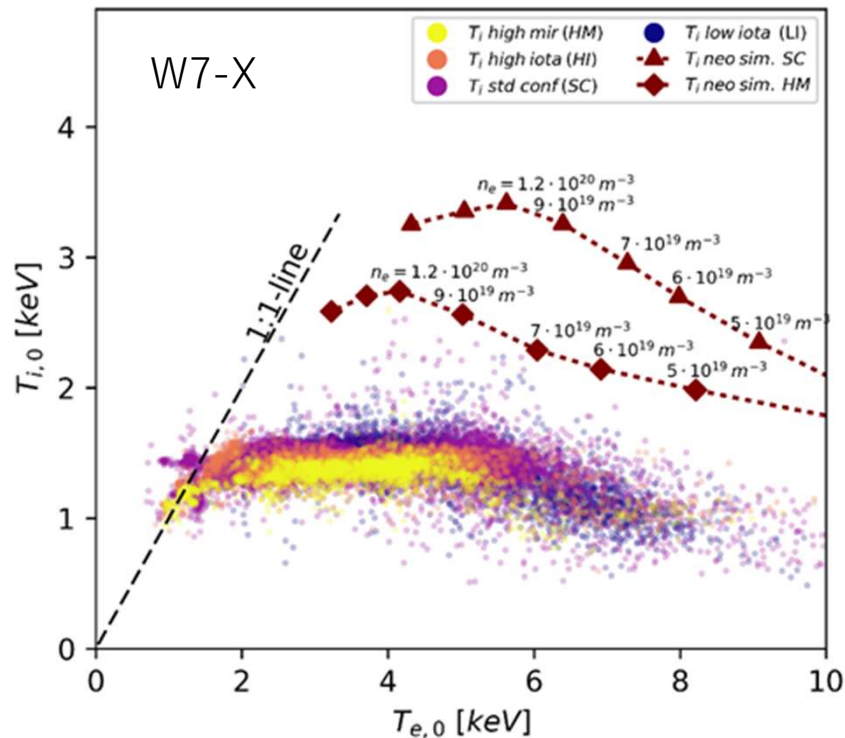


Ion transport in electron heated plasmas

Marc Beurskens, Kenji Tanaka, and Felix Warmer

Report Experiment 15 January

Ion temperature clamping generally observed in W7-X with ECRH



In ECRH plasmas with gas fuelling, generally

- Ion temperature is clamped at $T_i \sim 1.5 \pm 0.2$ keV
- It is well below neoclassical predictions
- It is virtually configuration independent

As well as lots of evidence for turbulent transport:

- Density profiles are flat to slightly peaked (not hollow)
- Impurity transport is turbulent
- Turbulent electron heat transport is diffusive

Key question to this paper:

→ what is causing T_i to be clamped at 1.5 keV?

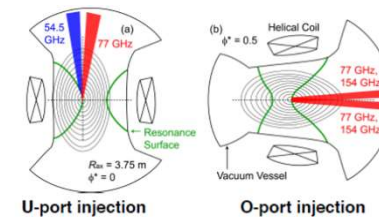
Electron heating in W7-X and LHD

- **W7-X :**

- ECRH heating is the main heating source with $P_{\text{ECRH}} < 7\text{MW}$ and with steady state capability (140GHz)

- **LHD:**

- ECRH heating has breakdown as main purpose. Available sources $P_{\text{ECRH}} \sim 3\text{ MW for } 2\text{s}$
 - U port One 77GHz 1MW for 2sec 0.8MW for 3sec
 - O port: One 77GHz 1MW for 2sec 0.8MW for 3sec
Two 154GHz 1MW for 2sec 0.8MW for 3sec
one 154/116GHz dual freq 1MW for 1sec



- NNBI: negative ion beams: mainly electron heating: $P_{\text{NNBI}} = 15\text{ MW for H and } 10\text{ MW for D}$

	H [MW]	D [MW]
NNBI-1	5.5	3.3
NNBI-2	5	3
NNBI-3	5	3.3
PNBI-4	5 - 6	9
PNBI-5	6	9

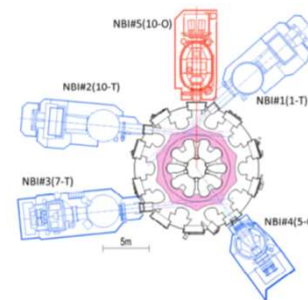


Fig. 1. Top view of LHD and arrangement of NNBI.

Ion transport under electron heating (M. Beurskens (IPP), K. Tanaka, F. Warmer(IPP))

Experimental conditions:

(R_{ax} , Polarity, B_t) = (3.6 m, CW, 2.85 T) H2 gas puff

Shot #: 167117-167185 (59 + 9 shots)

Background and motivations

Increase of Ti under electron heating is essential for the fusion reactor. In W-7X, and AUG, the saturation of increase of Ti is observed. This is called “Ti clamping”. This can be interpreted as the enhancement of the Ti stiffness with higher T_e/T_i . In particular, this study is important in high density with electron heating, where the ion heating is equipartition heating, since such condition is similar to the one in the reactor. We aimed density scan and power scan to obtain the dataset of this study.

Results

We try to use high power of electron heating. Thus, we planned to use two 154GHz balanced tangential ECRH, one 77GHz perpendicular ECRH, NNB and PNB. All systems worked very well without much failure and we could achieve most session goals in 59 pulses as well as conduct 9 addition density modulation experiments.

This time we did not use pellets as repetitive pellet injection was not possible. With gas puffing, CXS7 worked at up to $\sim 5 \times 10^{19} \text{m}^{-3}$. Perhaps data can be improved by averaging of spectra over multiple beam blips?

- In Pure ECRH we have a good density scan dataset to be compared to W7-X at $1-5 \times 10^{19} \text{m}^{-3}$
- At $3 \times 10^{19} \text{m}^{-3}$ we have a very complete power scan from 3-12MW
- At $5 \times 10^{19} \text{m}^{-3}$ and above more useful power scans were conducted

Pre selection of shots for analysis exp 15.01.2021.

