Orbit-averaged approach to fast-ion transport in stellarators

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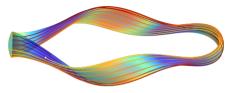
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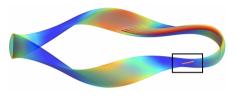
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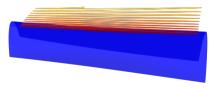
Orbit-averaged approach to fast-ion transport in stellarators

- In a fusion reactor, alpha particles must be confined long enough.
- Passing particles are well confined.
- In non-optimized stellarators, trapped orbits are not confined: large neoclassical transport of thermal particles. Worse for alpha particles...
- ... as they do not not enjoy the confining effect of the E × B drift.
- Good fast-ion confinement is a demanding criterion in stellarator optimization.
- The understanding of fast-ion transport and the development of efficient codes are very important for the design of stellarator reactors.

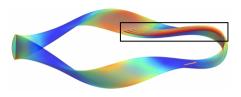


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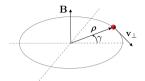


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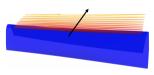




- Usually, Monte Carlo codes that solve a full-orbit kinetic equation or a drift-kinetic equation (DKE) for guiding centers are employed.
 - ASCOT [Hirvijoki, CPC 2014], ANTS [Drevlak, NF 2014], BEAMS3D [McMillan, PPCF 2014], GNET [Masaoka, NF 2013], SIMPLE [Albert, JPP 2020]...



 Guiding centers move rapidly along B and drift across the magnetic field. For many applications, averaging over the rapid motion along B (i.e. over lowest-order orbits) should work.



This talk

- Derivation of an **orbit-averaged DKE** for stellarators*.
 - Reduced phase-space dimensionality.
- Implementation in a Monte Carlo code, KNOSOS-MC.

*Previous work for tokamaks in [Eriksson, PoP 1994], [Falessi,

PoP 2019], [Meng, arXiv 2024] and for model stellarator fields in [Kolesnichenko, PoP 2006].

Orderings and assumptions

Plasma consisting of bulk ions with mass m_i and charge $Z_i e$, electrons with mass m_e , and fast ions with mass m_h , charge $Z_h e$ and characteristic speed v_h .

$$Z_i \sim Z_h \sim 1, \ m_i \sim m_h, \ v_{ti} \ll v_h \ll v_{te}.$$

- Strongly magnetized fast ions: ρ_{h*} = ρ_h/L₀ ≪ 1, where ρ_h is the fast-ion gyroradius and L₀ ~ R ~ a is a characteristic length of the order of the device size.
- Small fast-ion density n_h : the electrostatic potential φ is determined by bulk species and fast-ion self-collisions are negligible.
- $\varphi \simeq \varphi_0$, where φ_0 is a flux function.
- $\rho_{h*} \sim \nu_{h*}$, where ν_{h*} is the fast-ion collisionality.

n		R	a	В	T_i	T_e	$\frac{1}{2}m_h v_h^2$	v_{ti}/v_h	v_h/v_{te}	$ ho_h/a$
	W7-X	5.5	0.5	2.6	1.5	3	60	0.158	0.104	0.027
	$\mathrm{HSR4/18}$	18	2	5	15	15	3500	0.083	0.178	0.024

Typical values of NBI hydrogen ions in W7-X and alpha particles in a Helias reactor HSR4/18 [Beidler, NF 2001].

Full-orbit kinetic equation

• Under the above assumptions, the equation for the fast-ion distribution $f_h(\mathbf{x}, \mathbf{v}, t)$ is

$$\partial_t f_h + \mathbf{v} \cdot \nabla f_h + \frac{Z_h e}{m_h} \left(\mathbf{v} \times \mathbf{B} - \nabla \varphi_0 \right) \cdot \nabla_v f_h = C_h[f_h] + s_h$$

where s_h is a source term and the collision term reads [Helander, CUP 2002]

$$C_h[f_h] = \frac{1}{2\tau_s} v_b^3 \nabla_v \cdot \left(\nabla_v \nabla_v v \cdot \nabla_v f_h \right) + \frac{1}{\tau_s} v_c^3 \nabla_v \cdot \left(\frac{\mathbf{v}}{v^3} f_h \right) + \frac{1}{\tau_s} \nabla_v \cdot \left(\mathbf{v} f_h \right).$$

- Here, τ_s is the slowing-down time, and v_c and v_b are the velocities below which the drag and the pitch-angle scattering of the bulk ions start to matter.
- The E × B drift is negligible in our ordering and certainly for alpha particles, but we keep it to be able to check its influence in current experiments.

