



CHALMERS
UNIVERSITY OF TECHNOLOGY

A decorative graphic consisting of five yellow five-pointed stars scattered around the word 'DREAM'. To the right of the word is a large, stylized crescent moon with a rainbow gradient from purple on the outside to yellow in the center.

DREAM

Five flavours of hottail formation

O. Embreus and M. Hoppe

Flat-top plasma (“ASDEX-like”):

- Core temperature of 5.5 keV
- Pure deuterium plasma with uniform $n_e = 2.6 \times 10^{19} \text{ m}^{-3}$
- $R_0 = 1.65 \text{ m}$, $a = 0.5 \text{ m}$
- $I_p = 800 \text{ kA}$, $\left[j_{\parallel} / B = (j_0 / B_{\min})(1 - (r/a)^4)^{3/2} \right]$
- Elongation $\kappa \in [1, 1.15]$

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Disruption induced by instantaneous deposition of deuterium-argon mixture:

- Uniform density $n_{Ar} = 7.8 \times 10^{18} \text{ m}^{-3}$, $n_D = 2.6 \times 10^{20} \text{ m}^{-3}$
- Constant heat diffusion $D = 4000 \text{ m}^2/\text{s}$.

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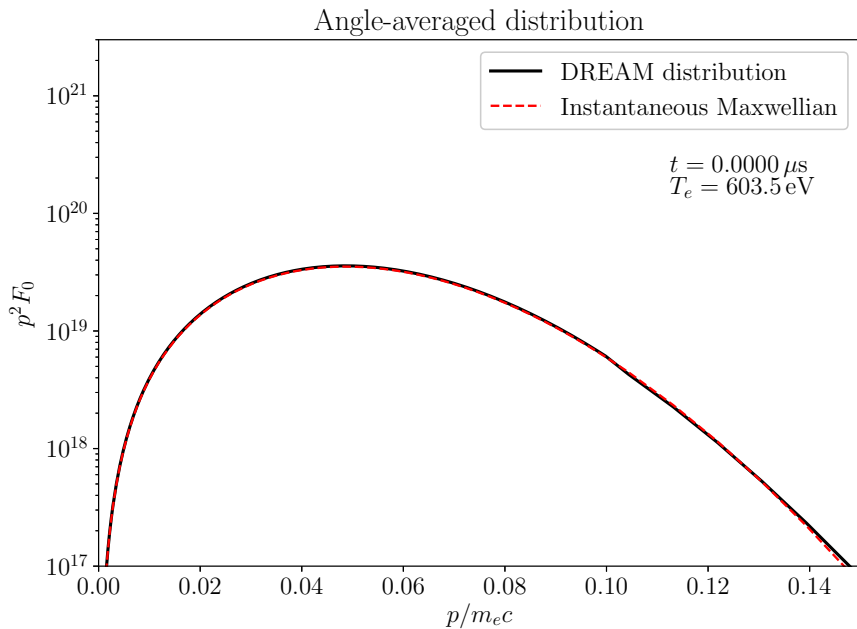
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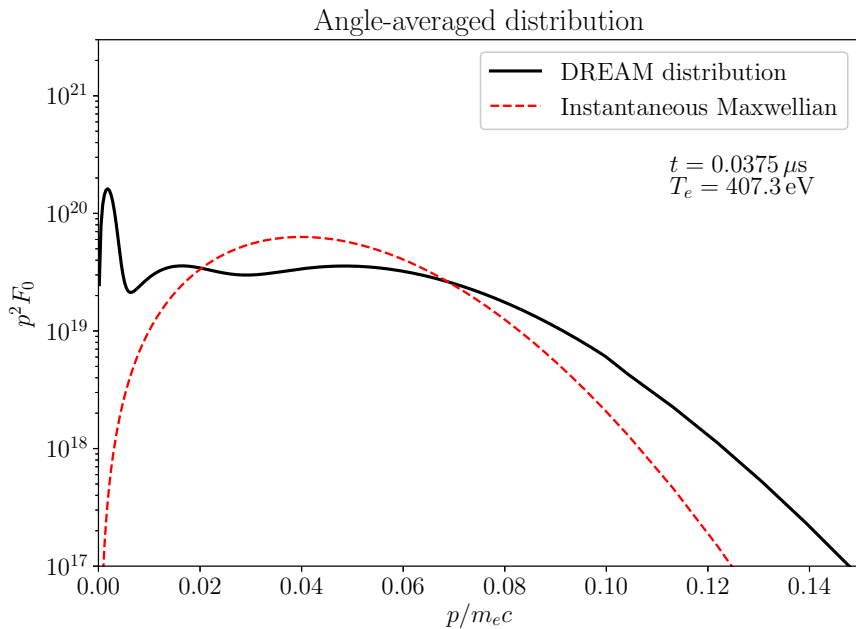
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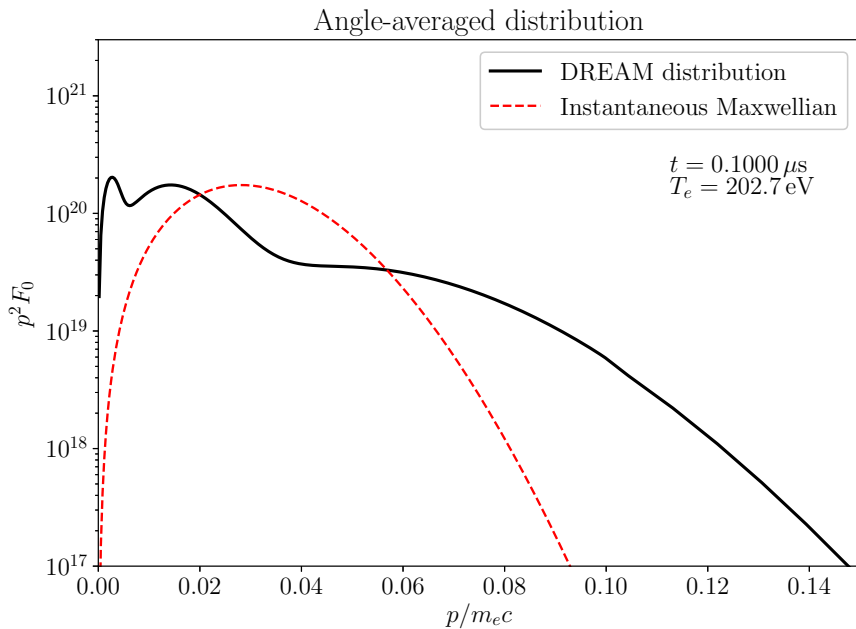
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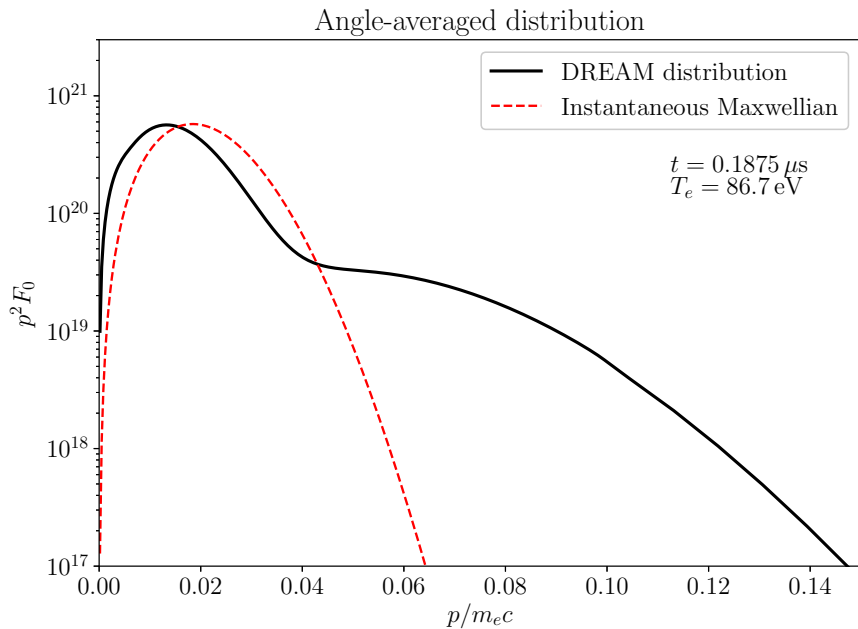
Kinetic electron model:

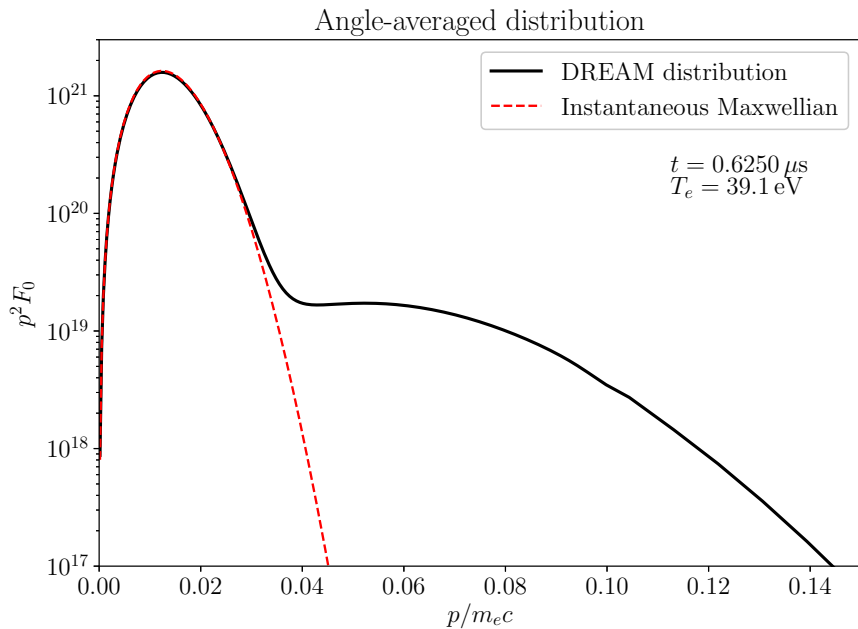
- Bounce averaged Fokker-Planck equation in zero-orbit-width limit
- Electric field acceleration and relativistic screened test-particle collisions
- Particle source ensuring density conservation
- Evolved self-consistently: Current density; electron and ion temperatures; electric field; ion charge states; poloidal flux