- ECRH power scan was not done as Power was limited to $P_{\text{ECRH}} = 2\text{-}2.5\text{MW}$
- Negative NBI (1,2,3) excellent performance 3 sources: $P_{\text{n-NBI}} < 9\text{ MW}$
- Pre-selection for discussion today:

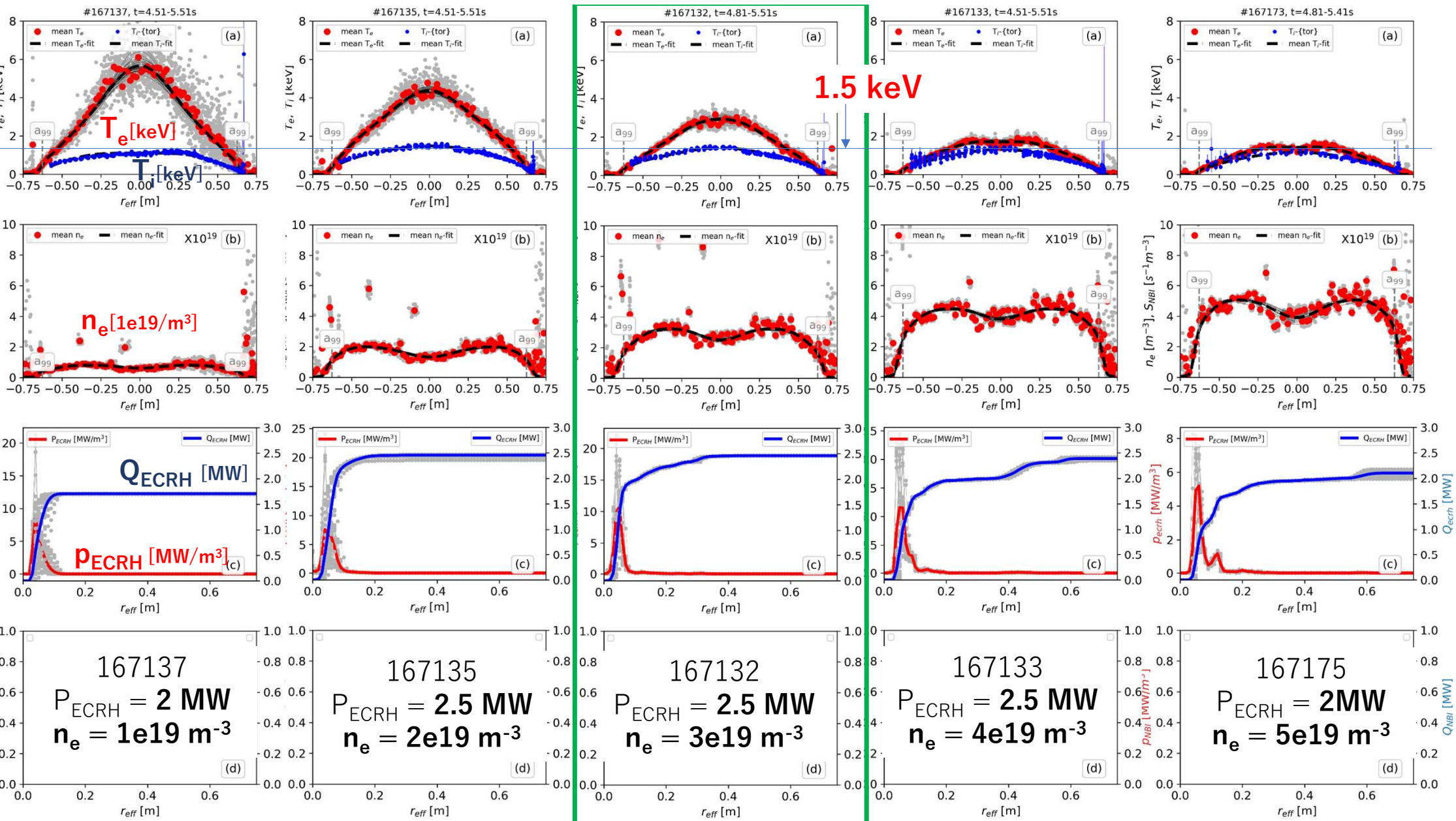
Density scans

- ECRH: $n_{\text{e,max}} = 1\text{-}6\ 10^{19}\ \text{m}^{-3}$, $P_{\text{ECRH}} = 2\text{-}2.5\ \text{MW}$
- nNBI: $n_{\text{e,max}} = 2\text{-}7\ 10^{19}\ \text{m}^{-3}$, $P_{\text{n-NBI}} = 9\ \text{MW}$

Power scans:

- nNBI: $n_{\text{e,max}} = 3\ 10^{19}\ \text{m}^{-3}$, $P_{\text{n-NBI}} = 3\text{-}9\ \text{MW}$
- ECRH+nNBI: $n_{\text{e,max}} = 3\ 10^{19}\ \text{m}^{-3}$, $P_{\text{ECRH}} = 2.5\ \text{MW}$, $P_{\text{n-NBI}} = 0\text{-}6\ \text{MW}$