Drift-kinetic equation

- Expanding the full-orbit kinetic equation in ρ_{h*} ≪ 1, one can average out the motion of the fast ions around lines of **B**. The result is the DKE for the guiding centers [Hazeltine, PoF 1973], [d'Herbemont, JPP 2022].
- Velocity coordinates $\{\mathcal{E}, \mu, \sigma, \phi\}$, where $\mathcal{E} = v^2/2 + Z_h e \varphi_0/m_h$, $\mu = v_{\perp}^2/2B$, $\sigma = v_{\parallel}/|v_{\parallel}|$ and ϕ is the gyrophase. Here,

$$egin{aligned} & \mathcal{V}_{||}(\mathbf{x},\mathcal{E},\mu,\sigma) = \sigma \sqrt{2\left(\mathcal{E}-U(\mathbf{x},\mu)
ight)}\,, \quad \mathbf{v}(\mathbf{x},\mathcal{E}) = \sqrt{2\left(\mathcal{E}-rac{Z_h e arphi_0(\mathbf{x})}{m_h}
ight)}\,, \ & U(\mathbf{x},\mu) := \mu B(\mathbf{x}) + rac{Z_h e arphi_0(\mathbf{x})}{m_h}. \end{aligned}$$

Drift-kinetic equation

- One can show that $f_h \simeq F_h$, where F_h is the gyroaverage of f_h .
- The equation for F_h is

$$\partial_t F_h + \dot{\mathbf{x}} \cdot \nabla F_h = C_h[F_h] + S_h,$$

where S_h is the gyroaverage of the source term and the collision term reads

$$C_{h}[F_{h}] = \nu_{hi}^{D} \frac{v_{||}}{B} \partial_{\mu} \left(\mu v_{||} \partial_{\mu} F_{h} \right) + \frac{v_{||}}{\tau_{s}} \left[\partial_{\mathcal{E}} \left(\frac{v^{2}}{v_{||}} \left(1 + \frac{v_{c}^{3}}{v^{3}} \right) F_{h} \right) + 2 \left(1 + \frac{v_{c}^{3}}{v^{3}} \right) \partial_{\mu} \left(\frac{\mu}{v_{||}} F_{h} \right) \right].$$

• As for the guiding-center* trajectories, $\dot{\mathbf{x}} = v_{||}\hat{\mathbf{b}} + \mathbf{v}_d$, where $\mathbf{v}_d = \mathbf{v}_M + \mathbf{v}_E$ and

$$\mathbf{v}_M = rac{1}{\Omega_h} \hat{\mathbf{b}} imes (\mathbf{v}_{||}^2 \hat{\mathbf{b}} \cdot
abla \hat{\mathbf{b}} + \mu
abla B), \quad \mathbf{v}_E = rac{1}{B} \hat{\mathbf{b}} imes
abla arphi_0.$$

 $|\mathbf{v}_d|/|\mathbf{v}_{||}| \sim \rho_{h*} \ll 1.$

*In what follows, we often refer to guiding-center trajectories as particle trajectories.

