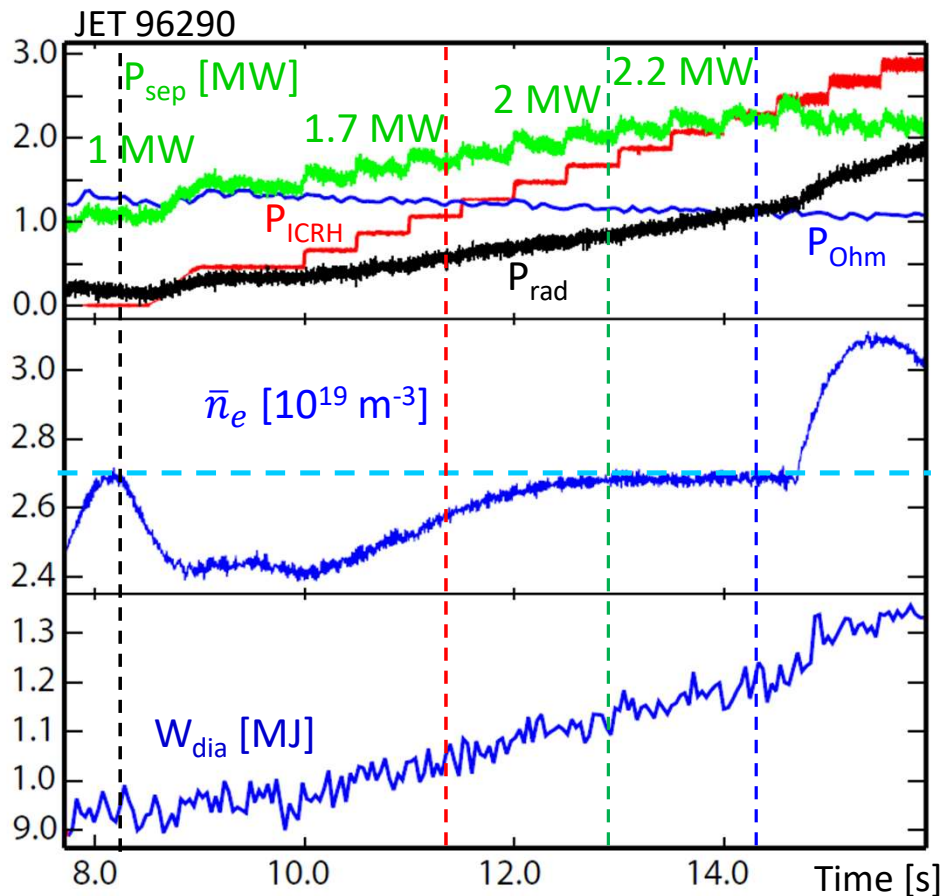
Doppler reflectometry measures  $v_{\perp}$ , rotation of fluctuations perpendicular to B

$v_{\perp}$  shear is exactly what will tear or stretch turbulence eddies', if that is the mechanism for L-H transition

$$v_{\perp} \sim E \times B / B^2$$



## New: evolution of $E_r$ along power ramp in Deuterium



The dominant understanding of the L-H transition:

- the interaction between turbulence and radial electric field leads to an eventual critical electric field (shear) that stabilizes the turbulence  
KC Shaing, Biglari&Diamond, Hahm&Burrell, etc
- The expectation is that the  $E_r$  profile evolves along the power ramp, as  $\nabla p$  or  $\nabla p_i$  [Ryter NF 2014] increases, until a critical  $E_r$  well can stabilise turbulence