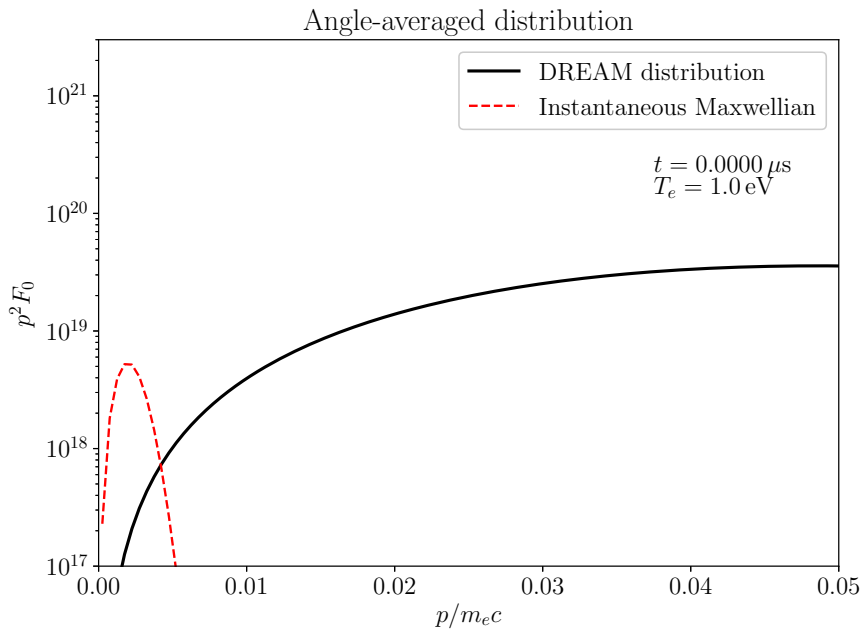


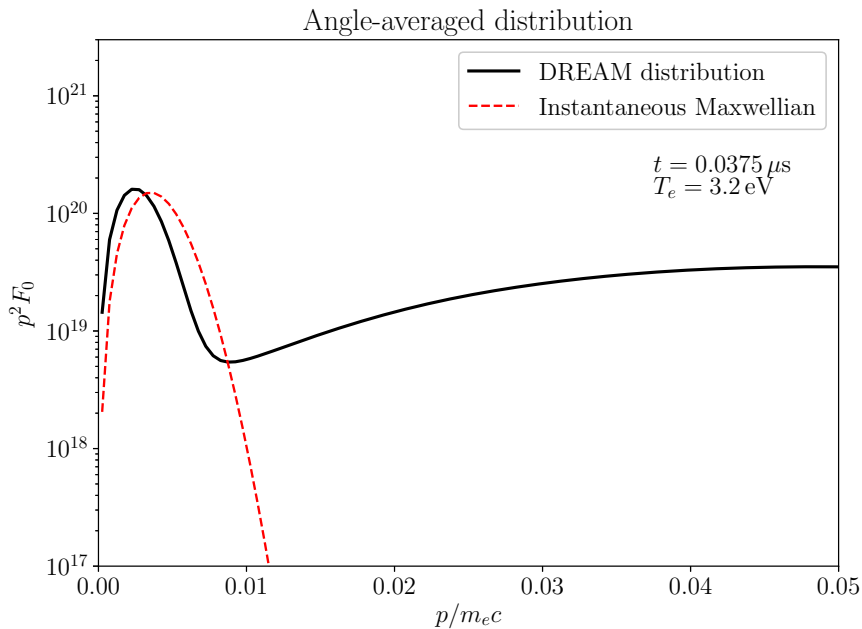


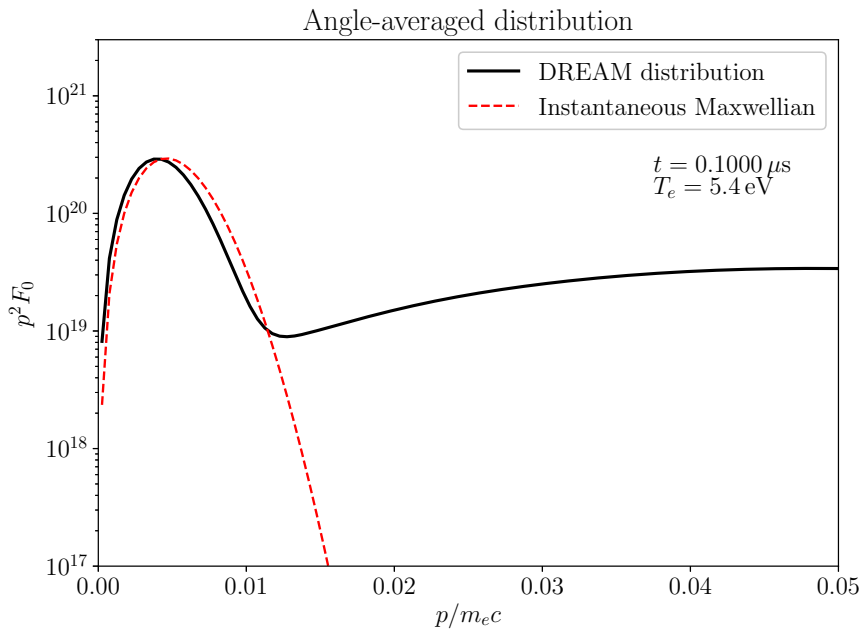


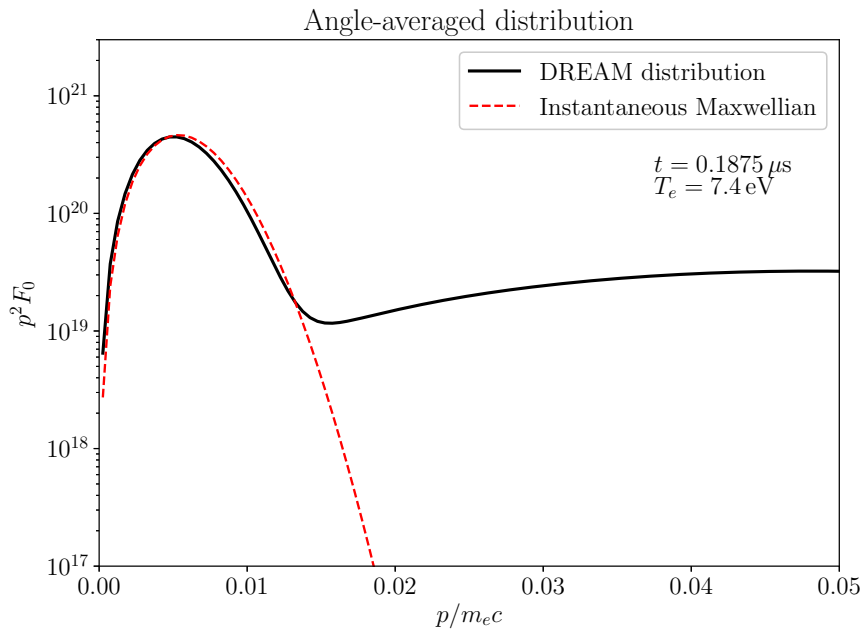
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- Non-linear electron collisions appear needed to describe the dynamics

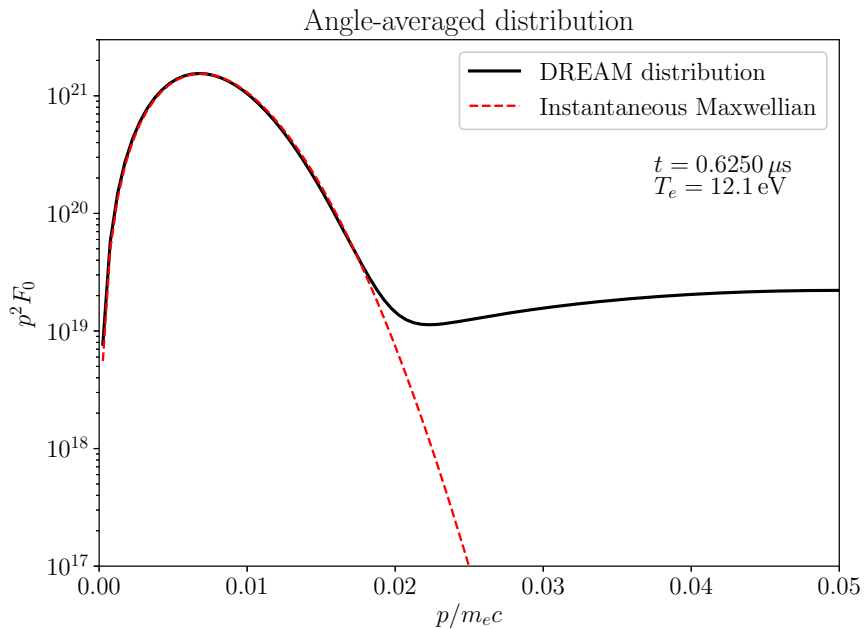
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- **Aleynikov & Breizman (NF 2017) presented an alternative picture:**
- Self-collisions among newly ionized electrons sufficiently fast to form cold Maxwellian
- Fluid electron density and temperature should represent that of newly ionized electrons (initialized at $n = T = 0$)
- Hot-tails form a second electron population which exchange energy with these *cold* electrons (+ ionize them)
- Energy still approximately conserved – eventually the approaches should end up at \approx same temperature