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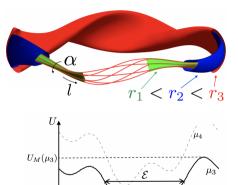
Orbit-averaged approach to fast-ion transport in stellarators

Orbit-averaged DKE: coordinates and lowest-order orbits

- Coordinates $\{r, \alpha, I\}$.
- Expand the DKE in $\rho_{h*} \ll 1$.
- $F_h = F_h^{(0)} + F_h^{(1)} + \dots$ To lowest order, orbits follow magnetic field lines and

$$\mathbf{v}_{||}\hat{\mathbf{b}}\cdot
abla F_h^{(0)} = 0.$$

- $U := \mu B + Z_h e \varphi_0 / m_h$ and let $U_M(\mu)$ be the maximum of U on the flux surface for fixed μ . If $\mathcal{E} < U_M(\mu)$, trapped. If $\mathcal{E} > U_M(\mu)$, passing.
- For trapped particles, $F_h^{(0)} \equiv F_h^{(0)}(r, \alpha, \mathcal{E}, \mu, t)$. For passing particles, $F_h^{(0)} \equiv F_h^{(0)}(r, \mathcal{E}, \mu, \sigma, t)$.
- $F_h^{(0)}$ obtained averaging next-order terms of the DKE.



 I_{h}

 $U_m(\mu_3)$

H2

Orbit-averaged DKE for trapped fast ions

• The equation that determines $F_h^{(0)}(r, \alpha, \mathcal{E}, \mu, t)$ for trapped particles is

$$\partial_t F_h^{(0)} + \overline{\mathbf{v}_d \cdot \nabla r} \, \partial_r F_h^{(0)} + \overline{\mathbf{v}_d \cdot \nabla \alpha} \, \partial_\alpha F_h^{(0)} = \overline{C_h[F_h^{(0)}]} + \overline{S_h} \,,$$

where $\overline{(\cdot)} = \tau_b^{-1} \sum_{\sigma} \int_{l_{b_1}}^{l_{b_2}} |v_{||}|^{-1}(\cdot) dI$ and $\tau_b = 2 \int_{l_{b_1}}^{l_{b_2}} |v_{||}|^{-1} dI$ is the orbit time. $J(r, \alpha, \mathcal{E}, \mu) = 2 \int_{l_{b_1}}^{l_{b_2}} |v_{||}| dI$ is called *second adiabatic invariant*.

• Relation between the average of \mathbf{v}_d and J:

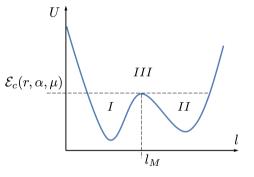
$$\overline{\mathbf{v}_d \cdot \nabla r} = \frac{m_h}{Z_h e \Psi_t' \tau_b} \partial_\alpha J, \quad \overline{\mathbf{v}_d \cdot \nabla \alpha} = -\frac{m_h}{Z_h e \Psi_t' \tau_b} \partial_r J,$$

where Ψ'_t is the derivative with respect to *r* of the toroidal flux.

In the absence of collisions, trapped particles move along curves of constant J.

Orbit-averaged DKE for trapped fast ions: junctures connecting wells

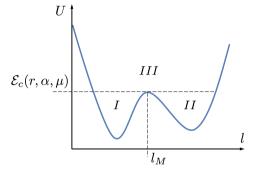
- The invariance of J can break at junctures, where particles undergo transitions between different types of wells.
- These collisionless transitions, where the value of J changes abruptly, are the cause of fast-ion stochastic transport [Beidler, PoP 2001], [Kolesnichenko, PoP 2022].





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- These collisionless transitions, where the value of J changes abruptly, are the cause of fast-ion stochastic transport [Beidler, PoP 2001], [Kolesnichenko, PoP 2022].
- For exactly zero collision frequency, F_h⁽⁰⁾
 can be discontinuous at junctures.