ER Solano et al, Nucl. Fusion **62** (2022) 076026

<https://doi.org/10.1088/1741-4326/ac4ed8>

**JET**

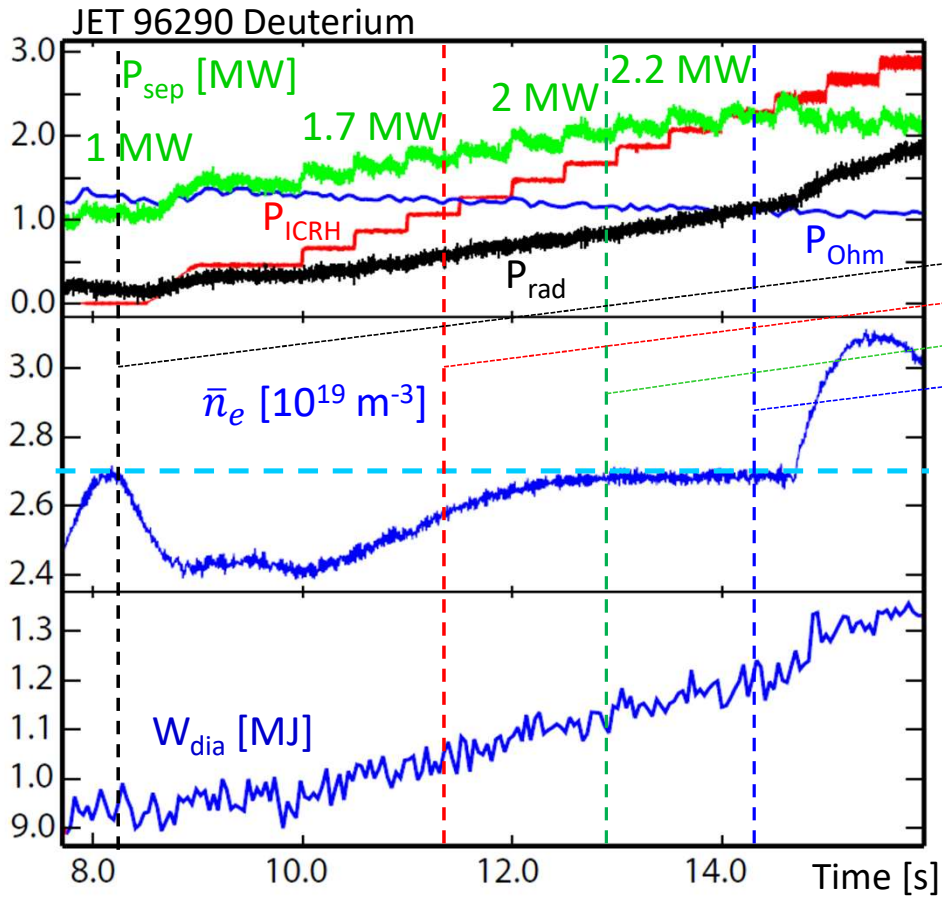
Emilia R. Solano

TSVV1 | Er measurements along power ramp in JET

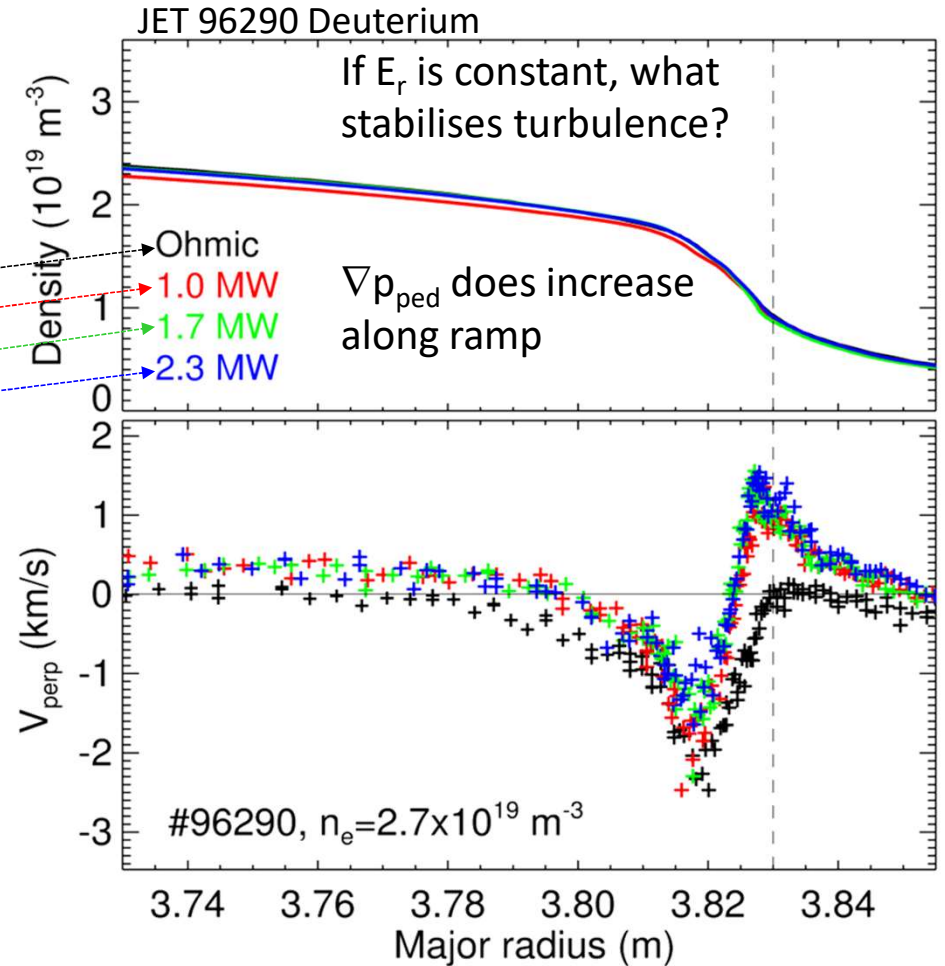
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# Doppler reflectometry along power ramp in Deuterium



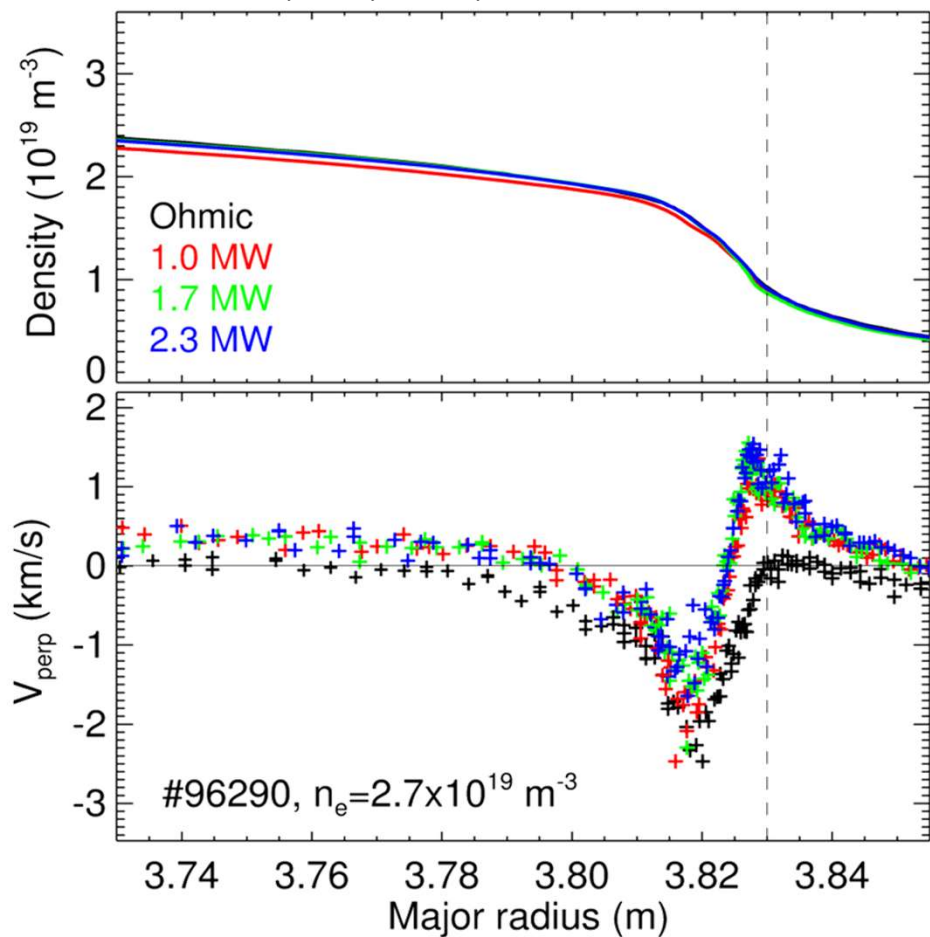
C. Silva, Nucl. Fus. 61 126006 (2021)