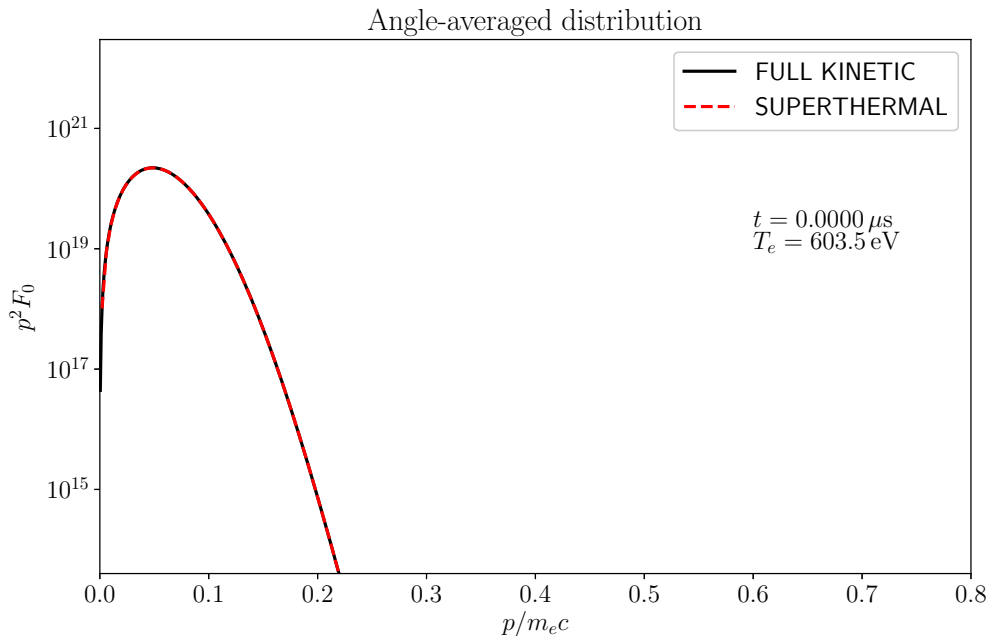


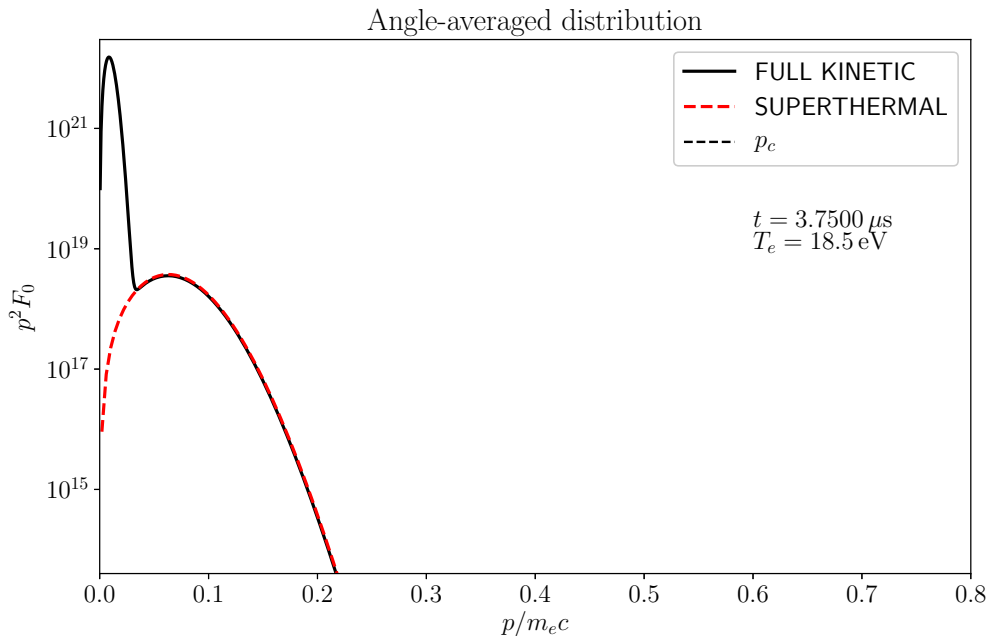




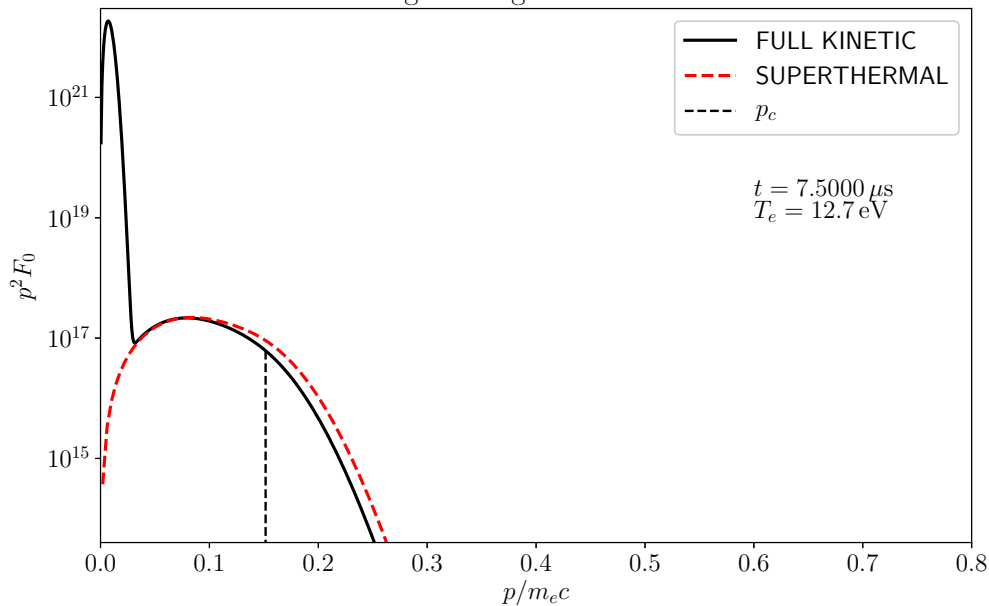


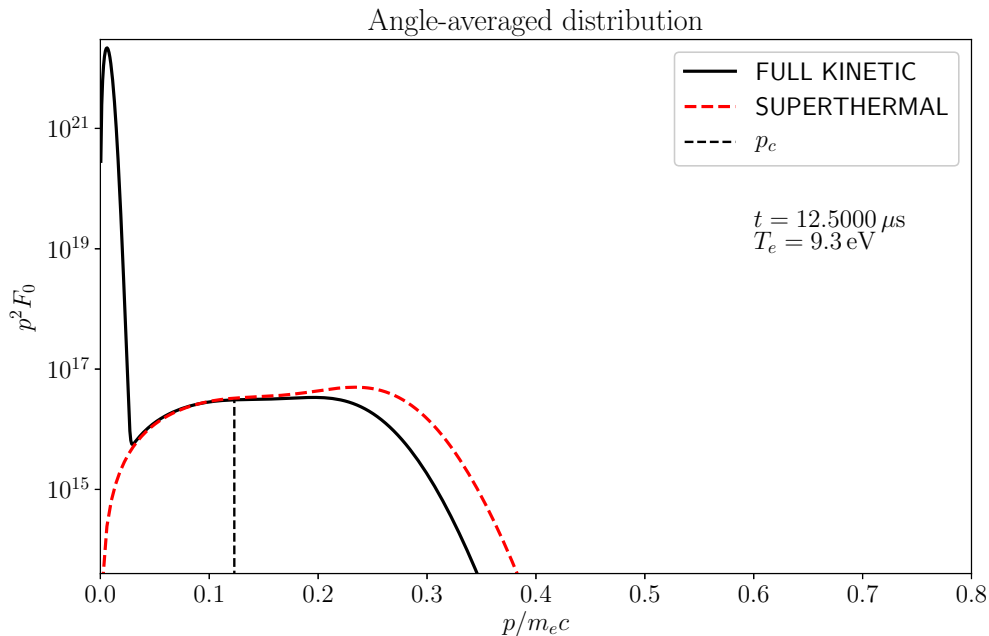
- “cold” population is essentially Maxwellian with $T \ll \langle W \rangle$: replace by fluid!
- Replace collision frequencies by superthermal limit
 - ▶ $\nu_s \propto n_{\text{cold}}/v^2\rho$
 - ▶ $\nu_D \propto n_{\text{cold}}/v\rho^2$
- Natural particle sink in $p = 0$: join electron flux with n_{cold}
- $n_{\text{cold}} = \sum_i Z_i n_i - \int f_{\text{hot}} d\mathbf{p} - n_{\text{RE}}$
- $j_{\parallel} = \sigma(T_{\text{cold}})E - e \int v_{\parallel} f_{\text{hot}} d\mathbf{p} + ec n_{\text{RE}}$