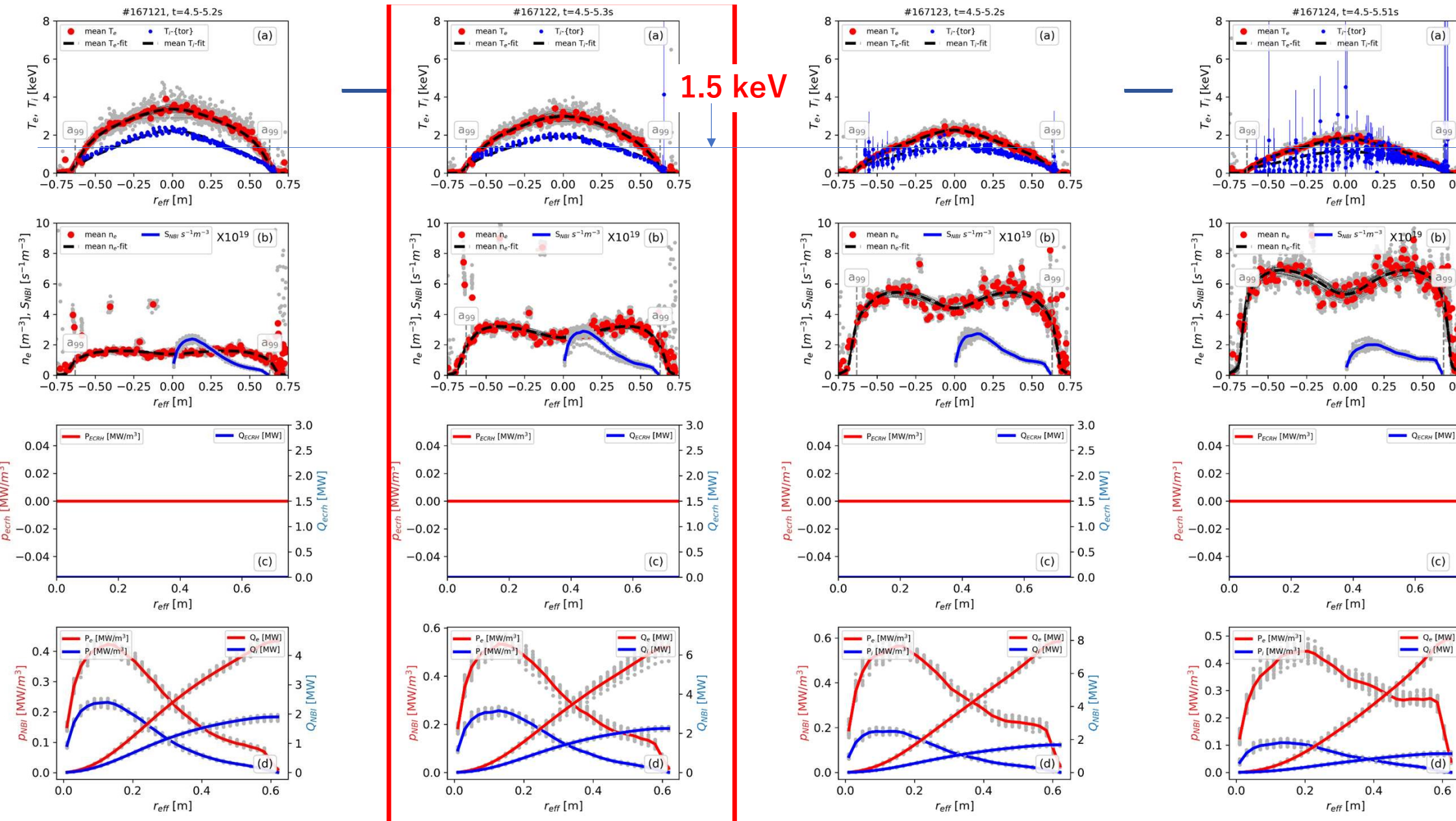
ECRH: Density scan from $n_{e,\max} = 1-6 \cdot 10^{19} \text{ m}^{-3}$, $P_{\text{ECRH}} = 2-2.5 \text{ MW}$



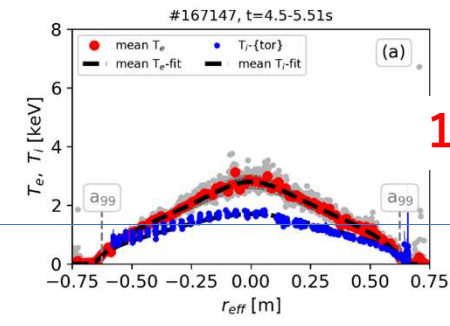
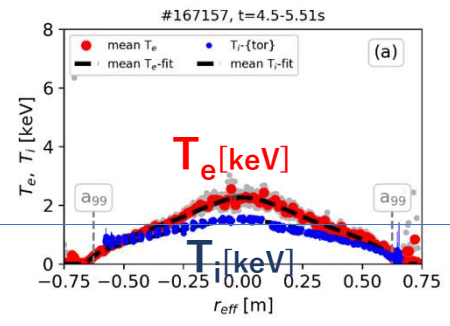
N-NBI: Density scan from $n_{e,\max} = 2-7 \cdot 10^{19} \text{ m}^{-3}$, $P_{\text{n-NBI}} = 9 \text{ MW}$ (3 sources)

$\sim 2/3 P_{\text{n-NBI}} \rightarrow \text{electron}$

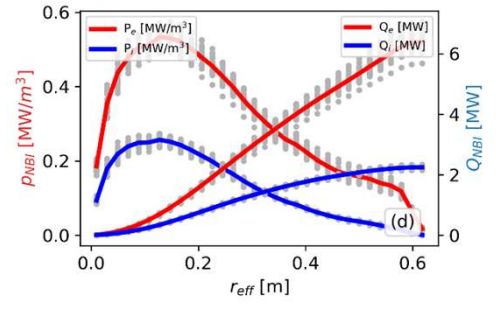
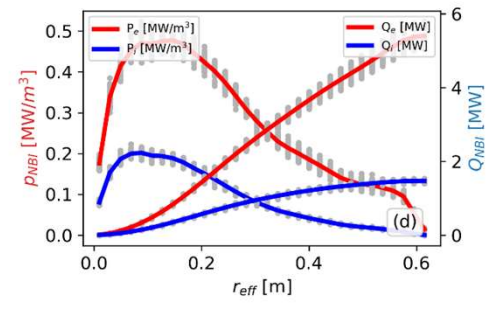
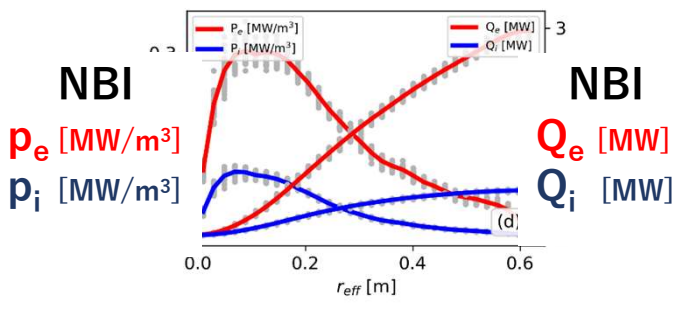
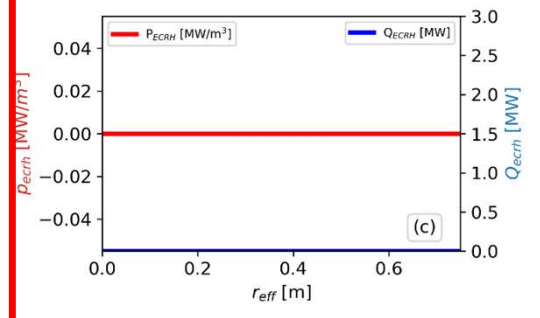
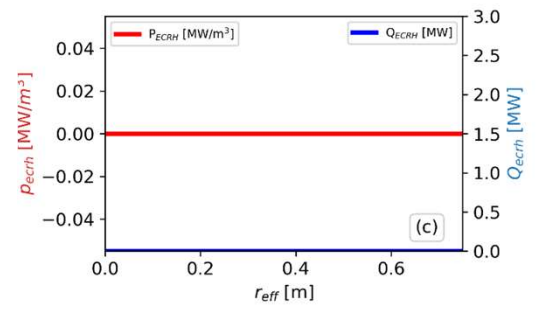
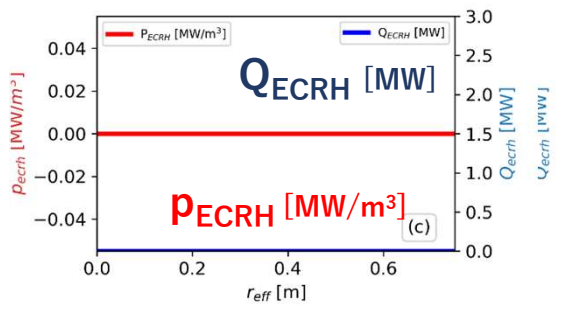
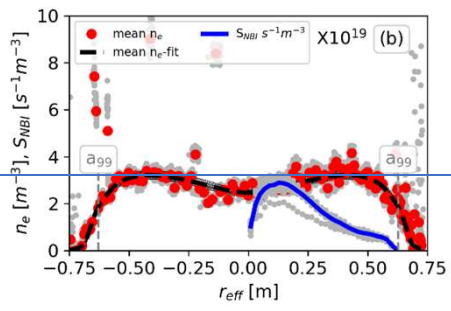
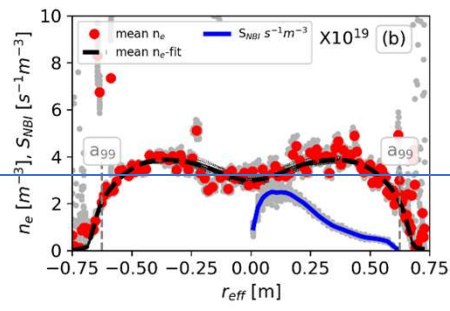
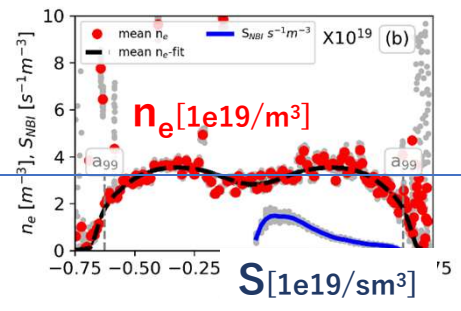
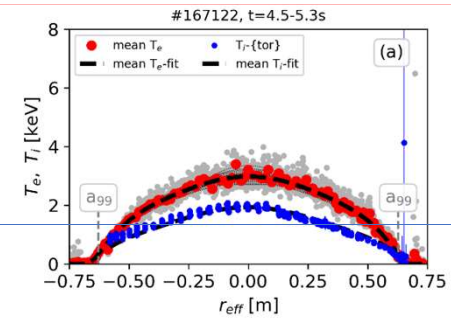
$\sim 1/3 P_{\text{n-NBI}} \rightarrow \text{ions (varies)}$