# Deuterium: $E_r$ measurements along ICRH power ramp, at $n_{e,\min}$ (D)



JET 96290, 2.4T, 2 MA, ICRH



**Evolution** of  $v_{\perp} \sim E_r/B$  measured with Doppler reflectometry along **especially slow** RF power steps (200 kW every 0.5 s). Time resolution: 300 ms (no momentum input)

**Ohmic:** low  $v_{\perp}$  at separatrix/SOL, deep well

**During power ramp:**

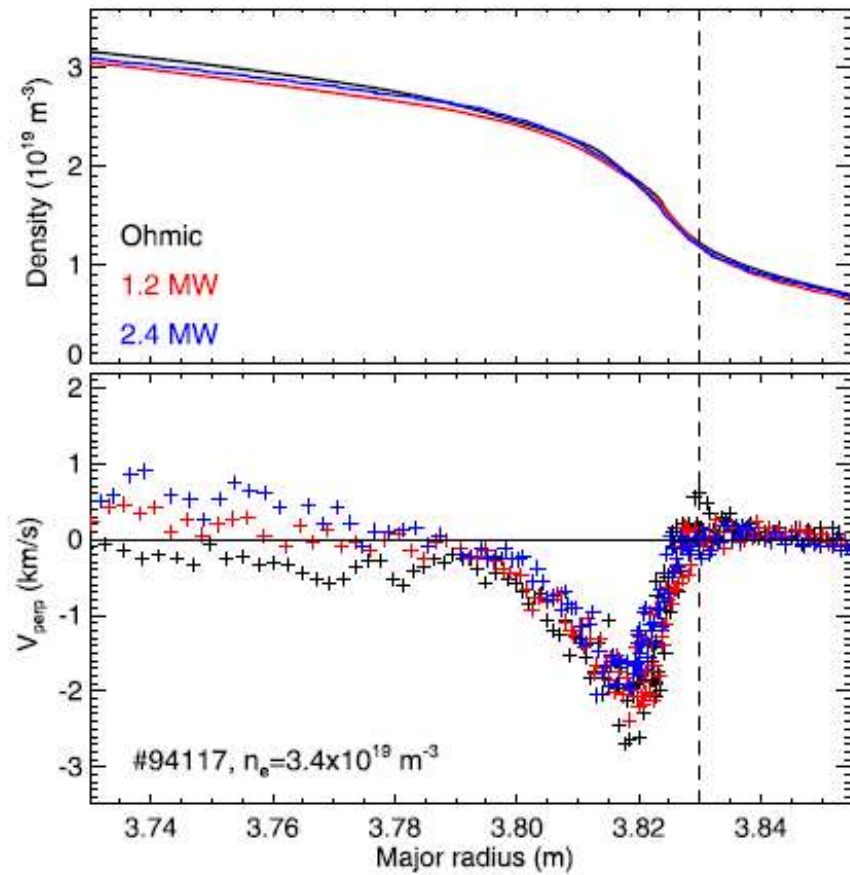
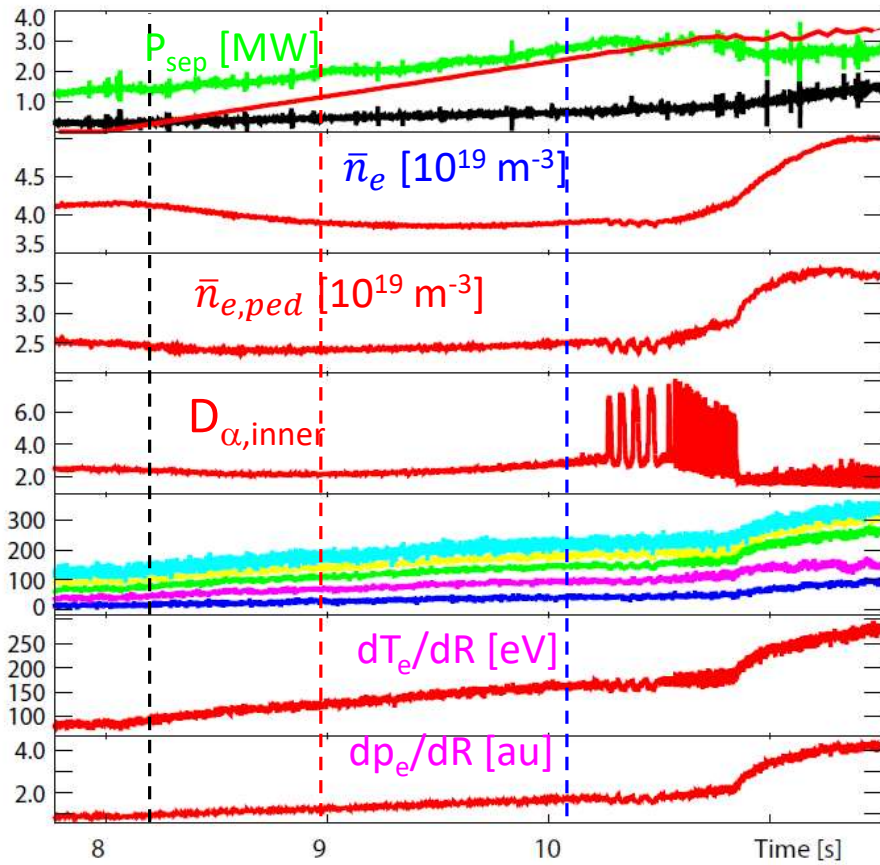
- high  $v_{\perp}$  at separatrix/SOL when ICRH on
- reduction in depth of  $v_{\perp}$  well with ICRH
- similar  $v_{\perp \text{ maximum}}$  shear during power ramp
- L-H: 200 ms after last  $v_{\perp}$  profile, 2.5 MW