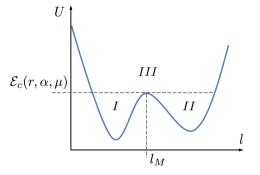


Apply techniques from [d'Herbemont, JPP 2022] to derive the discontinuity condition by imposing conservation of the collisionless particle flux:

$$F_{h,I}^{(0)}\left(\partial_{\alpha}J_{I}\partial_{r}\mathcal{E}_{c}-\partial_{r}J_{I}\partial_{\alpha}\mathcal{E}_{c}\right)+F_{h,II}^{(0)}\left(\partial_{\alpha}J_{II}\partial_{r}\mathcal{E}_{c}-\partial_{r}J_{II}\partial_{\alpha}\mathcal{E}_{c}\right)=F_{h,II}^{(0)}\left(\partial_{\alpha}J_{II}\partial_{r}\mathcal{E}_{c}-\partial_{r}J_{II}\partial_{\alpha}\mathcal{E}_{c}\right).$$

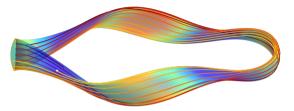
Orbit-averaged DKE for trapped fast ions: junctures connecting wells

- For finite collision frequency, $F_h^{(0)}$ is continuous, but $\partial_\mu F_h^{(0)}$ is not.
- The relation between the values of \(\partial_\mu}\mathcal{F}_h^{(0)}\) on each side of the juncture is obtained from conservation of the collisional particle flux:



$$\left(\int_{l_{b_1}}^{l_{b_2}} \frac{|\mathbf{v}_{||}|}{B} \mathrm{d}I\right)_I \partial_\mu F_{h,I} + \left(\int_{l_{b_1}}^{l_{b_2}} \frac{|\mathbf{v}_{||}|}{B} \mathrm{d}I\right)_{II} \partial_\mu F_{h,II} = \left(\int_{l_{b_1}}^{l_{b_2}} \frac{|\mathbf{v}_{||}|}{B} \mathrm{d}I\right)_{III} \partial_\mu F_{h,III}$$

Orbit-averaged DKE for passing fast ions

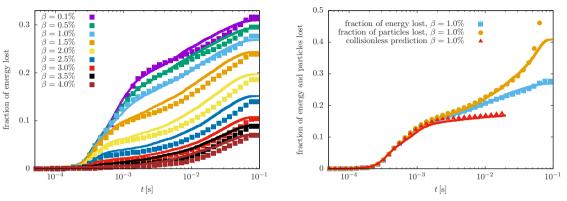


• The equation that determines $F_h^{(0)}(r, \mathcal{E}, \mu, \sigma, t)$ for passing fast ions is

$$\partial_t F_h^{(0)} = \left\langle \frac{B}{v_{||}} \right\rangle_r^{-1} \left\langle \frac{B}{v_{||}} C_h[F_h^{(0)}] \right\rangle_r + \left\langle \frac{B}{v_{||}} \right\rangle_r^{-1} \left\langle \frac{B}{v_{||}} S_h \right\rangle_r$$

Here, $\langle \cdot \rangle_r$ denotes flux-surface average and we have used that, for passing particles, $\partial_{\alpha} F_h^{(0)} \equiv 0$ and $\langle \mathbf{v}_d \cdot \nabla r \rangle_r \equiv 0$.

Implementation of the orbit-averaged DKE in a code: KNOSOS-MC



- Alpha particle transport in a Helias reactor configuration, HSR4/18. Alpha particles born at mid-radius.
- KNOSOS-MC (markers) vs guiding-center simulations with ASCOT (solid curves).
- In these simulations, KNOSOS-MC is one order of magnitude faster than ASCOT.

Conclusions and outlook

- Orbit-averaged drift-kinetic equation for fast-ion transport in general stellarator geometry derived.
 - ▶ Radially global, includes collisions, and accounts for both trapped and passing particles.
 - Careful treatment of junctures between different types of wells.
- Equation implemented in a new Monte Carlo code, KNOSOS-MC.
- Comparisons between KNOSOS-MC and guiding-center calculations with ASCOT support the validity of the orbit-averaged approach.
- KNOSOS-MC seems to be sufficiently fast to include direct simulations of fast-ion transport in stellarator optimization codes.

Possible routes for future work

- Improve numerical methods in KNOSOS-MC and carry out a more complete benchmark.
- Integrate KNOSOS-MC into stellarator optimization codes.
- Finite-difference code that directly calculates the steady state of the new equation.