N-NBI: Power scan at $n_{e,\max} = 3 \cdot 10^{19} \text{ m}^{-3}$, $P_{\text{n-NBI}} = 3\text{-}9 \text{ MW}$ (1-2-3 sources)



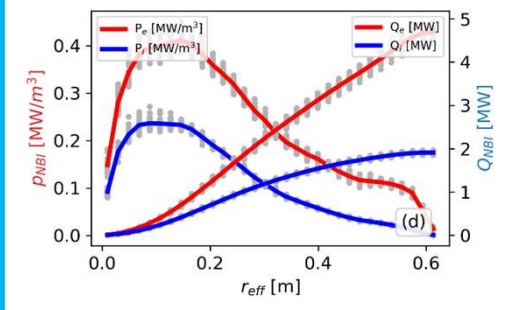
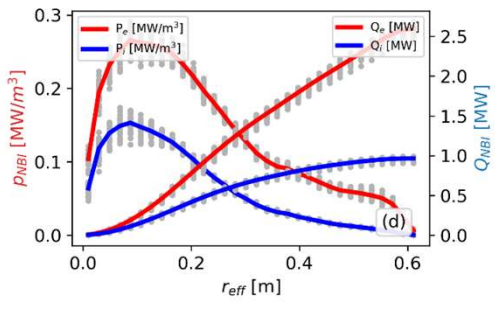
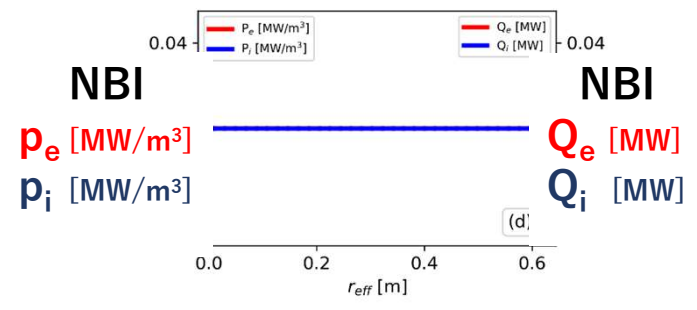
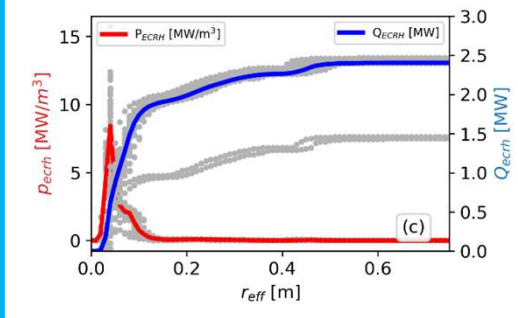
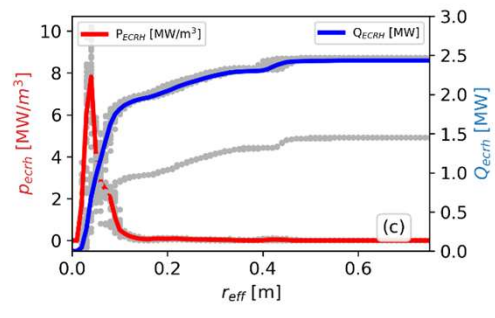
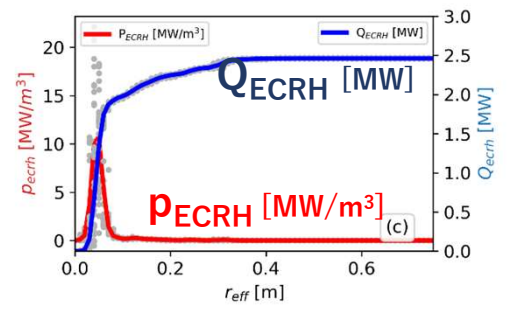
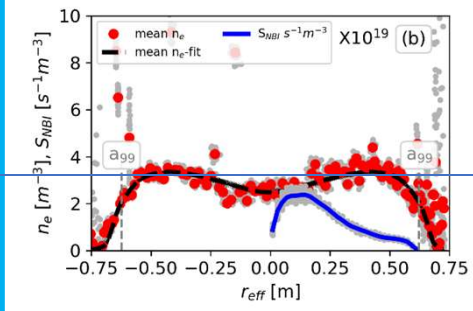
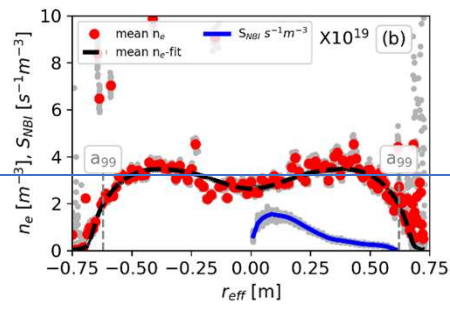
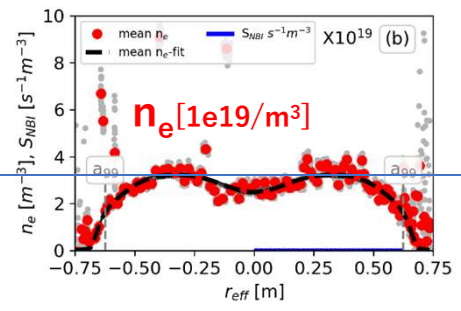
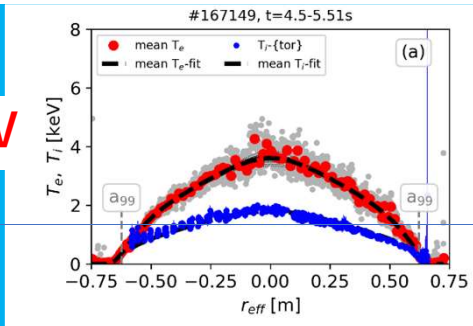
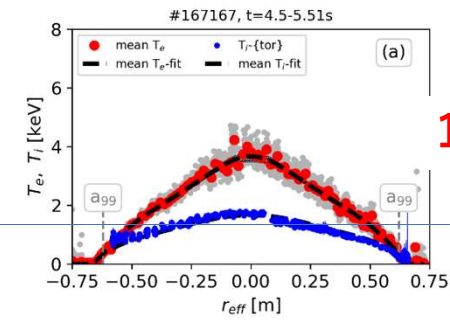
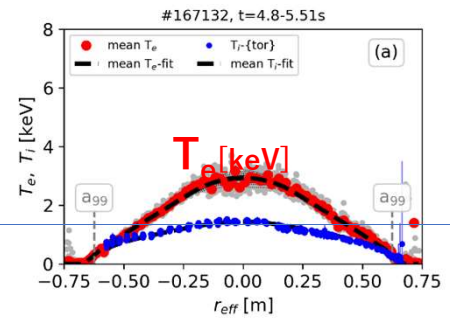
1.5 keV