**Neither  $E_r$  well nor  $E_r$  shear appear to increase during the power ramp.**

**Is this  $E_r$  shape characteristic of L-mode?**

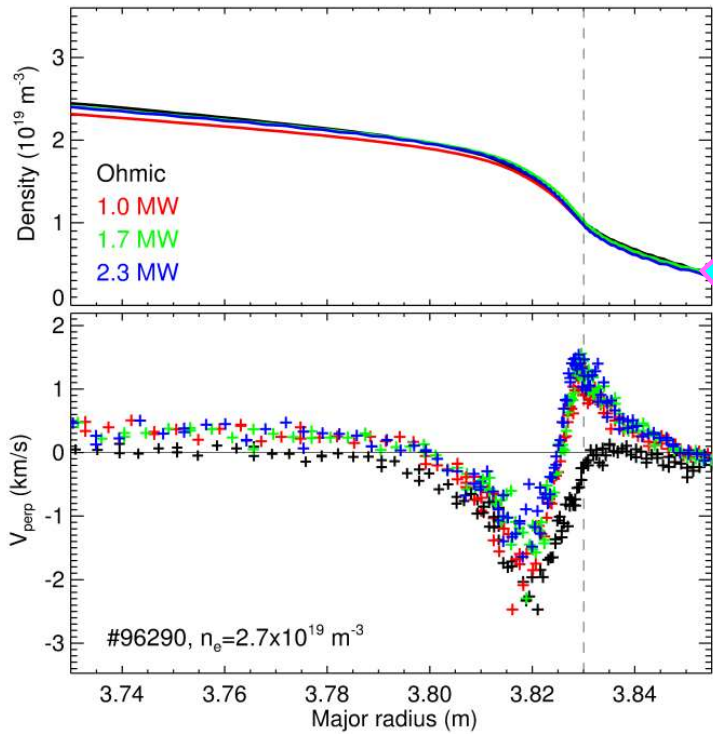
**If so, what triggers the L-H transition?**

# Similar lack of $v_{\perp}$ evolution in high density branch

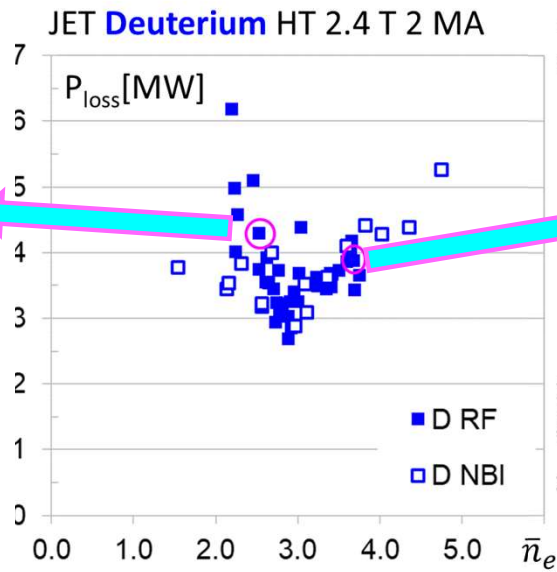




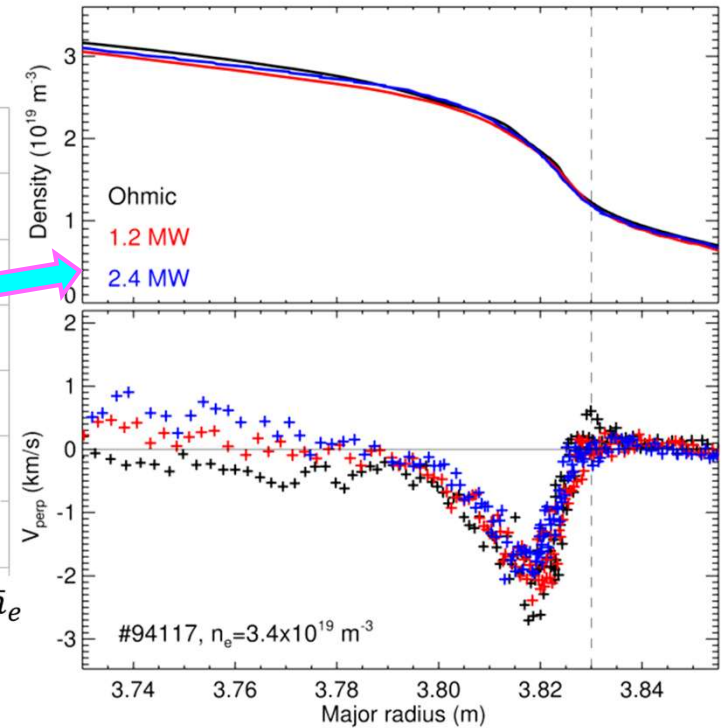
Low density branch



ICRH: no momentum input



High density branch



Same result in high density

# Conclusions




If  $E_r$  profile barely evolves along power ramp, why didn't the transition happen earlier?

We propose  $\nabla p$  or  $F'$  drive a magnetic phase transition from para- to dia-magnetism.