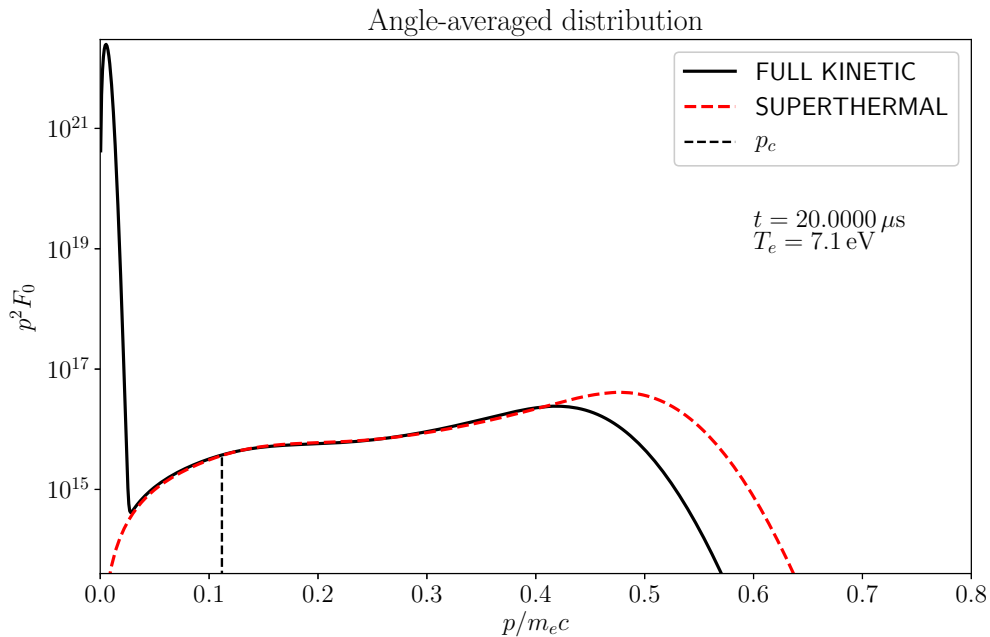




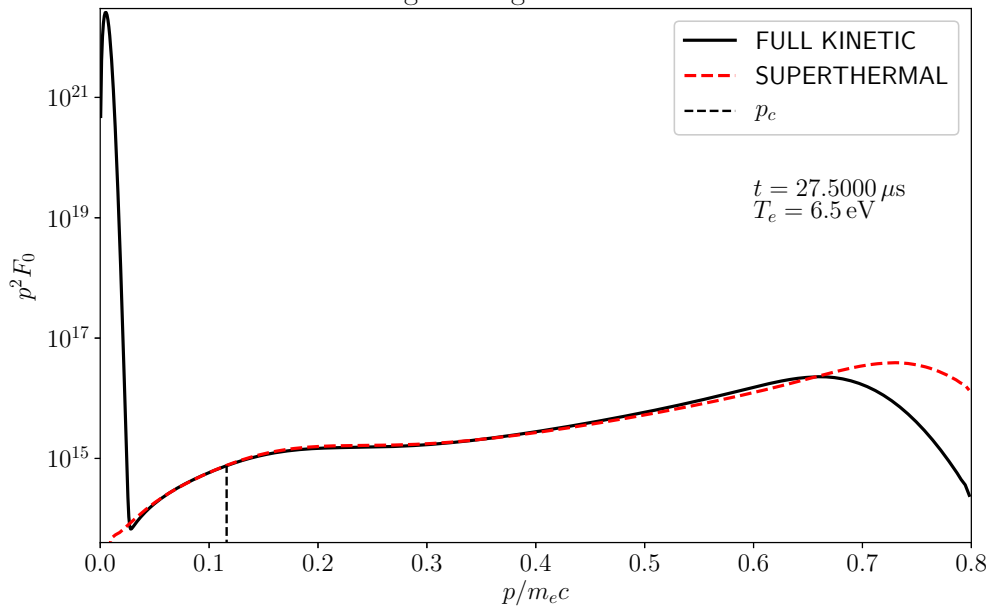
Angle-averaged distribution







Angle-averaged distribution



Approximations used to derive the avalanche growth rate [Rosenbluth & Putvinski 1997] can be applied to model hot tail:

- Asymptotic expansion with the ordering (“strong pitch-angle scattering”)

$$Z_{\text{eff}} = \mathcal{O}(1), \quad E = \mathcal{O}(\delta), \quad [\text{everything else}] = \mathcal{O}(\delta^2)$$

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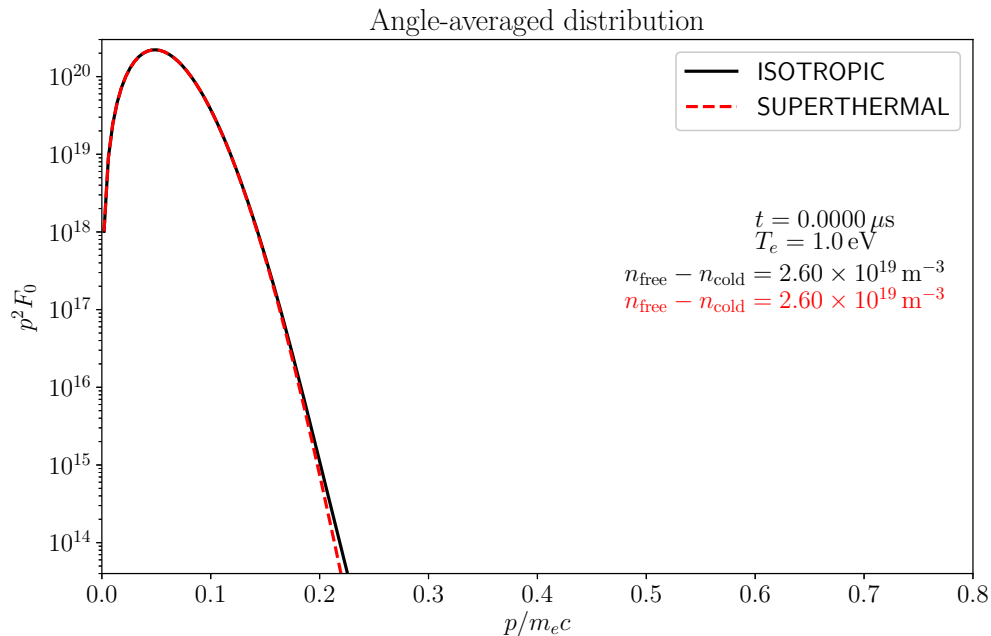
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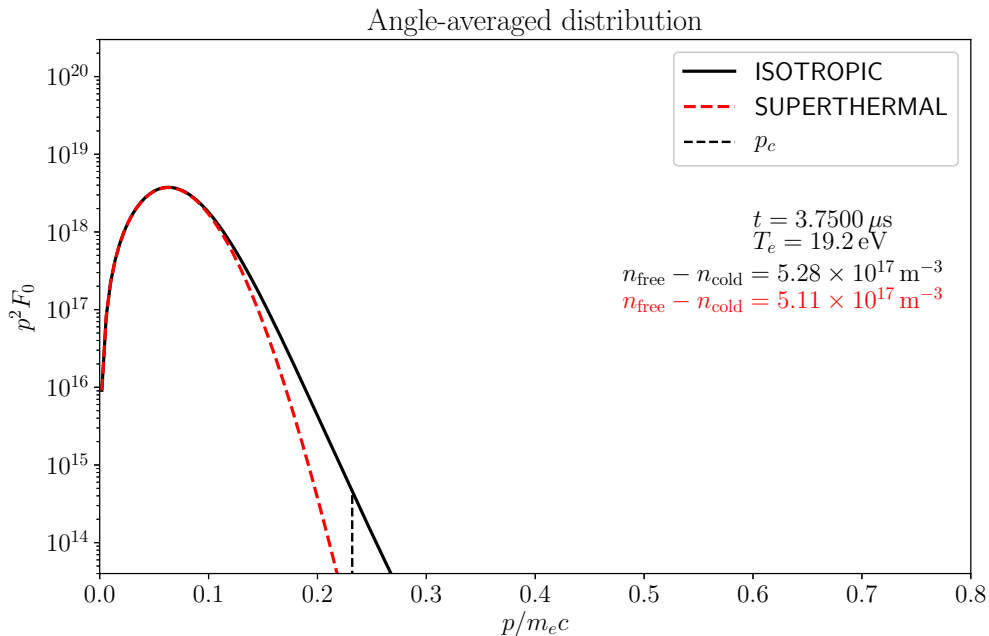
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- Leading order equation: $f_0 = f_0(t, r, p)$ (isotropic)
- Next-order equation yields