ECRH+nNBI:

Power scan at $n_{e,\max} = 3 \cdot 10^{19} \text{ m}^{-3}$, $P_{\text{ECRH}} = 2.5 \text{ MW}$,

$P_{\text{n-NBI}} = 0-6 \text{ MW}$ (0-2 sources)

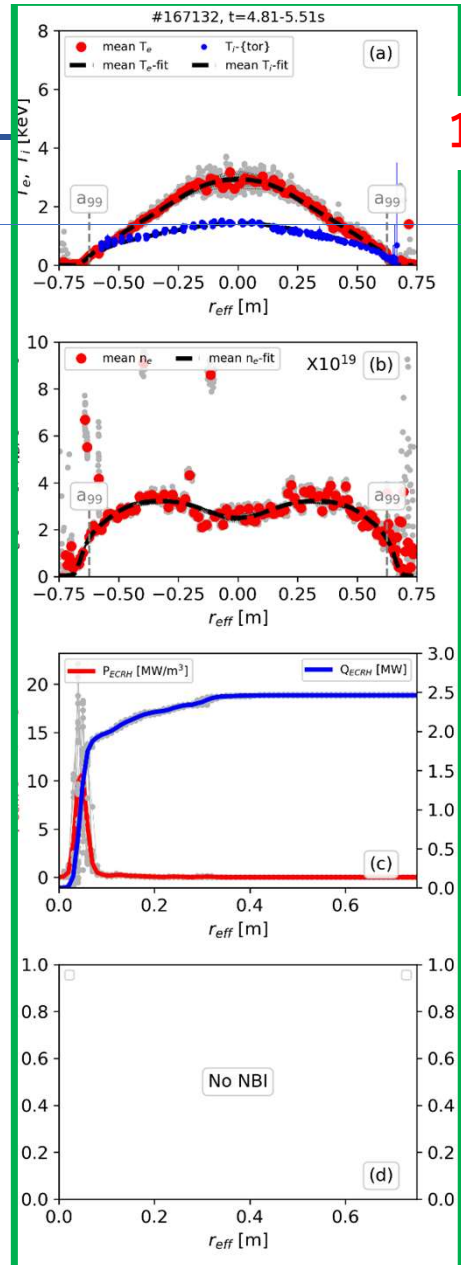
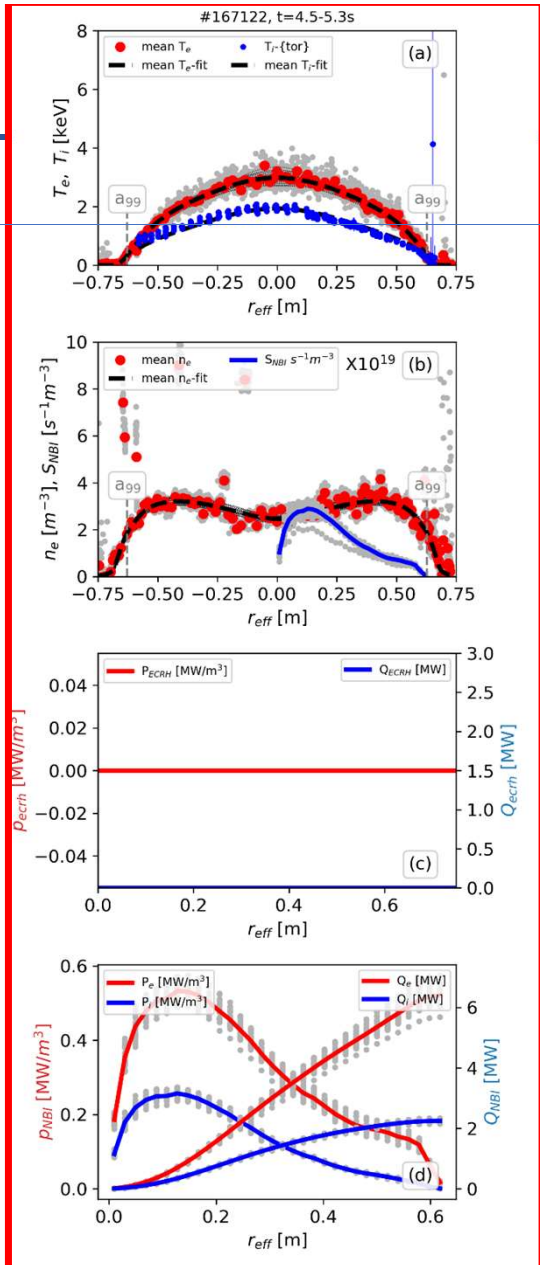