Note: after the L-H transition both  $\nabla p$  and  $E_r$  increase and ExB shear can reduce turbulence and further improve confinement

Technical note:

in the **Grad-Shafranov** equation the sign of  $F'$  (poloidal current density) matters

$$\underbrace{\frac{1}{\mu_0 R} \left( R \frac{\partial}{\partial R} \frac{1}{R} \frac{\partial \Psi}{\partial R} + \frac{\partial^2 \Psi}{\partial Z^2} \right)}_{L(\Psi)} + \underbrace{\left( R p' + \frac{(F^2)'}{2\mu_0 R} \right)}_{J(\Psi)} = 0$$


$F'$  sign affects magnetisation interchange transport, drives magnetization phase separation, maybe L-H transitions.

Different from Shafranov shift dependence  
Related to  $\tilde{B}_{||}$  Facundo Sheffield Heit's talk?

E.R. Solano, *PPCF* **46** (2004) L7–L13

E R Solano, R D Hazeltine (2012) *NF* **52** 114017



# JET L-H experimental results



- Critical profiles  $n_e$ ,  $T_e$ ,  $T_i$  determine access to L-H transition
- $P_{LH}$  depends on isotope due to L-mode isotopic dependencies
  - L-mode  $\tau_E$  scaling is VERY old (1989), needs revisiting, especially isotope effect
- Effective mass orders threshold, but not good scaling parameter
- $v_{\perp}$  profile doesn't evolve along power ramp

## And what about theory?

***Magnetization phase transitions:*** explored in

Equilibrium criticality when  $j_{\theta}=0$ : [ER Solano, PPCF 46 L7 \(2004\)](#)

Diamagnetism and ITB formation: [J Garcia, G Giruzzi PRL 104 205003 \(2010\)](#)

Magnetic phase transition, transport barriers: [ER Solano & RD Hazeltine NF 52 114017 \(2012\)](#)

# Plasma magnetization in a tokamak

Plasma force balance:

$$\nabla p = \vec{j} \times \vec{B} = \vec{j}_\zeta \times \vec{B}_\theta + \vec{j}_\theta \times \vec{B}_\zeta$$

In cylindrical approximation

$$\frac{d}{dr} \left( p + \frac{B_z^2 + B_\theta^2}{2\mu_0} \right) = -\frac{B_\theta^2}{r\mu_0}$$

$$\frac{dB_z}{dr} = -\mu_0 j_\theta$$

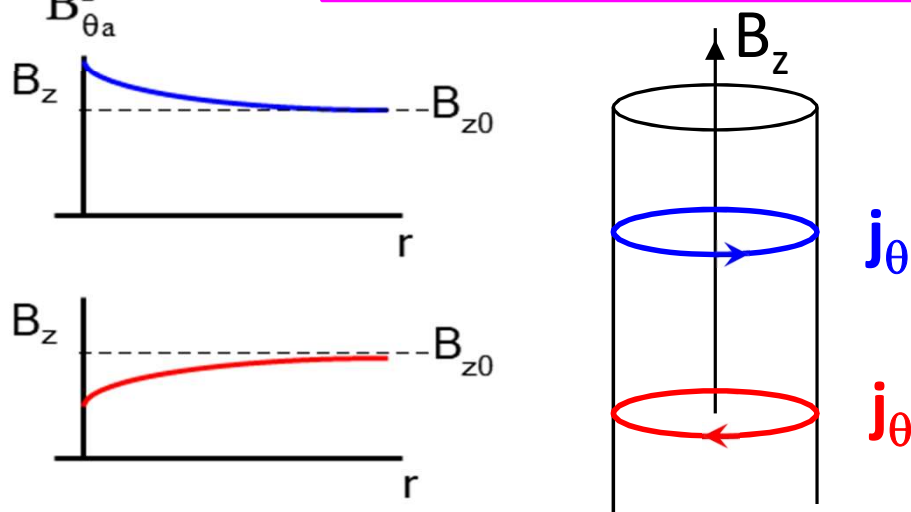
Integrating:

$$\beta_\theta \equiv \frac{\int_0^a p dS}{B_{\theta a}^2 / 2\mu_0} = \frac{B_{za}^2 - \langle B_z^2 \rangle}{B_{\theta a}^2} \simeq 1 + \frac{2B_{za} (B_{za} - \langle B_z \rangle)}{B_{\theta a}^2}$$

$\beta_\theta$  related to volume averaged plasma magnetization

$\beta_\theta < 1$   $B_z$  increased by  $j_\theta$ ,  
paramagnetism,  
low pressure

$\beta_\theta > 1$   $B_z$  reduced by  $j_\theta$   
diamagnetism,  
high pressure



# Plasma Magnetisation

The tokamak plasma is a magnet

$$\langle B_z \rangle - B_z^{\text{vac}} \cong \mu_0 \left( B_{\theta a}^2 / 2\mu_0 - \int_0^a p dS \right) / B_z^{\text{vac}}$$

The difference between poloidal magnetic and kinetic pressure determines if it is a **para**-magnet or a **dia**-magnet

## Paramagnets

increase the background magnetic field  
move towards high field

## Diamagnets

increase the background magnetic field  
move towards low field

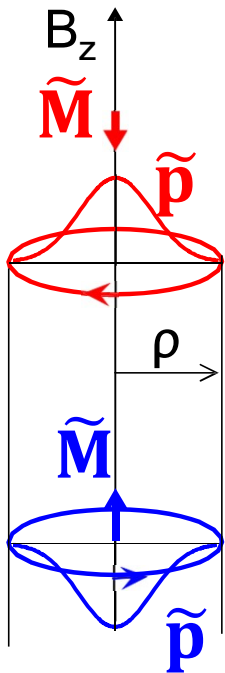


*Diamagnetic frog levitating  
in magnetic field*

Berry, Geim (Ig-Nobel)  
*Eur. J. Phys.* **18** 307 1997

# Magnetism in cylindrical blob with pressure **peak/hole**

$$\mathbf{F} = mn \frac{d\mathbf{v}}{dt} = -\nabla \tilde{p} + \tilde{\mathbf{j}} \times \mathbf{B} = 0 \quad \tilde{\mathbf{j}}_{\perp} = \frac{\mathbf{b} \times \nabla \tilde{p}}{B}$$