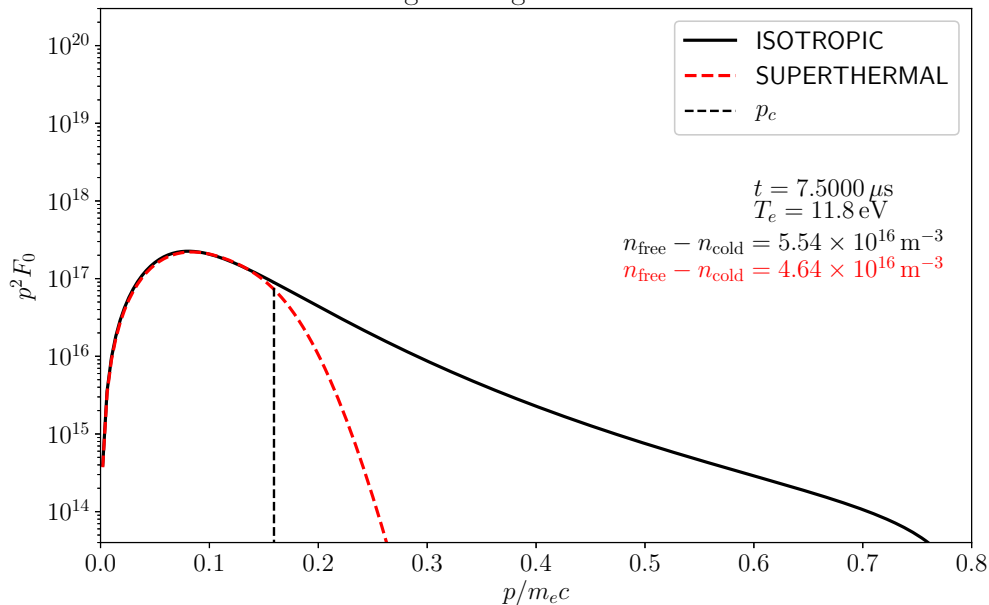
$$\frac{1}{\mathcal{V}'} \frac{\partial(\mathcal{V}' \{eE_{\parallel} \mathbf{b} \cdot \nabla p^i\} f)}{\partial p^i} \mapsto - \int_0^{1/B_{\text{max}}} \frac{\lambda d\lambda}{\langle \sqrt{1 - \lambda B} \rangle} \frac{(e\langle \mathbf{E} \cdot \mathbf{B} \rangle)^2}{4} \frac{1}{p^2} \frac{\partial}{\partial p} \left(\frac{p^2}{\nu_D} \frac{\partial f_0}{\partial p} \right)$$

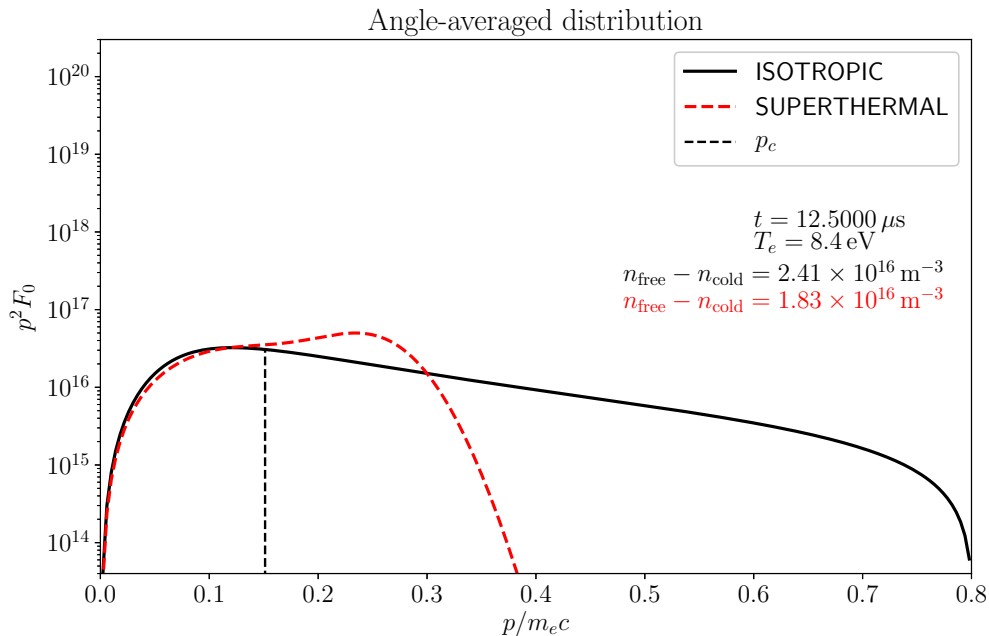
(bounce averaged E -field term becomes 1D energy diffusion term)



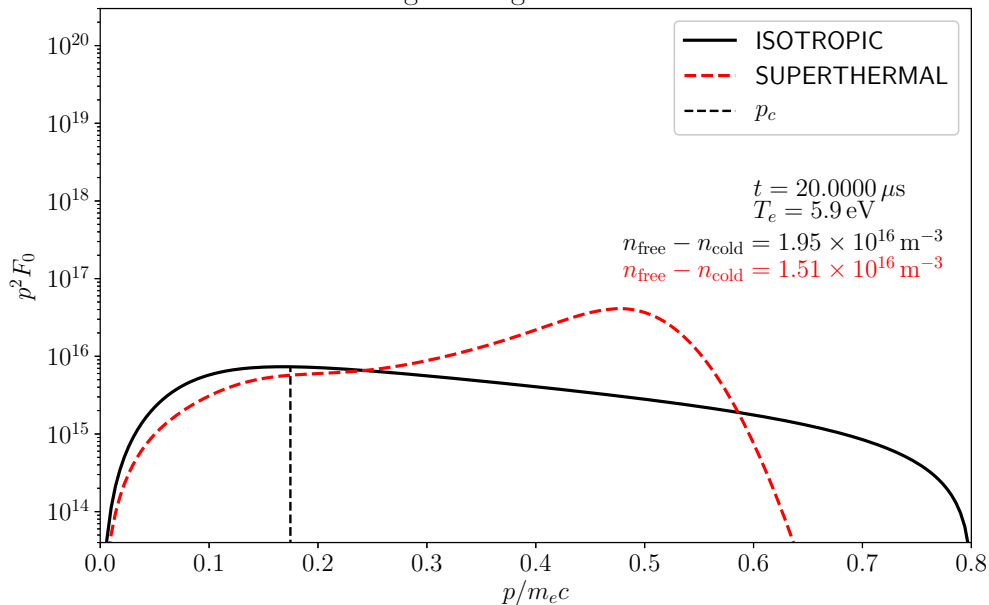


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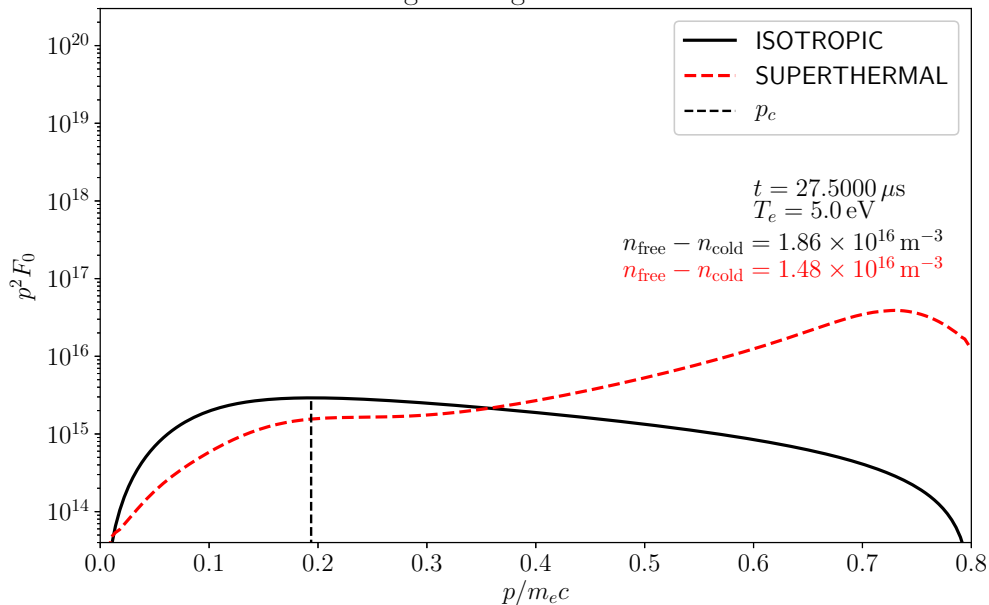