Do we see evidence for ion temperature clamping?

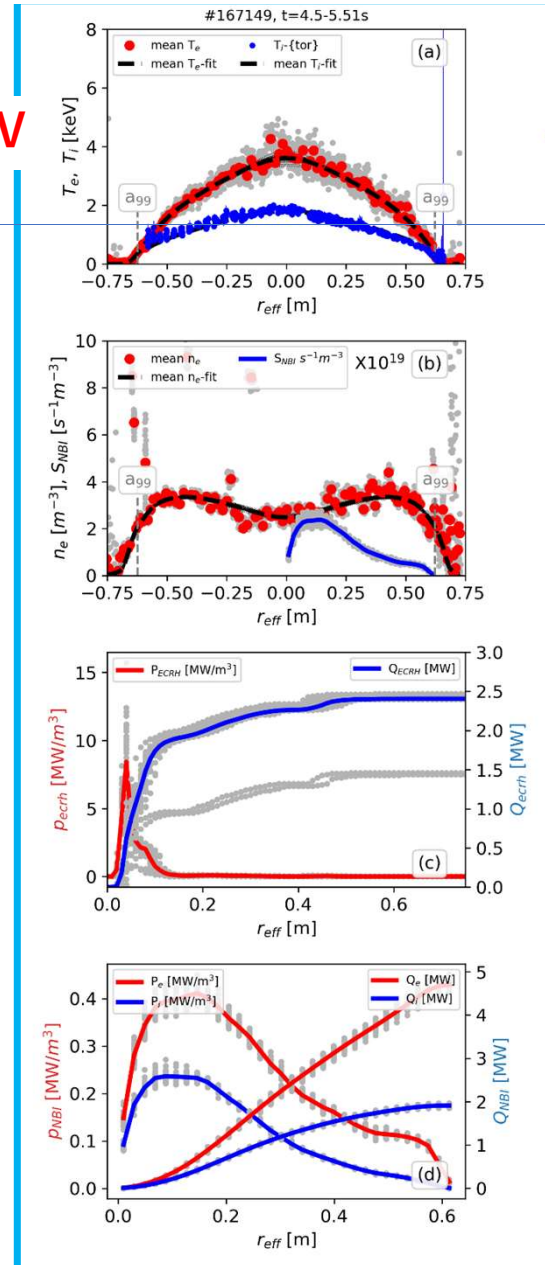
- In ECRH heated plasmas $T_i \sim 1.5\text{keV}$ $T_e > T_i$
- In n-NBI heated plasmas T 1.5 \rightarrow 2 keV $T_e \sim T_i$ or closer at least
(be-it with additional ion heating compoment)

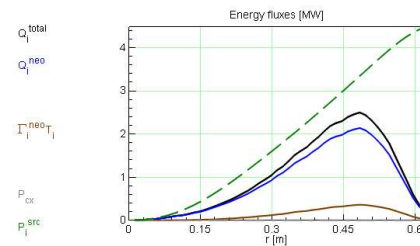
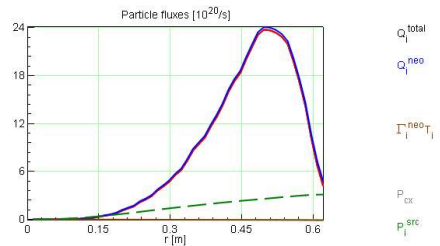
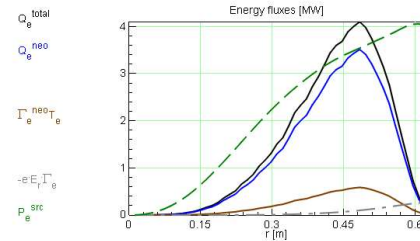
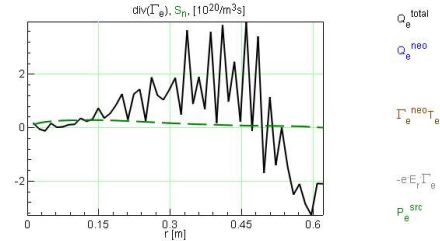
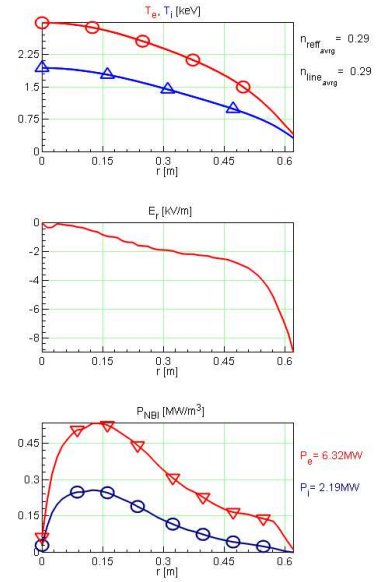
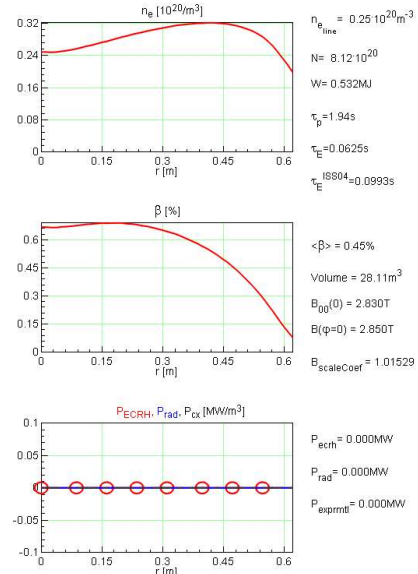
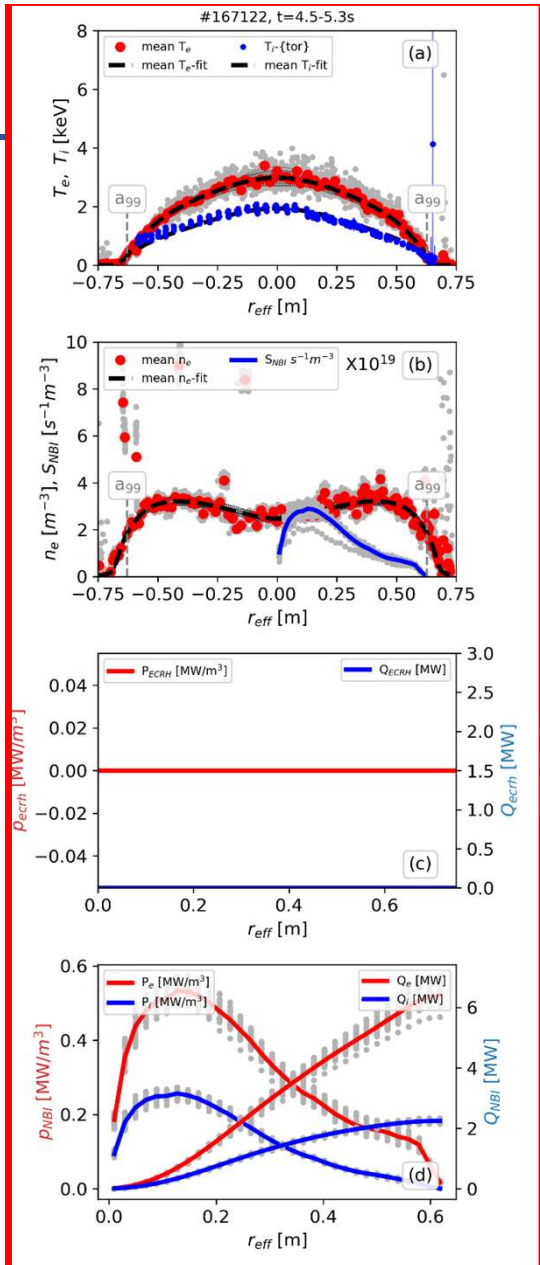
Combined ECRH and N-NBI $T_i > 1.5$ keV