**Diamagnetic** current: if inside the tube there is a pressure **peak**, the associated  $\tilde{\mathbf{j}}_{\perp}$  **reduces**  $B_z$ : **diamagnetism**

**Paramagnetic** current: if inside the tube there is a pressure **hole**, the associated  $\tilde{\mathbf{j}}_{\perp}$  **increases**  $B_z$ : **paramagnetism**

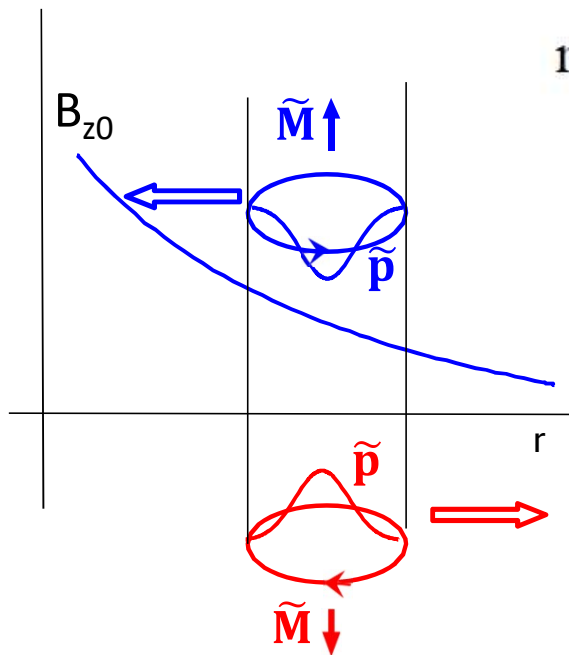
## Magnetization of the blob:

$$\nabla \times \mathbf{M} = \mu_0 \frac{\mathbf{b} \times \nabla \tilde{p}}{B} = -\frac{dM}{dr} \hat{\mathbf{r}}$$

$$\tilde{\mathbf{M}} = \frac{1}{\lambda_{\parallel}} \int_0^{\rho} \frac{\mathbf{b}}{B} \frac{\partial \tilde{p}(\rho')}{\partial \rho'} \lambda_{\parallel} d\rho' \approx -\frac{\tilde{p}}{B} \mathbf{b} \quad \left\{ \begin{array}{l} < 0, \text{ dia} \\ > 0, \text{ para} \end{array} \right.$$

# Movement of magnetised blobs in paramagnetic plasma

Jackson, Classical Electrodynamics



$$m_V \frac{d\mathbf{v}}{dt} \Big|_V = \int (\nabla(\tilde{\mathbf{M}} \cdot \mathbf{B})) dV \simeq \left( \int (\mathbf{r} \times \mathbf{j}_{\text{mag}}) dV \right) \int \nabla B_{0z} dV$$

blob magnetization

$$\vec{B} = \vec{B}_0 + \vec{r} \cdot \nabla \vec{B}_0 + \dots$$

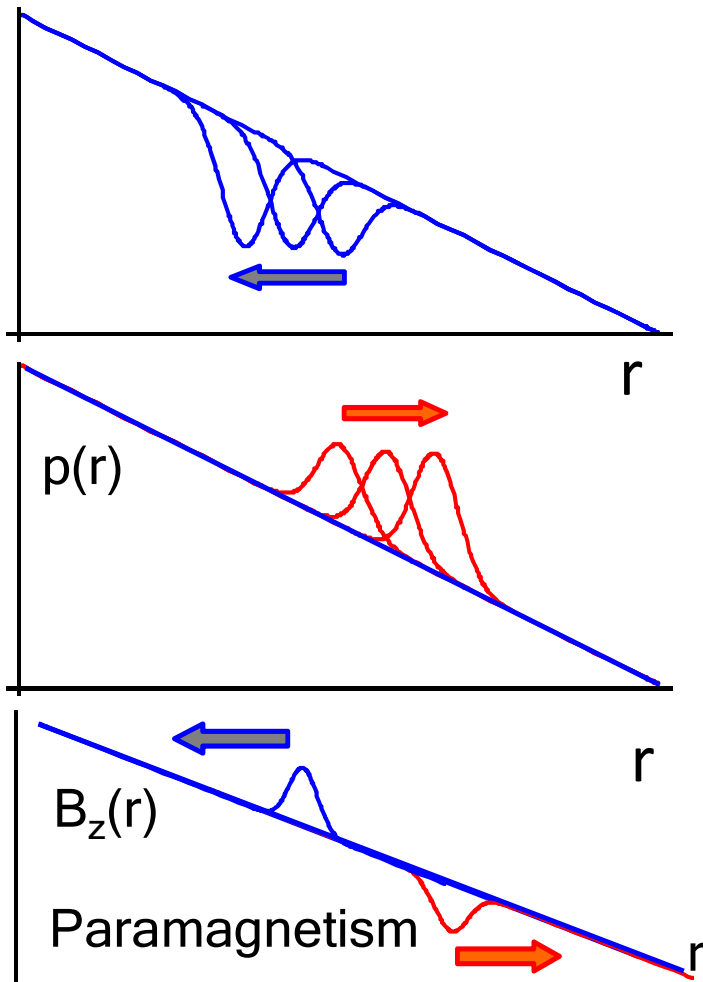
$$m_V \frac{d\mathbf{v}_r}{dt} \simeq \tilde{M}_z \nabla \bar{B}_{z0}$$

the **cold blob** (paramagnetic) seeks **high field**

the **hot blob tube** (diamagnetic) seeks **low field**

Blob averaged  $dB_z/dr$  controls motion of magnetised plasma  
blobs: Anti-potential leads to *magnetic phase separation*