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The isotropic equation is essentially

$$\frac{\partial f_0}{\partial t} = \frac{1}{\rho^2} \frac{\partial}{\partial \rho} \left[\rho^2 \left(\rho \nu_s f_0 + \frac{k}{\nu_D} \frac{\partial f_0}{\partial \rho} \right) \right],$$

where

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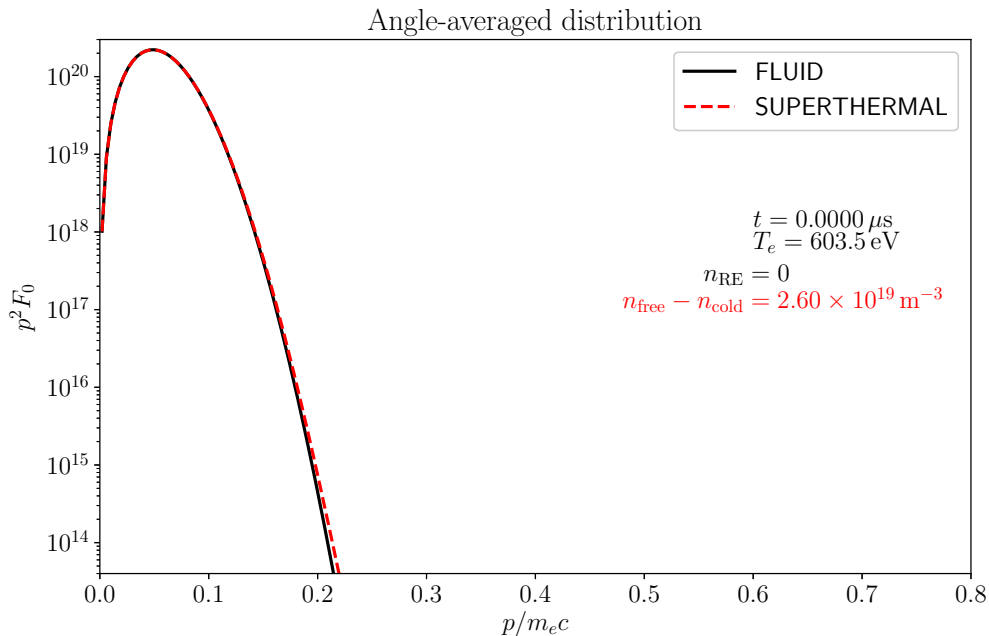
⇒ Sharp transition from ν_s -dominated to E -dominated at $p = p_0$:

$$\rho \nu_s f_0 + \frac{k}{\nu_D} \frac{\partial f_0}{\partial p} \Big|_{p=p_0} = 0$$

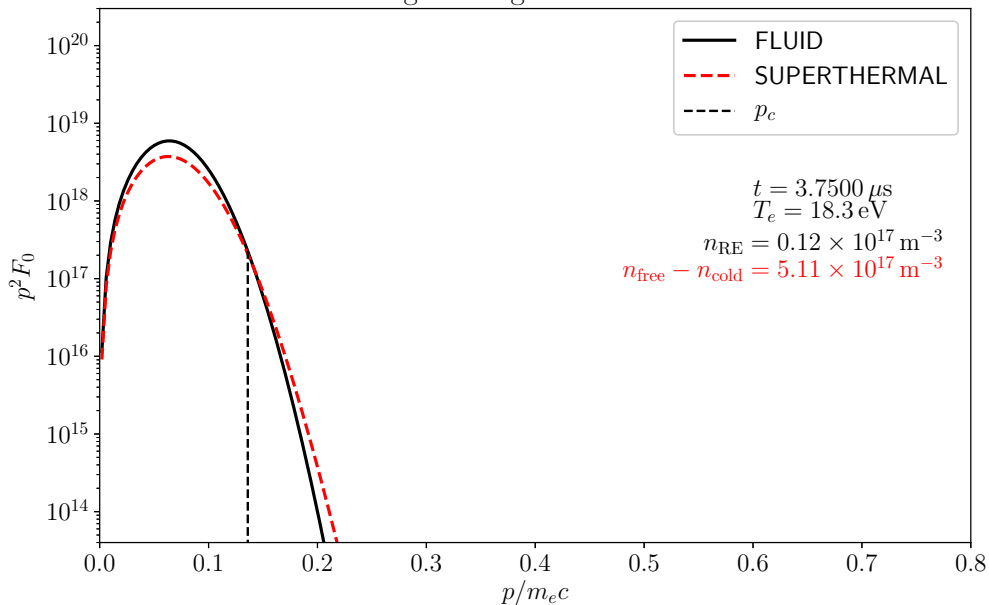
■ For $p < p_0$ the equation is [Smith & Verwiche (2008)]

$$\frac{\partial f_0}{\partial t} = \frac{1}{p^2} \frac{\partial p^3 \nu_s f_0}{\partial p}$$

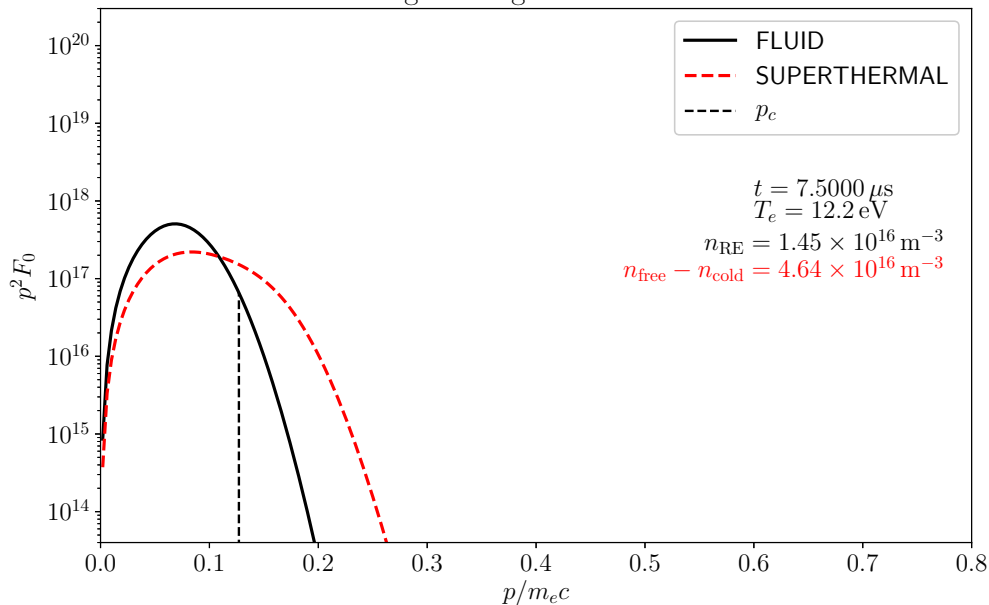
■ Analytic solution for f_0 , and $\partial n_{RE}/\partial t = -4\pi p_0^2 \dot{p}_0 f_0(p_0)$