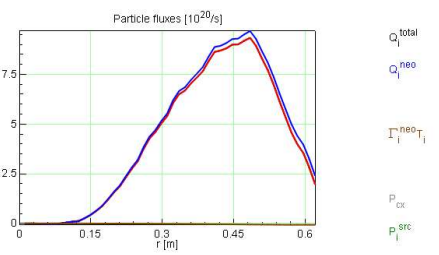
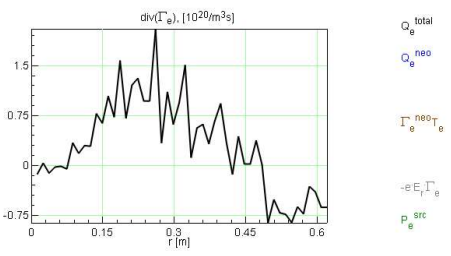
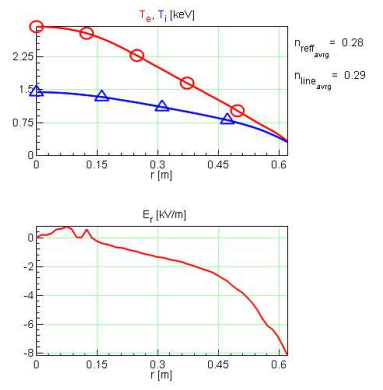
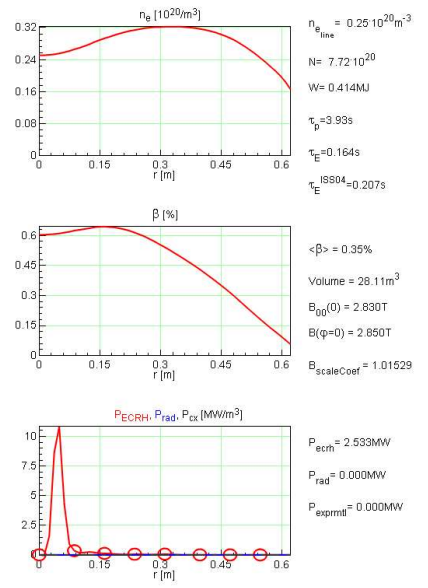
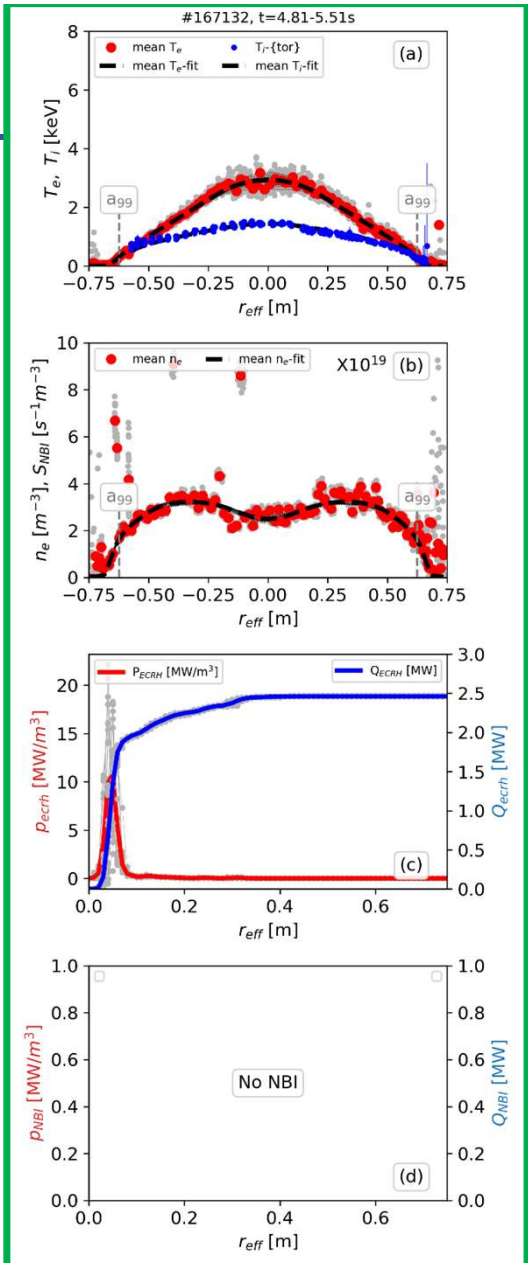
Look at power valance for three selected cases:

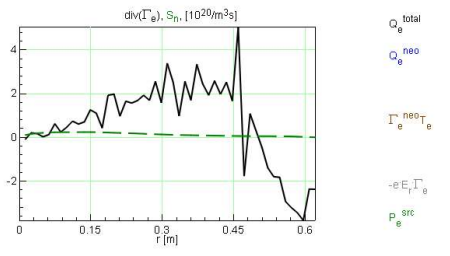
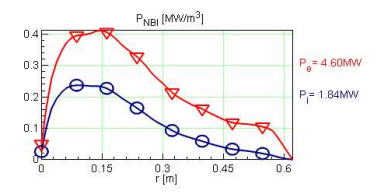
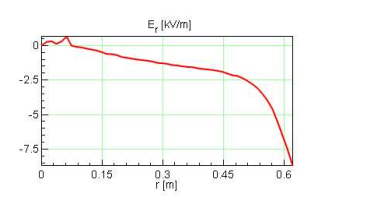
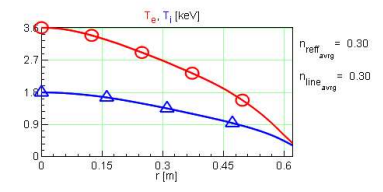
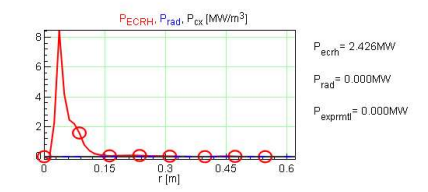
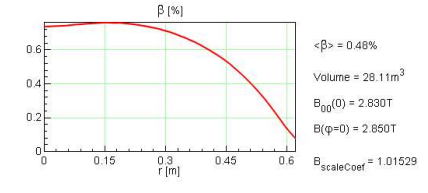
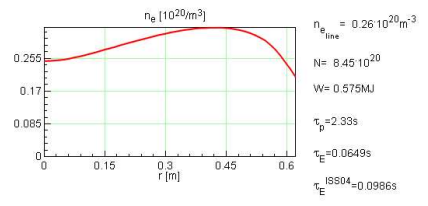
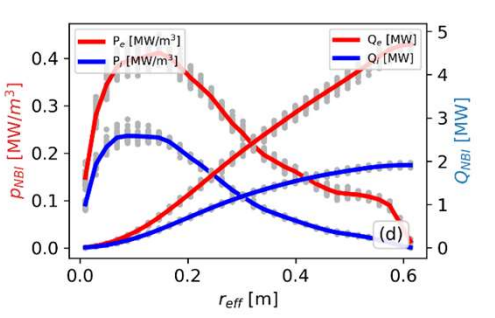
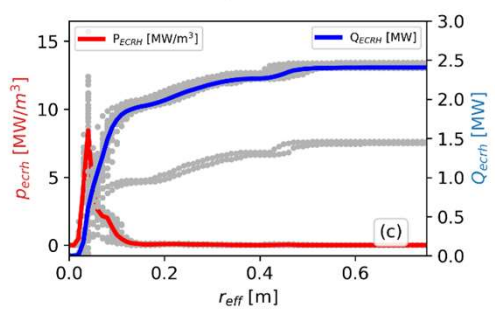
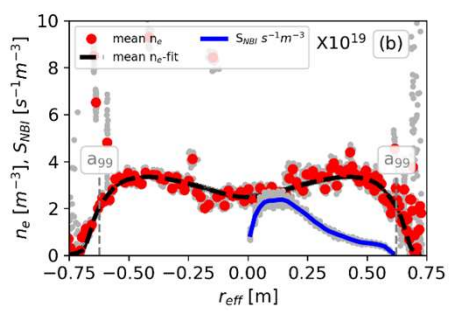
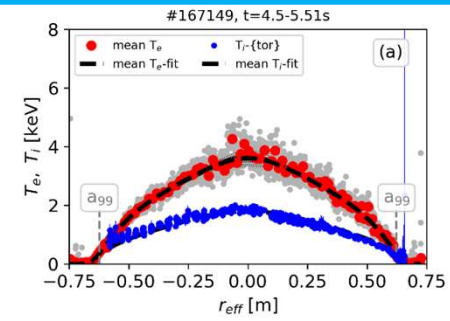


1.5 keV

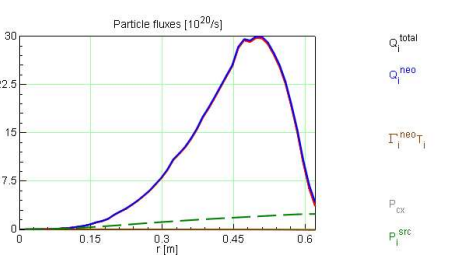
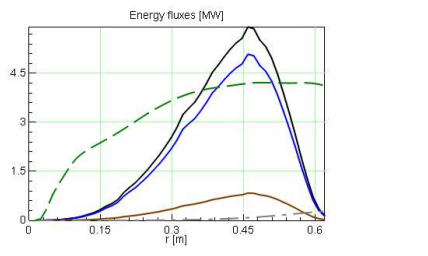




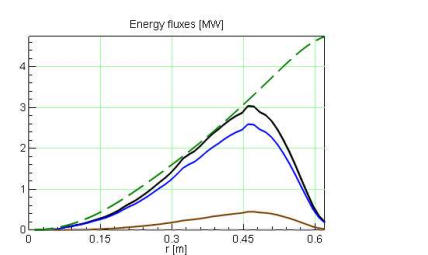




Q_e^{total}
 Q_e^{neo}
 $\Gamma_e^{neo} T_e$
 $-E_r \Gamma_e$
 P_e^{src}



Q_e^{total}
 Q_e^{neo}
 $\Gamma_e^{neo} T_i$
 P_{cx}
 P_i^{src}



Strong contribution by neoclassical transport

- Consistent with the draft PRL of Felix Warmer NC plays a bigger role in LHD than we usually see in W7-X
- Whether ITG plays a role at all in clamping T_i in these plasmas remains to be seen
- Work in progress