# Paramagnetic plasma: L-mode



Motion of pressure blobs depends on  $dB_z/dr$

$$mn_V \frac{d\vec{v}_r}{dt} \simeq \tilde{M}_\zeta \nabla_r \bar{B}_{\zeta 0}$$

paramagnetic cold blobs move inward,

diamagnetic hot blobs move outward

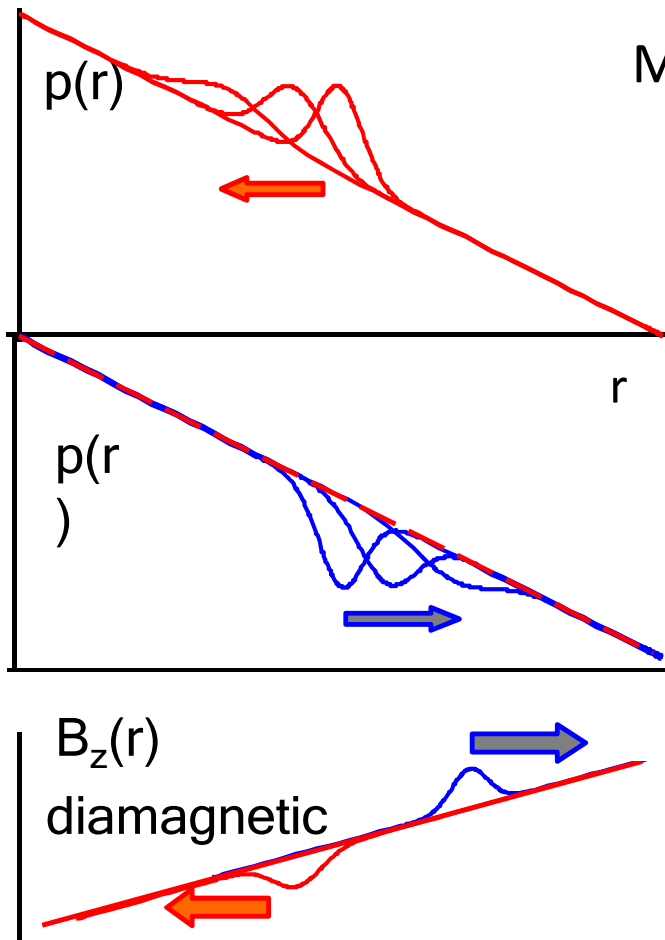
outward thermal energy convection at the expense of

inward magnetic energy convection

p blobs “grow”, “instability”

**L-mode**

# Diamagnetic plasma: H-mode



Motion of pressure blobs depends on  $dB_z/dr$

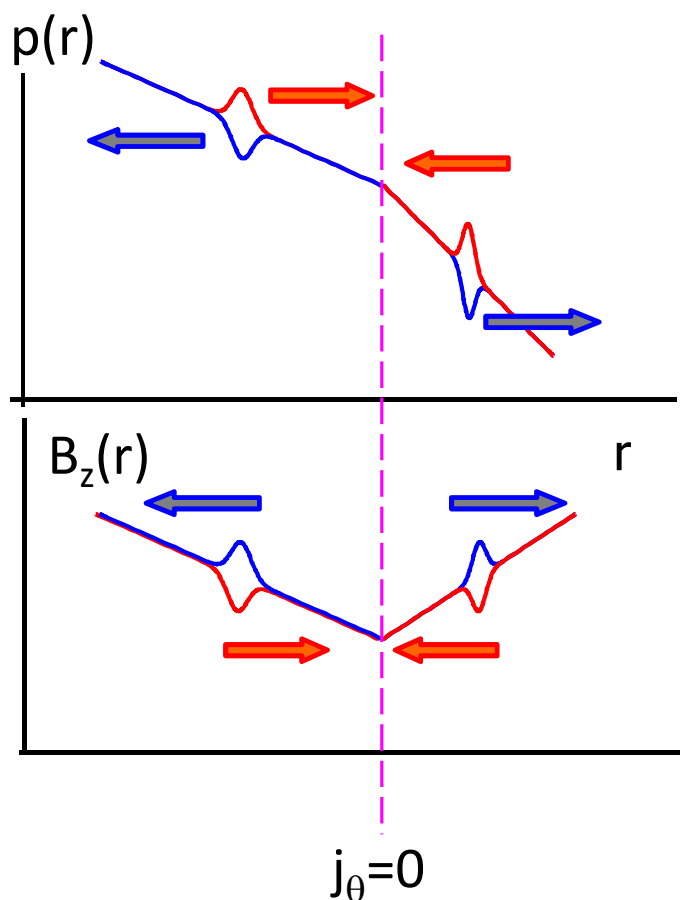
$$mn_{\nu} \frac{d\vec{v}_r}{dt} \simeq \tilde{M}_{\zeta} \nabla_r \bar{B}_{\zeta 0}$$

diamagnetic hot blobs move inward,  
 paramagnetic cold blobs move outward  
 inward thermal energy convection  
 at the expense of  
 outward magnetic energy convection

p blobs “decrease”, “saturation”

**H-mode**

## Magnetic Boundary: phase transition



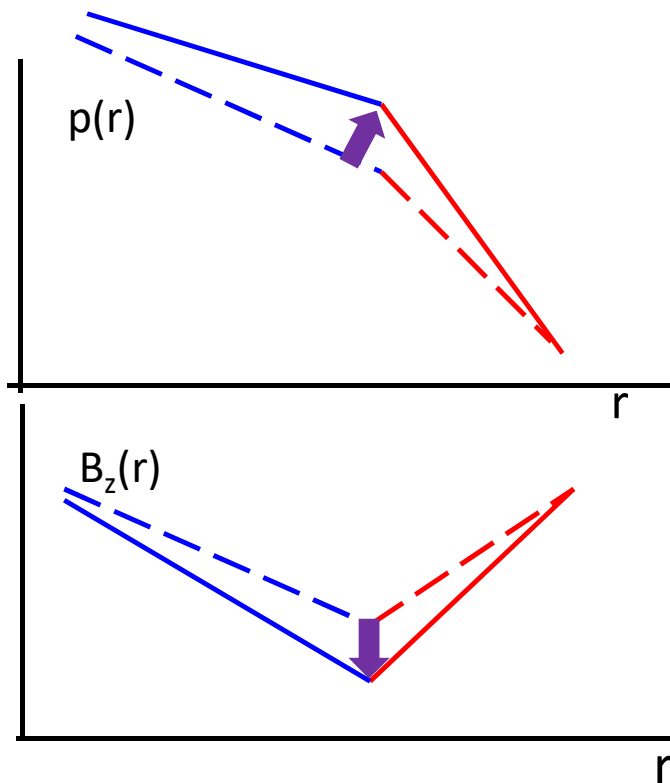
At a magnetic phase boundary blobs of the same type accumulate/separate

diamagnetic blobs (heat) seek magnetic wells  
paramagnetic blobs seek magnetic hills