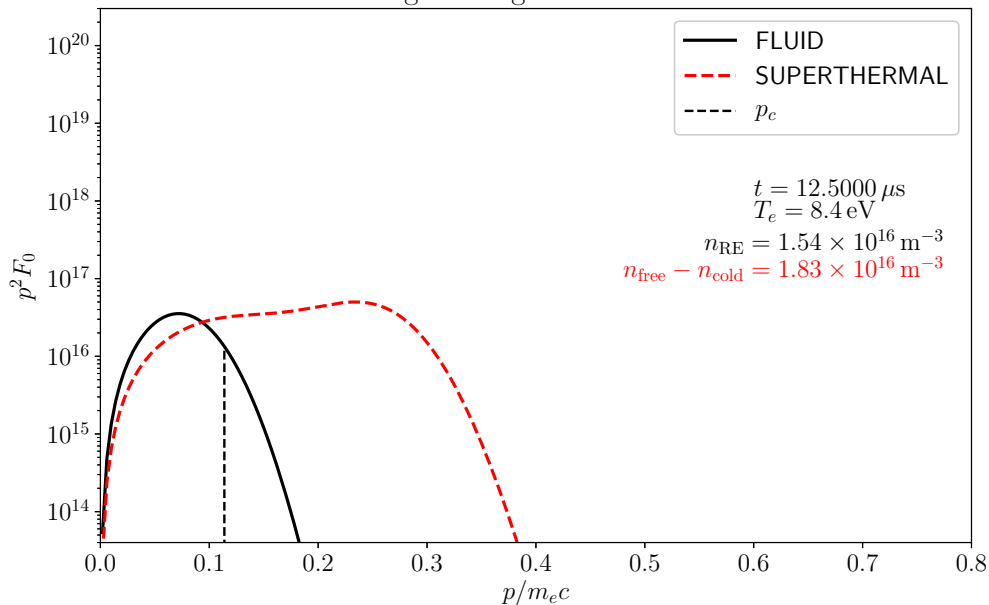
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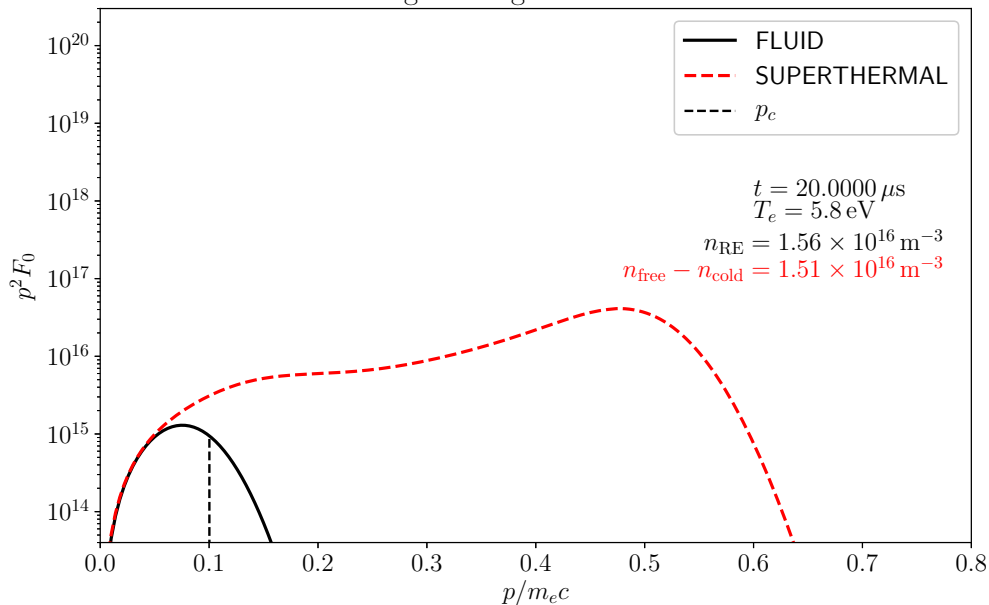
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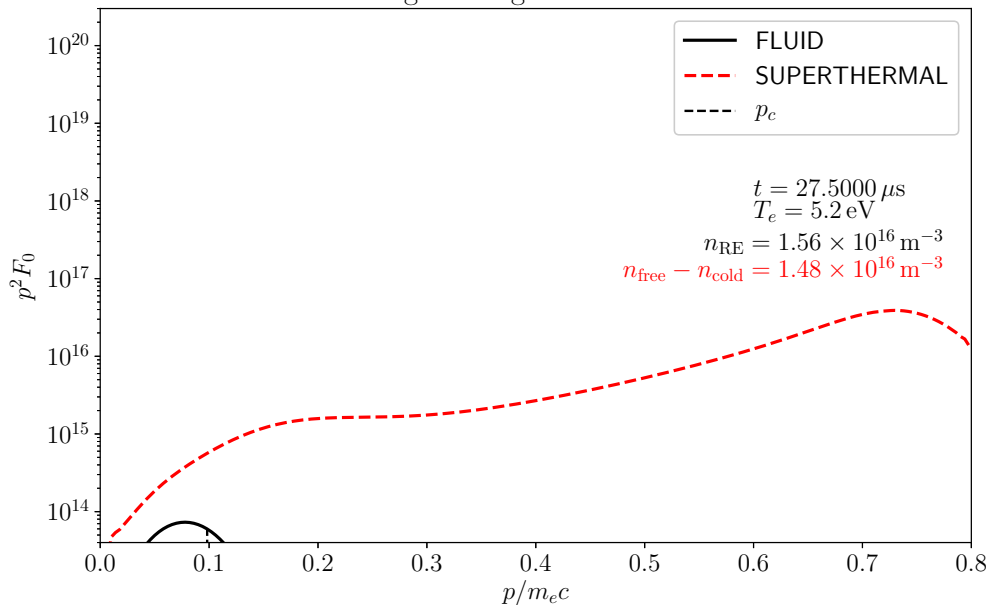
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- Ongoing “SUPERHERMAL FLUID” mode work:
 - ▶ Contribution to j_{\parallel} from hot distribution
 - ▶ Energy transfer from analytic slowing-down distribution to “cold” electrons
 - ▶ Ionization due to slowing-down distribution
- TODO: Screening effects (inelastic collisions) and Smith & Verwichte ρ_0 formula
- Note: Analytic hot tail model incompatible with fast-electron radial transport

- Resolution
 - $N_r \times N_p \times N_{xi} = 15 \times 140^* \times 68^*$
- Simulation time in illustrated test case:
 - ▶ “FULL KINETIC”: 4 hours
 - ▶ “SUPERHERMAL”: 1 hour 7 min
 - ▶ “ISOTROPIC”: 1 min 40 sec
 - ▶ “FLUID”: 25 sec
- Further developments in the works:
 - ▶ Q: with e.g. MGI, can the “cold” electrons equilibrate with the hot tail?
 - ▶ Stabilize “SUPERHERMAL” to support quiescent plasma before TQ
 - ▶ Future (probably): full non-linear “NORSE” collision operator