With multiple blobs moving,  
p and  $B_z$  profiles evolve,  
steepening magnetic hills, digging magnetic wells  
Developing pressure pedestal



## Magnetic Boundary: phase transition



$\nabla p$  increases somewhere, creating diamagnetic region at plasma edge.

Magnetization, of both signs, increases.

Phase transition is self-reinforcing.

Pressure pedestal forms, grows.

## Pedestal formation at magnetisation boundary

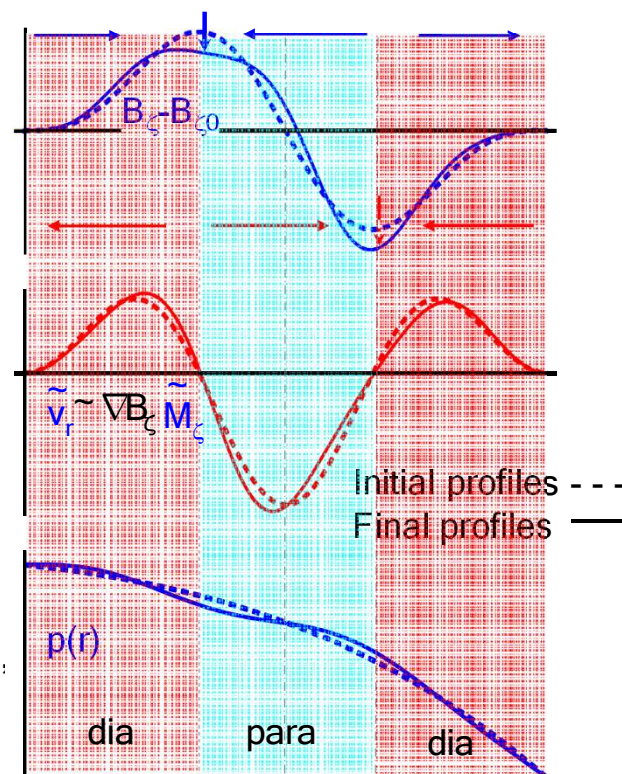
Assume dashed  $B_z(r)$ ,  $p(r)$  initial profiles

Ideal MHD with magnetization force

$$\bar{n}_V m_i \left. \frac{d^2 \xi_r}{dt^2} \right|_M = \tilde{M}_\zeta \nabla \bar{B}_{0z}$$

$$\left. \frac{\partial B_z}{\partial t} \right|_M = \nabla \times (\tilde{v}_r \bar{B}_{0z})$$

$$\left. \frac{3}{2} \frac{\partial p}{\partial t} \right|_M = -\nabla(\tilde{p} \tilde{v})$$



Integrating one temporal step

pressure steepens in **diamagnetic** regions, increases **diamagnetism**

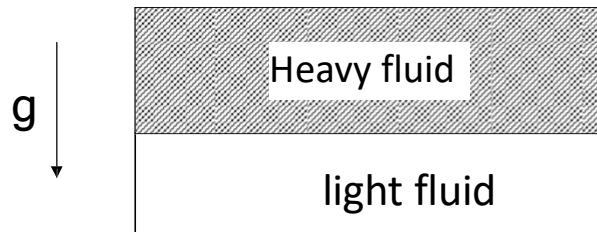
flattens in **paramagnetic** regions, increases **paramagnetism**

Magnetic phase separation drives pedestal formation

## Interchange instability<sup>1</sup>

- present when radial force acts equally on electrons and ions
- equivalent to the Rayleigh-Taylor instability in a fluid.
- magnetization gradient acting on magnetized plasma blobs replace “gravitational field” or “curvature”.

### Magnetization interchange



$$\begin{aligned}\gamma &= \sqrt{g\lambda_{\perp}} = \\ &= \sqrt{-\frac{1}{m_V} \frac{\tilde{p}}{2B^2} \frac{\partial B_{\zeta}^2}{\partial r} \lambda_{\rho}}\end{aligned}$$

Magnetization interchange growth faster for  
high magnetisation, blob amplitude & radius, low field & mass

<sup>1</sup>M.N. Rosenbluth and C.L. Longmire, Annals of Physics, Volume 1, Issue 2, May 1957,120

## Suydam criterion for interchange instability

B. R. Suydam, Proc. 2nd UN Conf. on Peaceful Uses of Atomic Energy, Geneva, 1958.

$$\beta' \left( \frac{Rq}{r_s} \right)^2 \left[ \frac{B^2 \kappa_r}{\mu_0} \right] > \frac{q'^2}{4q^2} \quad \text{magnetic shear opposes interchange of tubes driven by cylindrical curvature and } \nabla\beta$$

Generalization: add magnetization force to cylindrical curvature

$$\beta' \left( \frac{Rq}{r_s} \right)^2 \left[ \frac{B^2 \kappa_r}{\mu_0} + \tilde{M}_z \frac{dB_{0z}}{dr} \right] > \frac{q'^2}{4q^2}$$

In magnetically mixed states  $\tilde{M}_z \frac{dB_{0z}}{dr} < 0$

magnetisation force adds to curvature, instability,  
until the magnetic shear  $q'$  or the sign of  $dB_z/dr$  changes.